

The Case for Spatial Filtering CHARA Classic

Dr. Gerard T. van Belle

Science Community Development Lead Michelson Science Center California Institute of Technology

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CHARA Collaboration Year-Three Science Review







CHARA Collaboration Year-Three Science Review



















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PTI Combination Scheme

- PTI's combined beams are fed into a NICMOS detector dewar
- The *white light channel* is a combined beam fed directly into the can
 - Used for immediate fringe tracking
- The *spectrally dispersed channel* is routed through a prism first
 - Also sent through a single mode fiber
 - Used for group delay updates, and ultimately used for science
 - Detected flux is $\sim 30\%$ of the WL side
- This setup delivers both sensitivity and precision

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PTI "Classic"

Reference: Colavita et al. 1999



PTI Data on η Boo

- Recently published result (van Belle, Ciardi & Boden 2007 ApJ 657 1058)
- Spectral channel data used for the paper, but WL channel data preserved in the data stream
- Allows for a demonstration of the performance of the non-spatially filtered WL channel versus spectrometer channel





Incoherent WL V² Time Trace -- 100125.sum



Incoherent Spec V² Time Trace -- 100125.sum



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η Boo: Non-Spatially Filtered









η Boo: Spatially Filtered









Error Histogram: Non-SF





Error Histogram: SF





Performance, Quantitatively

- Median individual V² measurement relative error (σ_{V2}/V^2): 0.039 (SF) versus 0.129 (non-SF)
- η Boo size: 2.154±0.048 mas (SF) versus
 2.255±0.148 mas (non-SF)



The Downside

- Alignment
 - Possibly a challenge as a function of telescope pointing (wander of telescope beam on fiber head)
- Multi-r₀ Regime
 - Gains may not be as substantial as for PTI
 - Will throw away more 'bad photons'



Why CHARA Classic?

- Most sensitive beam combiner
- Most used beam combiner
- Easiest to use
- Possibly an only 'slight' modification to the existing beamtrain
- Can be done for near-term operations







The Exhortation

- Even within the context of the caveats, substantial gains in performance are possible
- Possibly no (or at least little) sacrifice in sensitivity













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