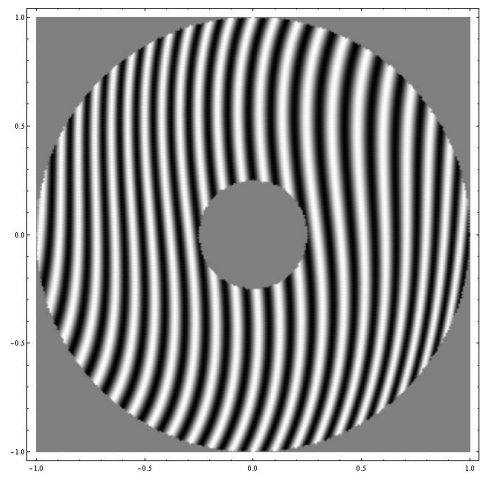




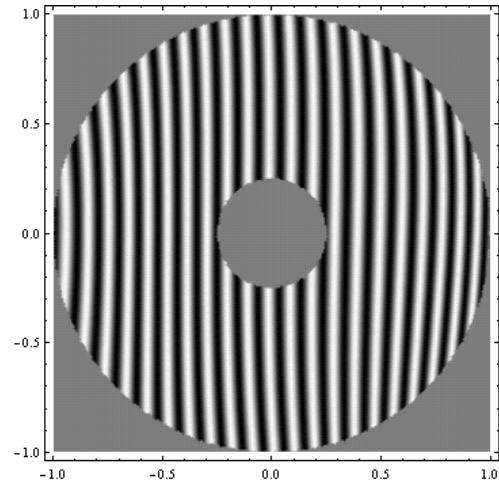
# Telescope Alignment

Laszlo Sturmann

How do we get from



to



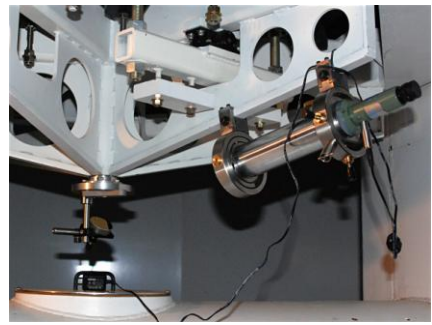
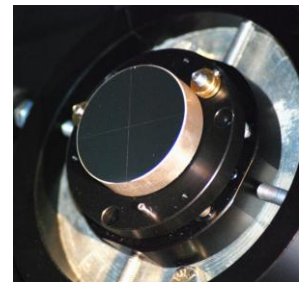
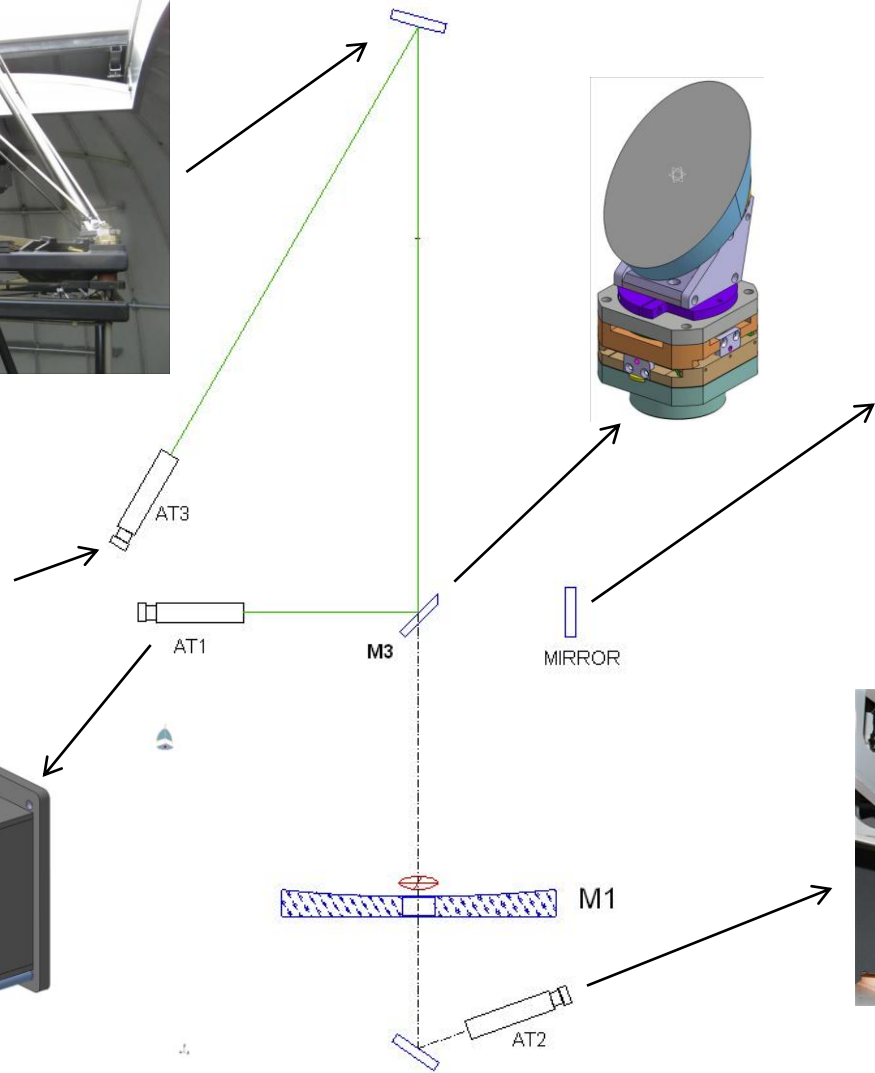
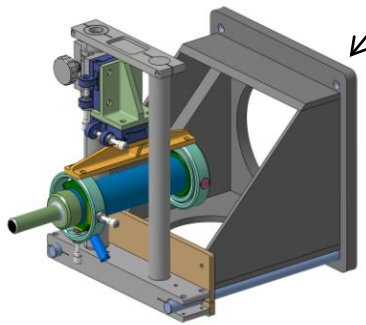
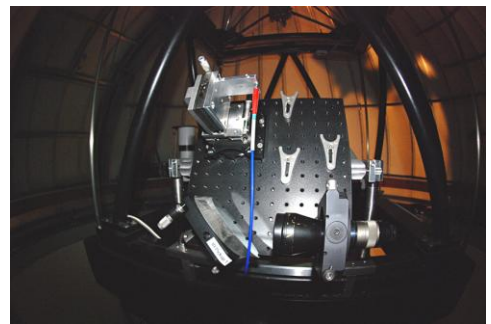
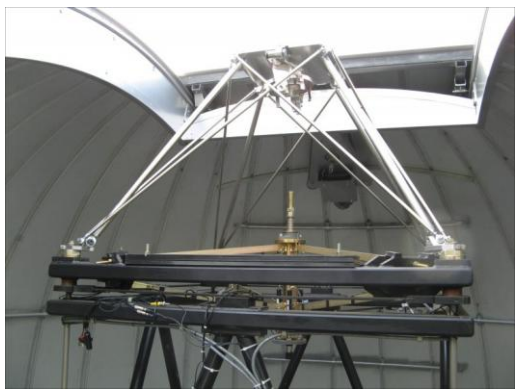
?

Perhaps Photoshop?

(Nils)

Pasadena 03/09/2010





a precise telescope alignment procedure was applied to E1, E2 and W1 in Spring 2009



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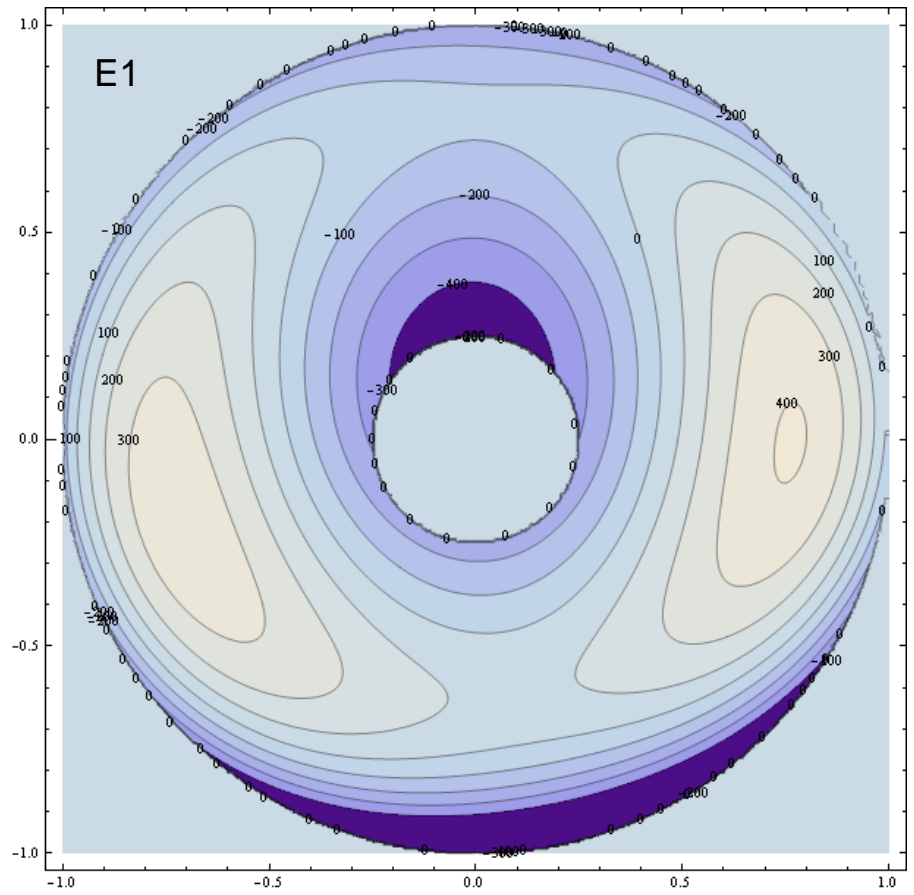
Observatoire de la CÔTE d'AZUR



# Wavefront Contour Plots

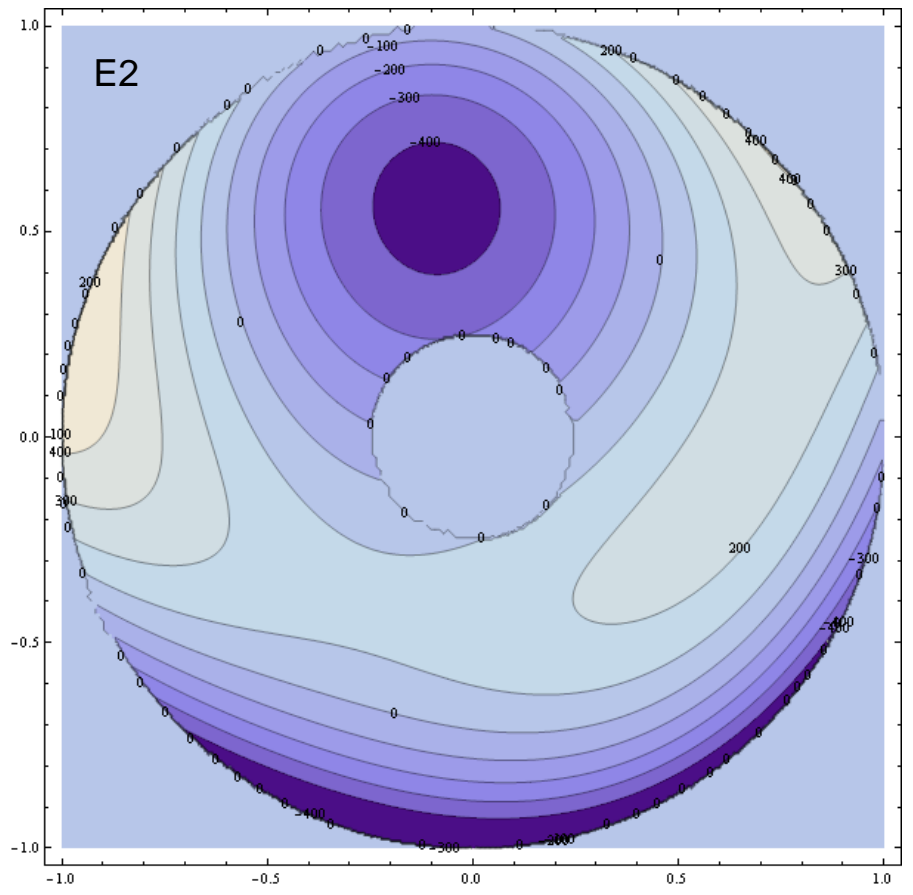
4/14/2009

P-V = 710 nm



5/18/2009

P-V = 960 nm



LESIA

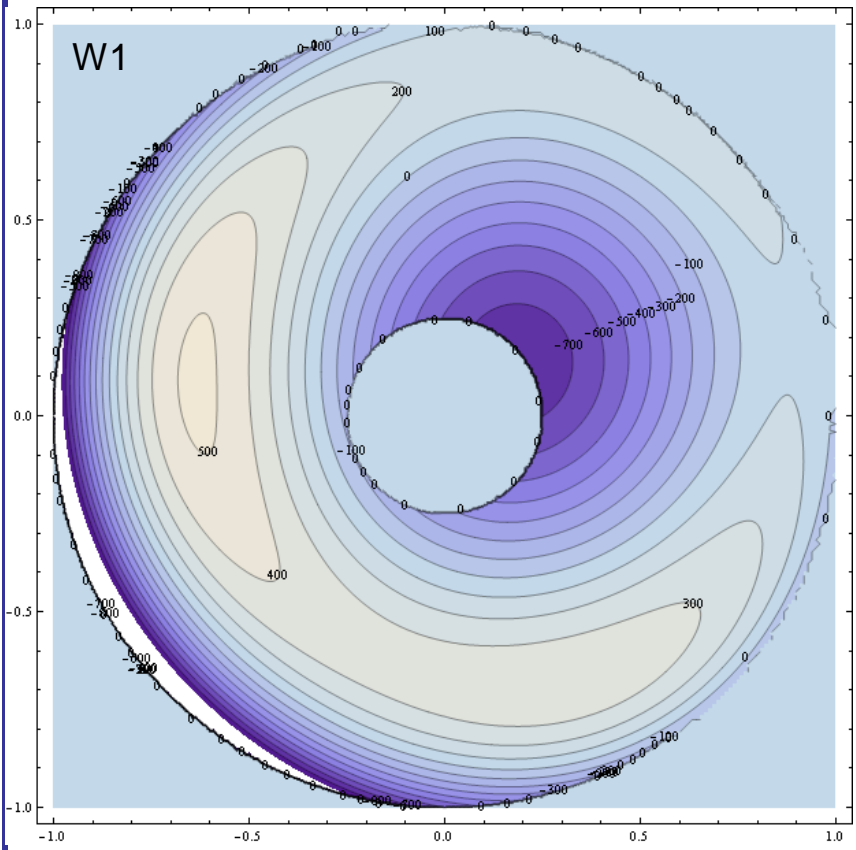


Observatoire de la CÔTE d'AZUR



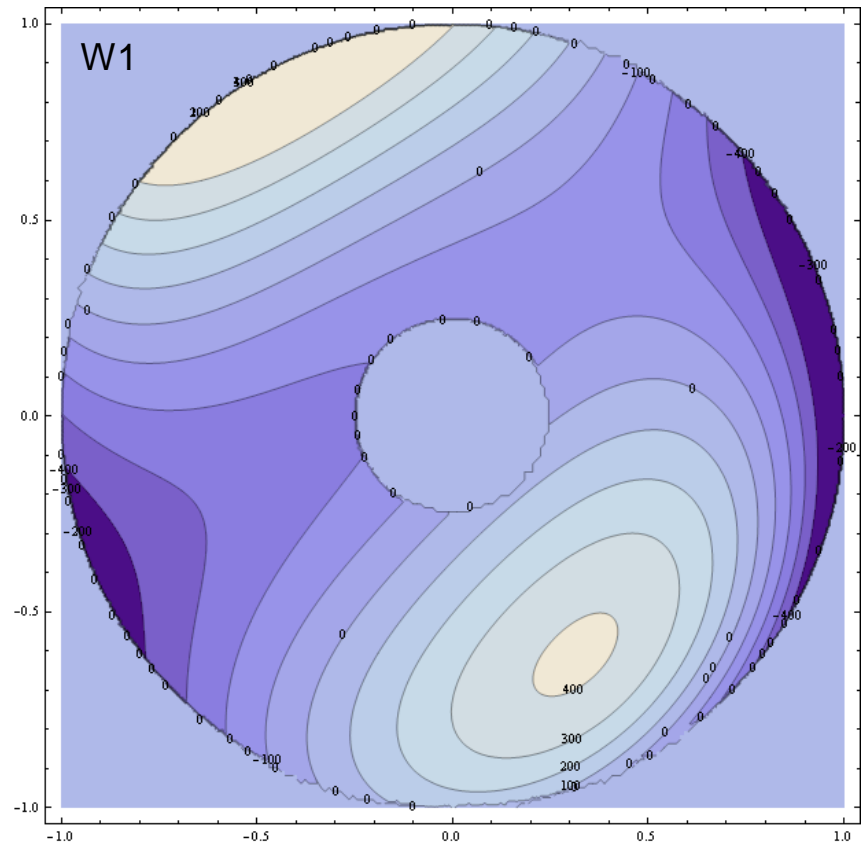
4/4/2009

P-V = 1940 nm



5/31/2009

P-V = 1280 nm



LESIA



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ef

4/14/2009

5/18/2009

5/31/2009

4/4/2009

Zernike [nm]	Name	E1	E2	W1-A	W1-B
4	focus	346	364	-842	-661
5	astigmatism	30	29	-96	-152
6	astigmatism	134	146	-31	-145
7	coma	93	154	136	70
8	coma	-6	-41	224	-81
9	trefoil	-12	45	13	22
10	trefoil	24	-48	-22	21
11	spherical	-152	-28	-191	-44

concentrating on these



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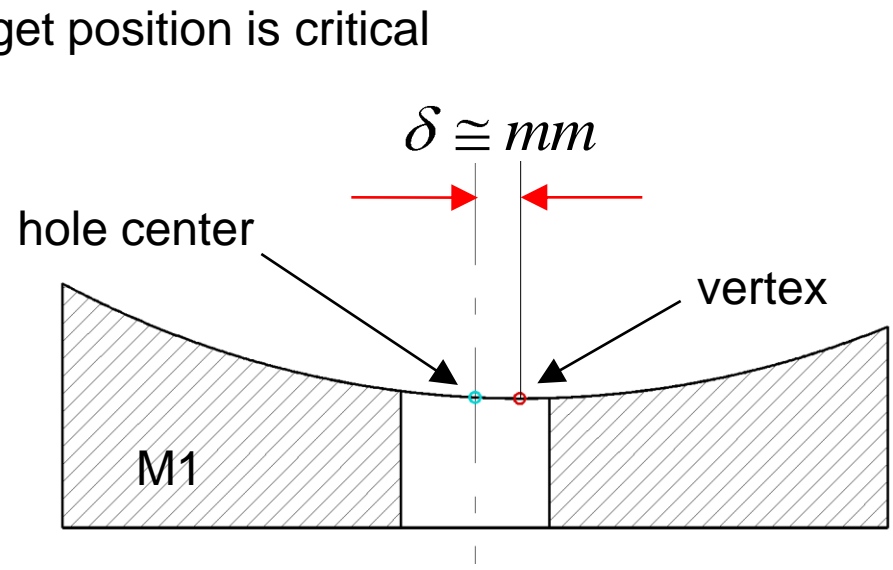
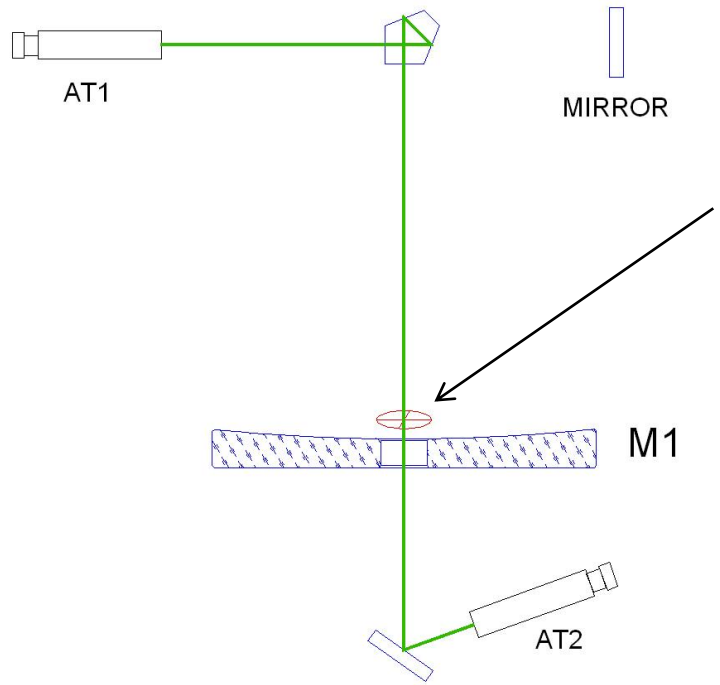


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The new alignment procedure is precise, but not necessarily accurate

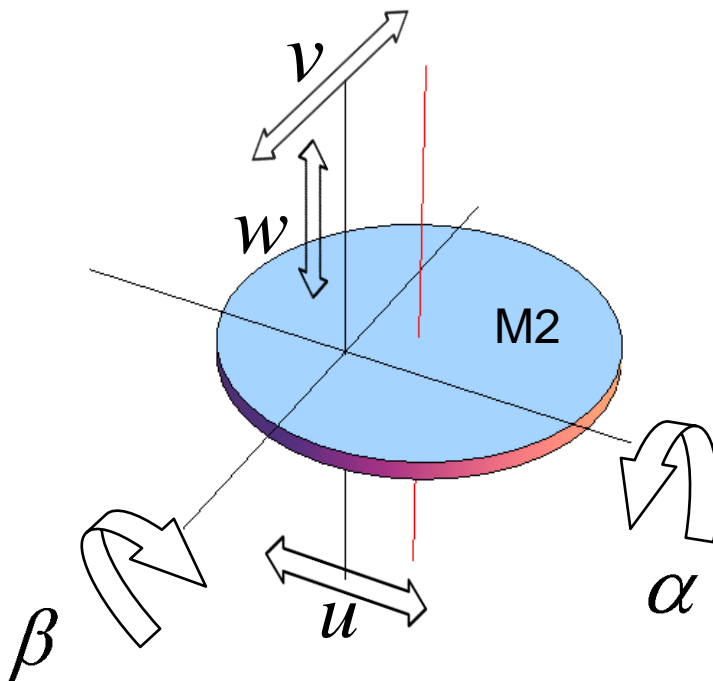
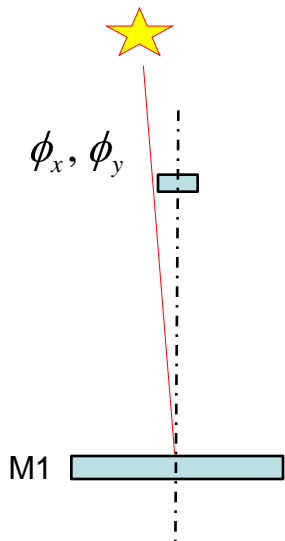


$\delta$  is unknown and may result in misalignment



# Hypothesis

M2 is misaligned and star is off-axis

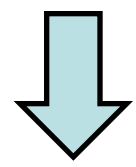


7 parameters

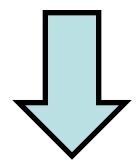


Action

slow beam quality evaluation due to seeing and 7 parameters



Random search for the best position/orientation of M2 is not feasible

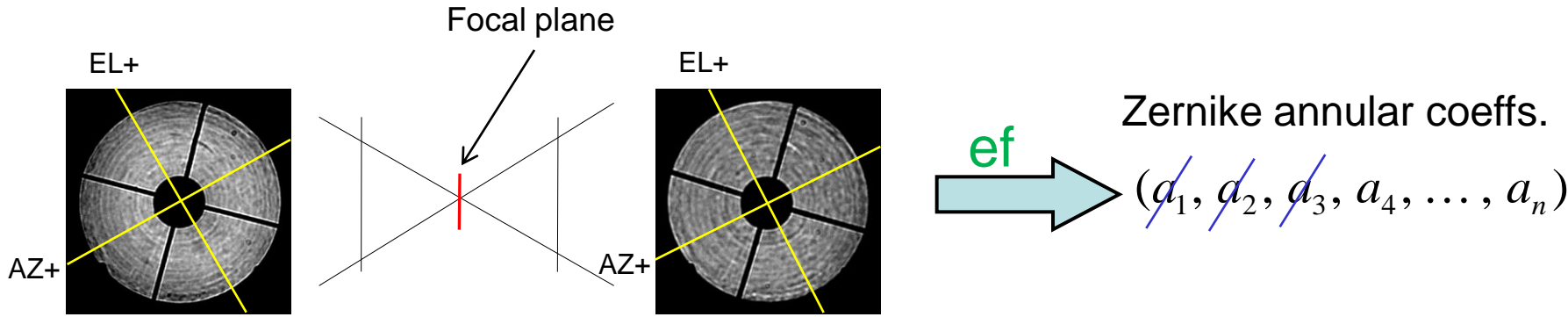


?



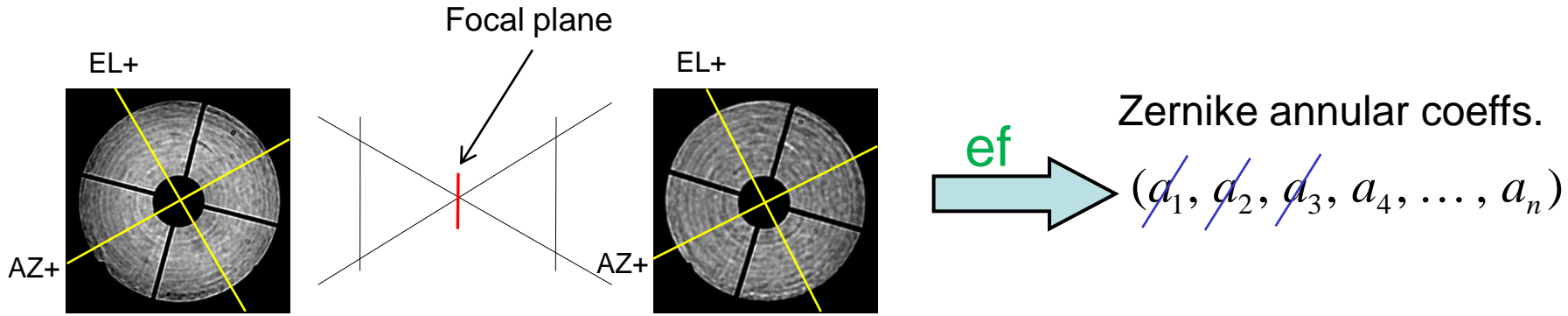


ef expands the wavefront into annular Zernike polynomials



$$W(\alpha, \beta, u, v, w, \phi_x, \phi_y) \cong \sum_i a_i Z_i(\rho, \theta, \varepsilon)$$

ef expands the wavefront into annular Zernike polynomials



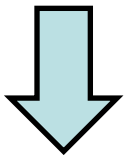
$$W(\alpha, \beta, u, v, w, \phi_x, \phi_y) \cong \sum_i a_i Z_i(\rho, \theta, \varepsilon)$$

$$a_i \xrightarrow{?} (\alpha, \beta, u, v, w, \phi_x, \phi_y)$$

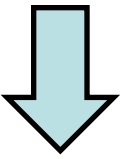


Action

slow beam quality evaluation because of seeing and 7 parameters



Random search is not feasible

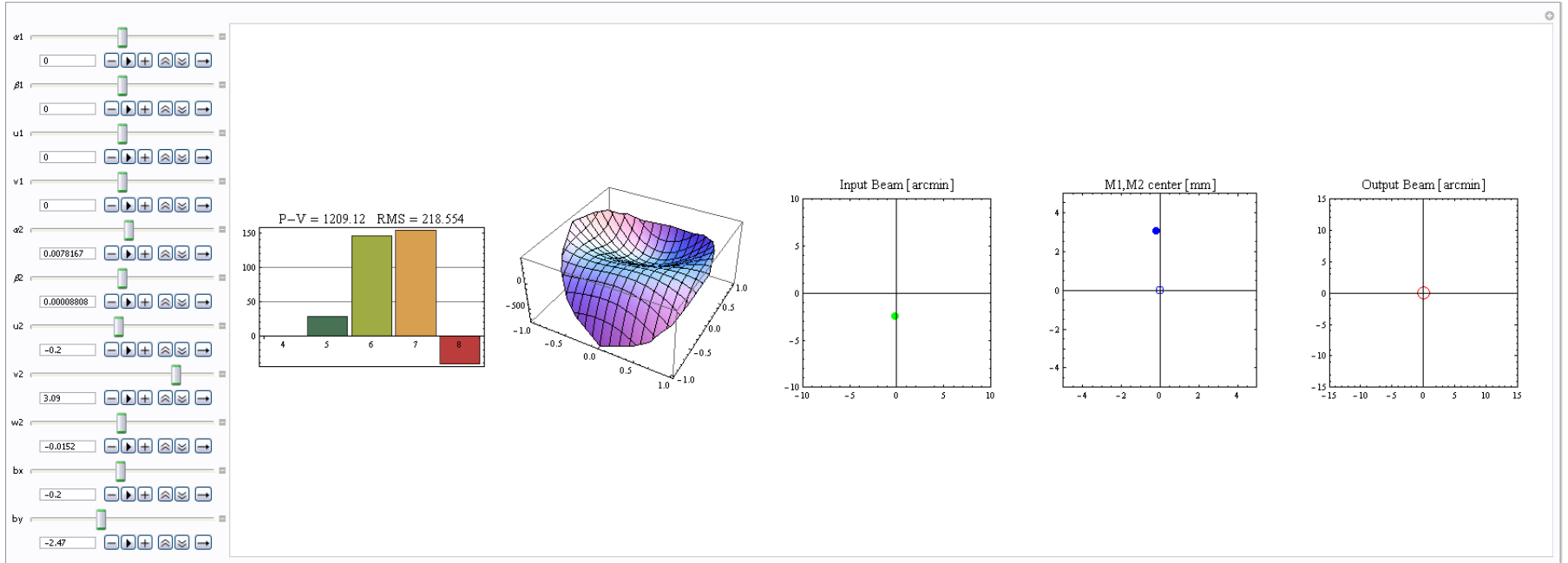


Estimating the position/orientation of M2 from Zernike coefficients





# SIMULATOR



Simulations showed:

$$(\alpha, \beta, u, v, \phi_x, \phi_y) \rightarrow a_5, a_6, a_7, a_8$$

$$w \rightarrow a_4, a_{11}$$

The effect on other coefficients is not significant.



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## GRID SEARCH

### Algorithm:

- given a set of  $\vec{a} = (a_5, a_6, a_7, a_8, a_{11})$  from measurements
- generate grid
- compute wavefront  $W$  in each grid points and expand it to annular Zernike polynomials  $\vec{b} = (b_5, b_6, b_7, b_8, b_{11})$
- compute the weighted average of grid coordinates of those points for which  $\Delta = (\vec{a} - \vec{b})(\vec{a} - \vec{b})$  is less than a certain limit


 $(\alpha', \beta', u', v', w', \phi'_x, \phi'_y)$



## Parameter Space:

- $5 \leq u[mm] \leq 5, \Delta u = 0.2mm,$
- $5 \leq v[mm] \leq 5, \Delta v = 0.2mm,$
- $3 \leq \phi_x[arc\ min] \leq 3, \Delta\phi_x = 0.2arc\ min,$
- $3 \leq \phi_y[arc\ min] \leq 3, \Delta\phi_y = 0.2arc\ min,$
- $40 \leq \alpha[arc\ min] \leq 40, \Delta\alpha = 1arc\ min,$
- $40 \leq \beta[arc\ min] \leq 40, \Delta\beta = 1arc\ min.$

**Brute force is not “brute” enough**

# of grid points:  $51 \times 51 \times 31 \times 31 \times 81 \times 81 = 16,399,619,721$

**Simplification is needed!**





Assumption:

the direction of the outgoing beam is close to be correct

$$\alpha = \alpha(v, \phi_y) = \frac{v - 0.735\phi_y}{628},$$

$$\beta = \beta(u, \phi_x) = \frac{0.735\phi_x - u}{628}.$$

5 free parameters

Treated separately  $\left\{ \begin{array}{l} (\alpha(v, \phi_y), \beta(u, \phi_x), u, v, \phi_x, \phi_y) \rightarrow a_5, a_6, a_7, a_8 \\ w \rightarrow a_4, a_{11}, \quad \text{All other } a_i \equiv 0 \end{array} \right.$

4 parameter grid search provides unique solution

$(u, v, \phi_x, \phi_y)$



## Simplified Algorithm

given  $\vec{a} = (a_5, a_6, a_7, a_8)$  from **ef**

grid points  $(\alpha(v_j, \phi_{y,l}), \beta(u_i, \phi_{x,k}), u_i, v_j, \phi_{x,k}, \phi_{y,l})$

$W$  in each grid point expanded into annular Zernike polynomials

$$\vec{b} = (b_5, b_6, b_7, b_8)$$

Compute  $\Delta = (\vec{a} - \vec{b})(\vec{a} - \vec{b})$  for each grid point

Select the grid point where  $\Delta$  is minimum

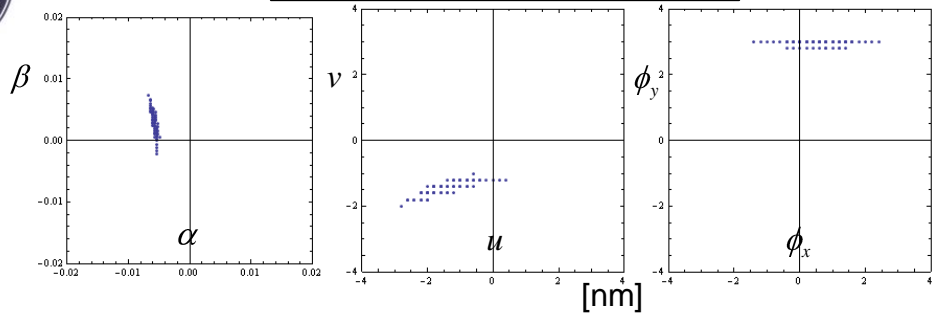
repeat it on finer grid around minimum

Adjust  $W$  to match  $a_{11}$

Found a unique solution!



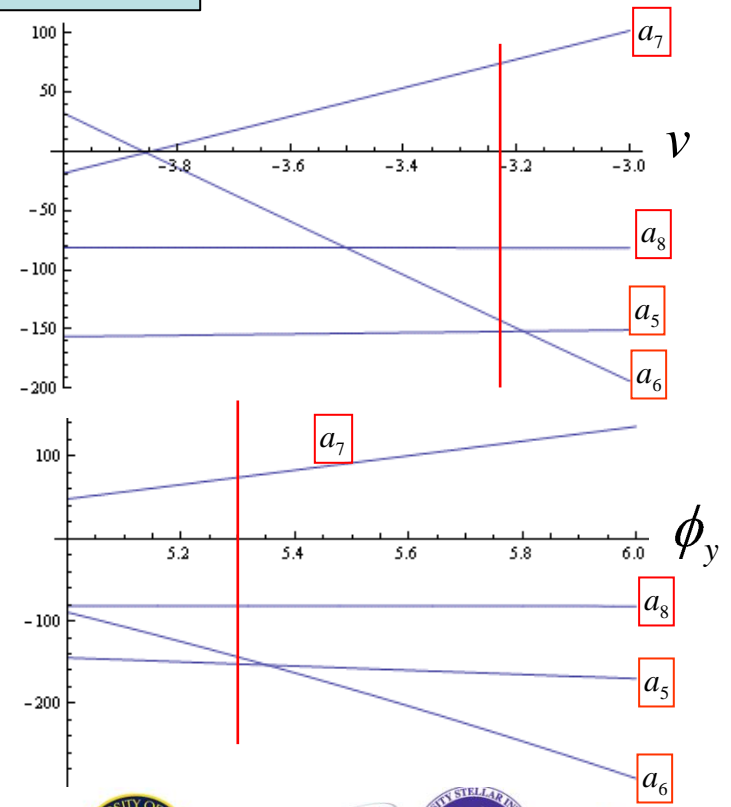
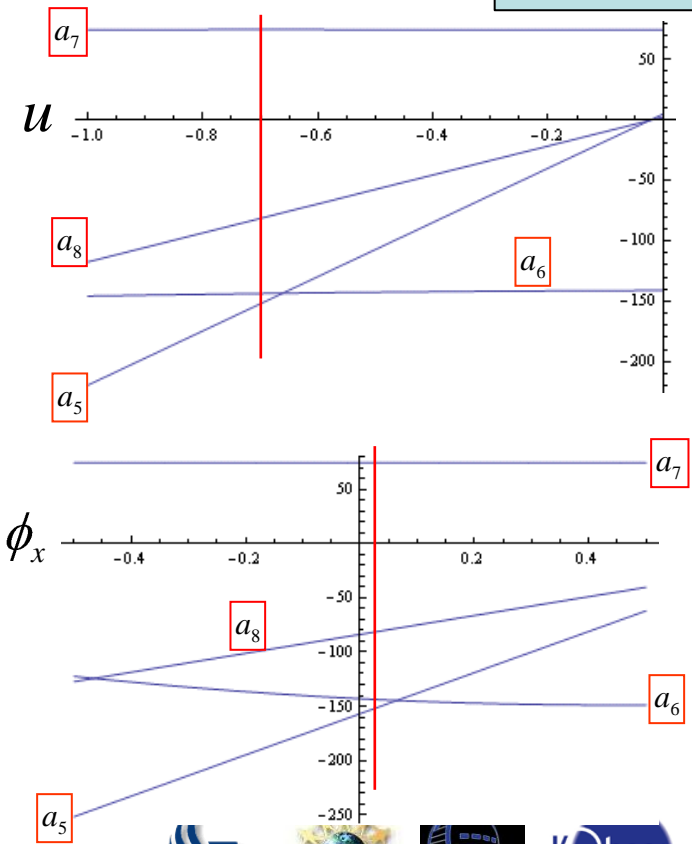
Clustering in parameter space



Target:  $a_5 = -152nm$   
 $a_6 = -145nm$   
 $a_7 = 70nm$   
 $a_8 = -81nm$

W1

Variation of Zernike coefficients





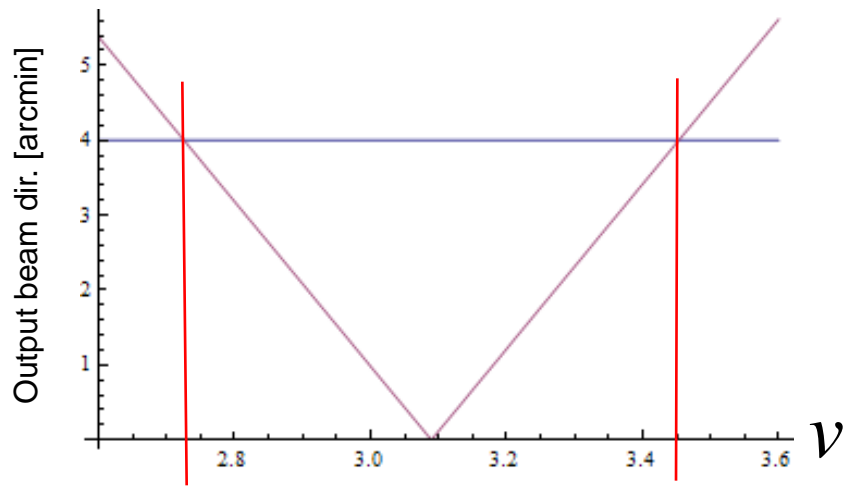
Error estimates:

Not done yet

- random error in  $a_i$  ~15%  
more measurements, S-H sensor

- the output beam direction is not (0,0,-1) → grid-point error →

Biased solution





Solutions found

	$\alpha$ [arcmin]	$\beta$ [arcmin]	$u$ [mm]	$v$ [mm]	$w$ [mm]	$\phi_x$ [arcmin]	$\phi_y$ [arcmin]
<b>E1</b>	36	8	-0.7	3.7	0.43	0.95	-4.1
<b>E2</b>	27±2	0.3±2	-0.2±0.4	3.1±0.4	0.065	-0.2±0.5	-2.5±0.5
<b>W1</b>	39	4	-0.7	-3.2	0.21	0.03	5.3

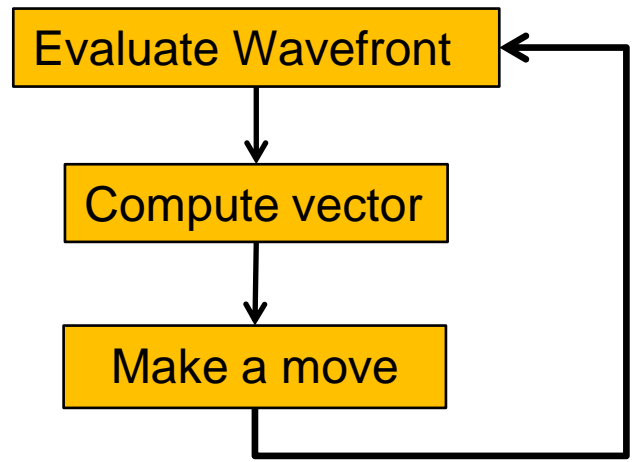
provides a way to estimate  $\delta \longrightarrow$  More accurate placement of M1 target

Complications:

lateral adjustments of M2 ( $u, v$ ) is difficult

pivot is not at vertex of M2 ( $\alpha, \beta$ )

coupling among ( $u, v, \alpha, \beta$ )





# Hardware Updates

- Ovenized crystal oscillator in the clock
- Three domes have been retrofitted with the new dome wheels
- Differential Image Motion Monitor (DIMM) on S1
- New M3 mount (W1 & W2)
- Coude-box cover prototype
- AZ limit switch mounting
- Particulate Monitor on S2
- New alignment telescope mount
- Serial bus on the telescope
- RS-485 interface for the finder and acq. cameras
- RS-485 interface for temperature/humidity sensor

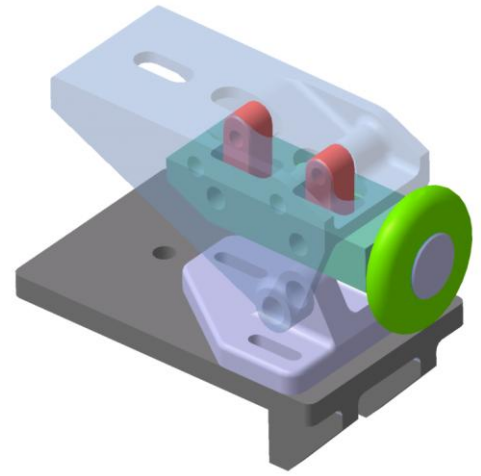
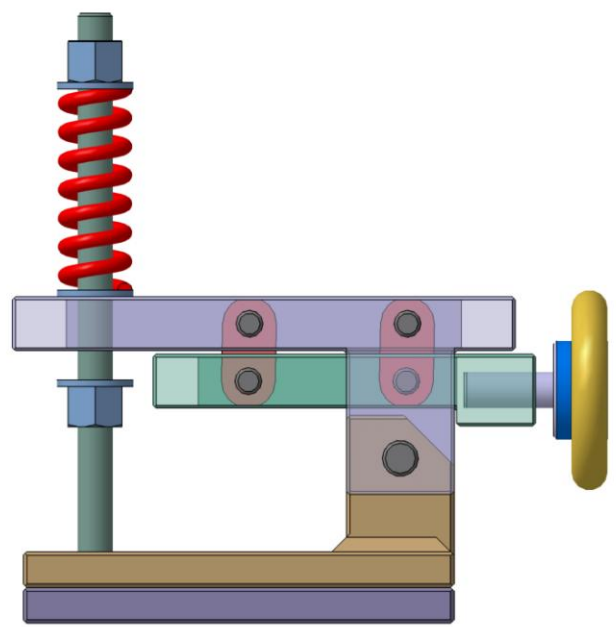
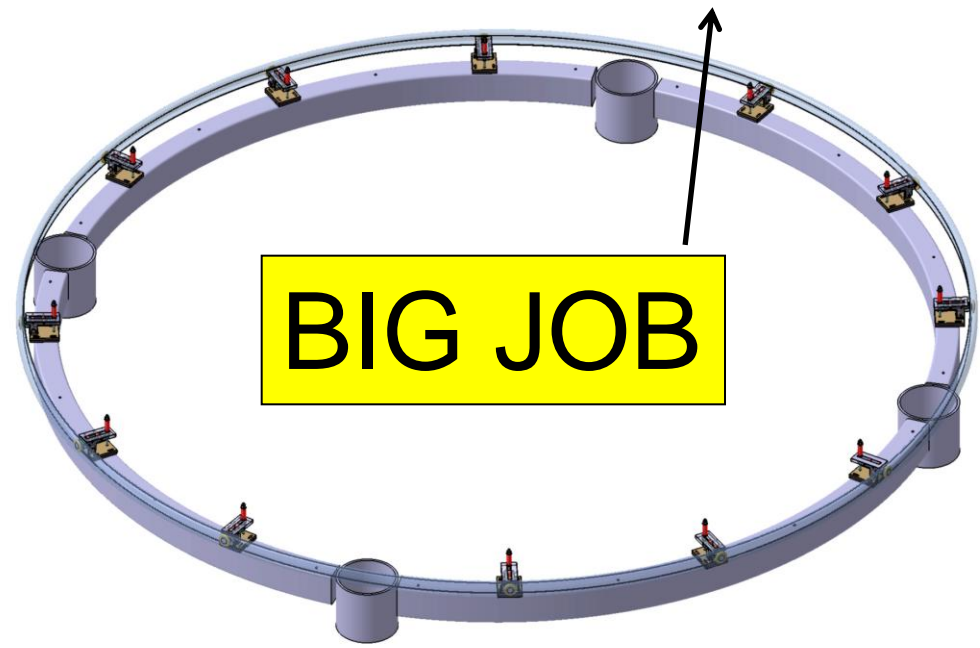




# Dome Rollers

three domes done, three to go (Golden, Webster)

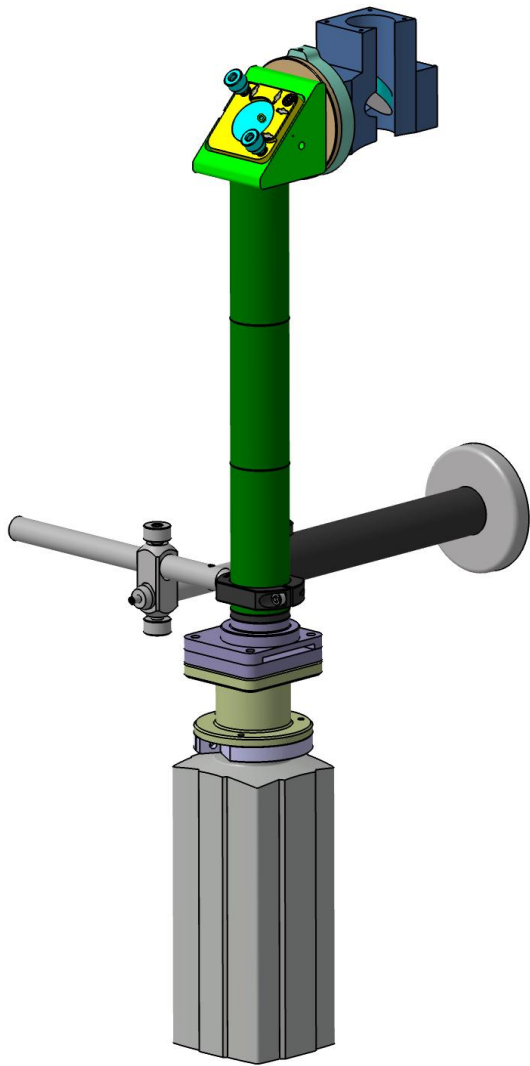
“floating” dome support equalizes load on the individual wheels, minimizes wear and load on the drive motor, quieter





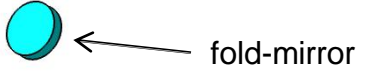
# DIMM

## Mounted on S1



Rotatable mirror in the acq. telescope

**Needs motor for remote operation**



fold-mirror

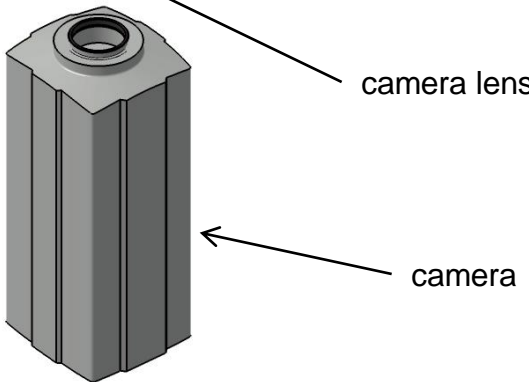


collimator



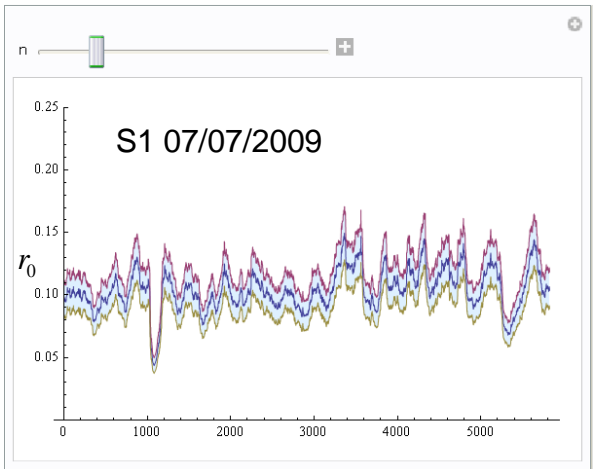
mask

wedge prisms



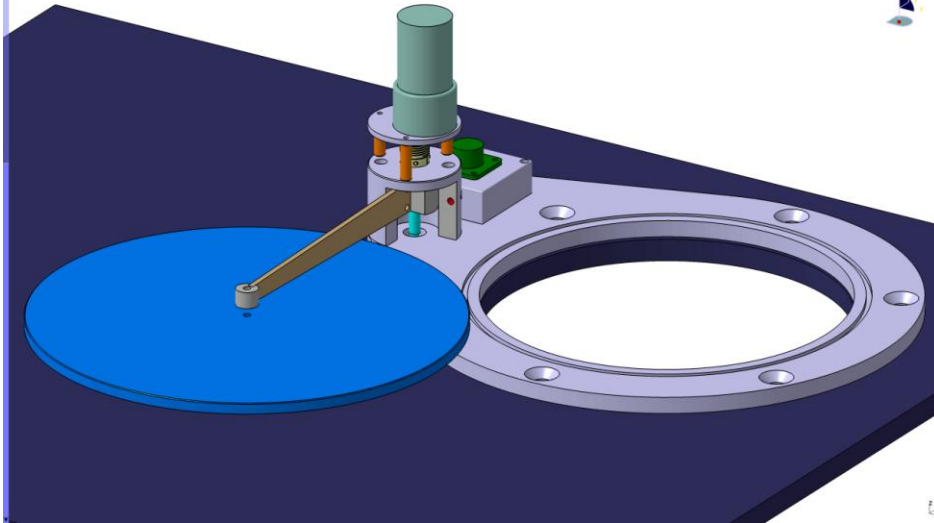
camera lens

camera

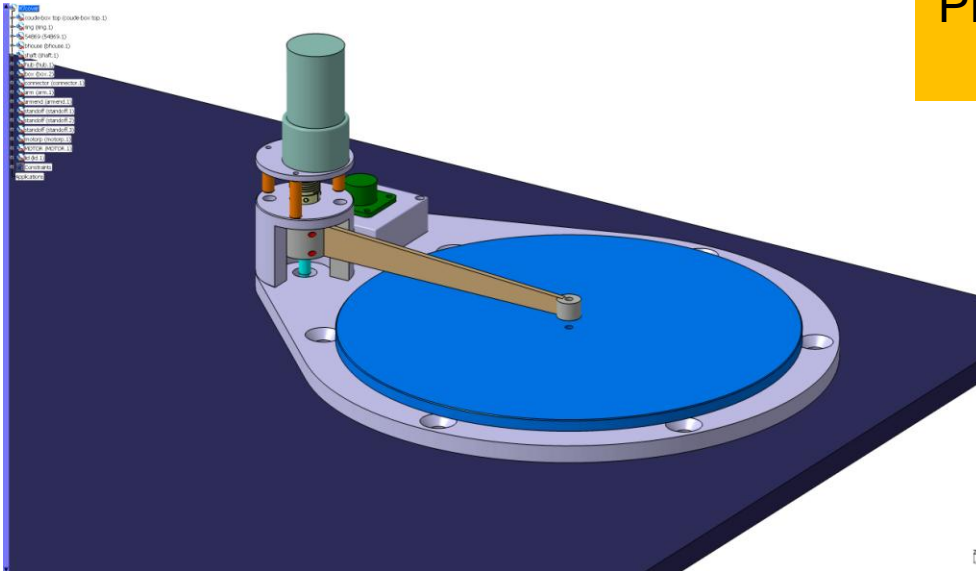




coude-box cover

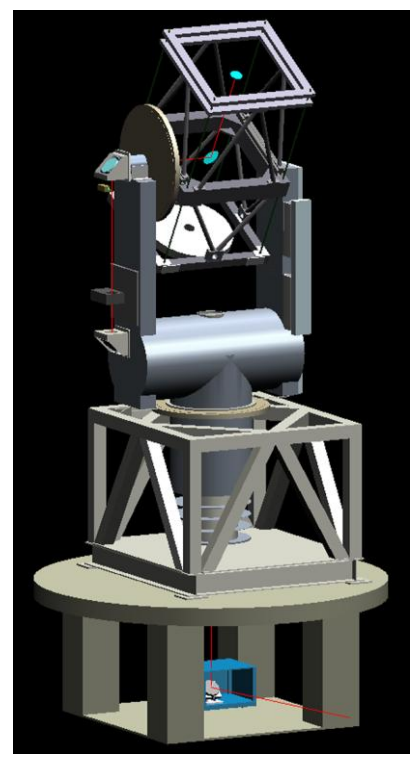
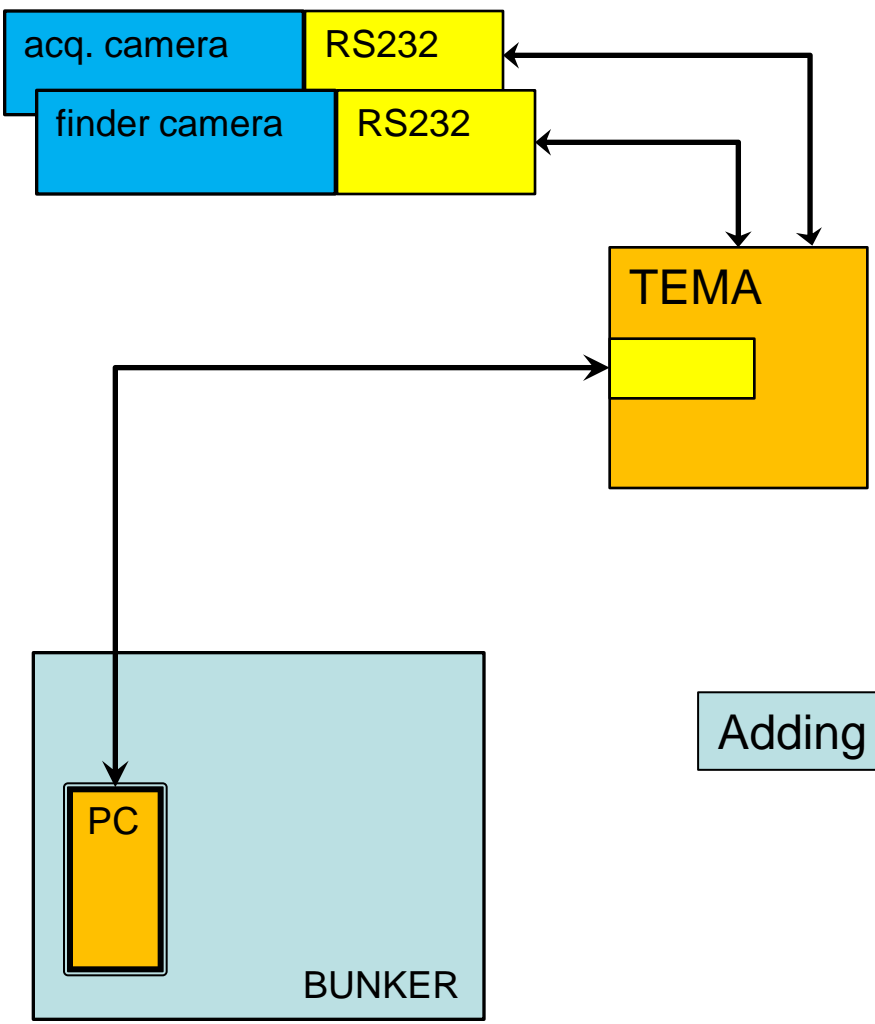


Prototype has been built but not tested yet.





# RS-485 Bus on the Telescope



Adding serial ports to TEMA is not practical.



# RS-485 Bus on the Telescope

Implementing a serial bus on the telescope and hooking up devices to the bus makes more sense. Simpler cabling, expandable.

