Progress at the Magdalena Ridge Observatory Interferometer

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Magdalena Ridge Observatory

Federally funded since 2000 EIS completed in 2003 Two facilities at MRO Fast-tracking 2.4m NIR/Optical I0-element interferometer 2.4m scope started full operations Aug, 2008 75% NASA/DoD funded

MROI is 10 1.4m movable afocal telescopes in equilateral Configuration Optical and near-IR operation Baselines from 7.5 to 340m **Minimized reflections** Design optimized for imaging mission 🚝

MROI Key Science Mission

- AGN:
 - Verification of the unified model
 - Determination of nature of nuclear/extra-nuclear starbursts
 - H = 14 gives > 100 targets.
- Star and planet formation:
 - Protostellar accretion, imaging of dust disks, disk clearing as evidence for planet formation
 - Emission line imaging of jets, outflows and magnetically channeled accretion.
 - Detection of sub-stellar companions.
- Stellar accretion and mass loss:
 - Convection, mass loss and mass transfer in single and multi-star systems
 - Bipolarity and collimation of circumstellar material, wind and shock geometries.
 - Pulsations in Cepheids, Miras, RV Tauris, etc.





03/01/2011

CHARA Science Meeting

Requirements Flowdown

- Telescope diameter of 1.4 m
 - H magnitude = 14 for group delay tracking limit
- Spatial scales of 0.3 to 30 mas
 - Baselines from 8 to 350 m (for 0.6-2.4 microns)
- Moderate-to-high spectral resolutions
 - Separate fringe tracking and science cameras
- High throughput to achieve sensitivity limit
 - Fifteen reflections from primary to detectors
 - Optimized coatings for 0.6-2.4 microns
- Large number of telescopes
 - Optimized for model-independent imaging



Walk through the Optical Path





Unit Telescopes

- Designed/built by AMOS
 - 1.4m aperture
 - afocal alt-alt design
 - polarization preserving

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- 62 nm rms wavefront
- UTI completed factory acceptance testing
- UT2-3 ordered





UT Enclosures

- Designed by EIE
- Build in "award"
- Houses and transports UTs
- Allows close-packed configuration to 30 deg elevation without vignetting for 6 hour tracks







Foundations and Beam Transport

- Designed M3 and built by MRO
- Supports 3 UTs per beamline with 0.5 mbarr vacuum from UT to BCA
- Install for piers for inner array began this past summer
- Houses all components of automated alignment system

Inner Array Install





03/01/2011

Beam Combining Facilities

- Design by M3/built KL House – delivered in 2008
- Thermal & vibrational stability
- Supports full array
- Single-pass DL section 190 m long
- Equipment install started 2010





Delay Lines

- Designed/built Cambridge
- Innovative approach
- Inductive pick-up & wireless communications
- Install begins this month











Fringe Tracker



Science Instrument

- MRO conceptual design
 SIRCUS J,H,K with R~30 and 300
- 4-way image plane combination with fastswitching to combine 6 beams in ~100 sec

Mag	J	Н	К
13	0.45	0.54	0.53
П	17.6	20.8	18.4
9	195	207	159

Performance: SNR per spectral channel in 100 sec at R~30 with 0.7" seeing and RN=5e-

A Science Meeting





Scientific Schedule for MROI

- Technical Phase Key observations that quickly demonstrate technical competencies
- Science Phase Scientific observations that produce transformational changes to understanding of astrophysical phenomena
- Open Time Phase Release of facility to broader community through public funding



Interferometry Workshop in NM

- MRO/NOAO/LANL/USIC co-sponsors
- March 28-31, 2011 in Socorro
- Wide variety of science topics presented and discussed over 4 days
 - Invited speakers including:
 - C. Haniff, R. Genzel, A. Speck, J. Monnier, D. Segransan, T. Storchi-Bergmann, M. Delbo, T. Metcalfe, J. Eisner, G. Torres, J. Aufdenberg, P. Young
 - Poster sessions, tour of MROI, Conference Proceedings
 - Website: www.mro.nmt.edu/workshop



Thank you for your attention!

- <u>PI</u>:Van Romero
- <u>Deputy PI</u>: R. Cervantes
- <u>Advisor:</u> D. Westpfahl
- <u>Prog. Director</u>: I. Payne
- <u>System Architects</u>: C. Haniff, D. Buscher
- <u>Proj. Scientist</u>: M. Creech-Eakman
- <u>Proj. Manager:</u> R. Selina



- <u>NMT Team</u>: V. Alvidrez, C. Anderson, M. Apodaca, L. Archuleta, D. Brown, A. Farris, E. Iker, S. Jimenez, C. Jurgenson, R. King, D. Klinglesmith, T. McCracken, J. Murphy, K. Nyland, A. Olivares, J. Pino, M. Richmond, C. Salicido, J. Sanchez, J. Sandoval, F. Santoro, J. Seamons, A. Shtromberg
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