Update on VISION at NPOI And Completely Unrelated Collaboration Opportunities

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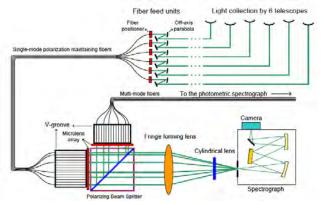
VISION: Visible Imaging System for Interferometric Observations at NPOI

- ► Construction: NSF MRI R² (American Reinvestment and Recovery Act) Funded
- ▶ Improve NPOI Data Product Precisions by $10 \times$
- Improved System Visibilities: Low Visibility Targets
- Avoid Systematic Effects:
 Delay Nonlinearities, APD Afterpulsing
- Monitor all baselines
- Spectral Resolution Flexibility
- Modernize Equipment

Design

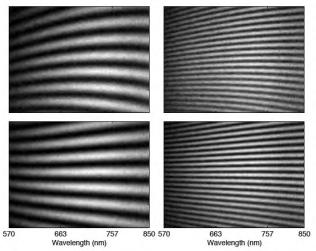
VISION Design

MIRC, With Modifications



VISION in the Lab

Lab Fringes



VISION in the Lab

VISION Design



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- ► Imaging Stellar Surfaces
- Emission Line Studies: High Spectral Resolution
- Baseline Bootstrapping: Beyond the first visibility zero
- Defending Democracy, Deterring Aggression, and Maintaining Freedom of the Seas (Imaging Satellites)
- Triple Star Systems



Complementarity and Divergence With MIRC

- Combined IR-Visible Studies
- VISION: Photon Counting, Baseline Bootstrapping, baseline flexibility
- MIRC: Sensitivity Advantage—IR/Telescope Size (until 1.8m....)
- High-Contrast Objects
 - Separated Objects: Want Lower Contrast
 - ► Faint is redder than Bright: IR (BD's, Planets)
 - ► Faint is bluer than Bright: Visible (CVs, Thanks Fabien)
 - Embedded Objects: Want Higher Contrast (Star Spots, Visible)



Fairborn Observatory: Fully Robotic Photometry and Spectroscopy



Automatic Photometric Telescopes (APTs)

- Operations:
 - ▶ 6 Telescopes (0.4, 0.75, 4 × 0.8 m)
 - Single Pixel (PMT), 2-channel (b-y), photometers
 - Sequential observations of calibration/check stars
- Benefits:
 - Nearly Maintenance Free (this year, OS updates during monsoon season, and a few mouse-chewed cables replaced).
 - ▶ Precision: 1 milli-mag/obs; 0.1 milli-mag seasonal repeatability
 - Long-term stability: 0.4m APT in operation since 1987
 - Cost-efficient: Under \$30k/year/telescope
- Tradeoff
 - Not versatile
- ► Also: Red-sensitive 14" Imaging (CCD) Telescope

VISION Overview

Robotic Spectroscopy

Automatic Spectroscopic Telescope

- ▶ 1 Telescope, *D* = 2 m
- Fiber-optic fed instruments (recent upgrades)
- ► Primary Instrument: R~30,000 Optical Echelle, 3800–8200 Å
- ▶ Also: Ca H/K spectrograph, Mini-Echelle
- Future Instruments: UF's FIRST, EXPERT. NSF-Funded "EDI-Testbed".
- Nearly Maintenance Free (2-3 trips/year)
- Partially Versatile (multiple instruments)



Contact



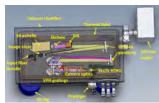
Robotic Spectroscopy

AST Complex, September 2012

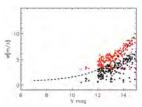


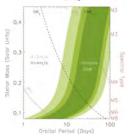


Future: FIRST



- $R = 68,000 (1.4-1.8 \ \mu m), R = 56,000 (0.8-1.35 \ \mu m),$
- M Dwarf Planet Survey
- Young Star Planet Survey
- ► SB1 → SB2
- Cold Stellar Atmospheres
- ► Young Stars





Observations by TSU's Fairborn Telescopes are generally available on a collaborative basis.

- For photometric measurements with the Fairborn APT's and new Imaging Telescope, contact:
 Greg Henry, gregory.w.henry@gmail.com
- For spectroscopic measurements from the Fairborn AST, contact:
 Matthew Muterspaugh, matthew1@coe.tsuniv.edu
- ► For technical questions, please contact: Michael H. Williamson, michael.h.williamson@gmail.com