PAVO Science Update

Daniel Huber (IfA Hawaii), Tim White (Aarhus), Mike Ireland (ANU), Peter Tuthill (Sydney), Iva Karovicova (Heidelberg)















PAVO Basics

- 3-Telescope Beam Combiner
- $\lambda \sim 0.6$ -0.8 um in ~20 channels
- Spatially modulated pupil-plane fringes
- Best performance: $R \sim 8 \text{ mag}$
- No major hardware upgrade since ~2009







PAVO Software

- Reduction & analysis tools available (on cvs ... to be ported to Github soon)
- 2T: essentially the default operation mode
- 3T: bias reduction issues; we welcome collaborators willing to contribute to reduction efforts (chopper data available)

http://pavo.wikispaces.com/













TULAR RESE.		CHARA	2017: Year 13 Science Review – Adaptive	Optics and Open Access
2	016	Gordon	Gies, Schaefer	Fundamental Properties of O- and B- Type Stars
		Huber	White, Collet, Ireland, Tuthill, Bedding	Measuring Limb Darkening at visibile wavelengths with PAVO
	P3	Ireland	Karovicova, White, Gilmore, Jofre, Worley, Hourihane, Lind, Bergemann, Heiter, Wittkowski, Asplund, Chiavassa, Creevy, Thevenin	Interferometric Observations of Benchmark Stars for calibrating large stellar surveys of the Milky Way
	P4	Quinn	Jones, White	Open Cluster Ages from Sizes of Giants and A Stars
	P5/C4	Boyajian/vonBraun	ten Brummelaar, Farrington, McAlister, Gies, van Belle, White, Jones, Ireland, Huber, Fischer	Diameters and Temperatures of Main Sequence FG Stars (PAVO/Classic)
	P6/C5	Huber	White, Ireland, Tuthill, Bedding, Stello, Boyajian, Aguirre, Creevy, Nardetto, Mourard	Paving the way for Galactic Archeology: Angular diameters of oscillating red giants(PAVO/Classic, VEGA/Classic)
	P7/CL3	Jones	White, Boyajian, Schaefer, Baines, Ireland, Patience	How Old are the nearest A-stars?
	P8/C6	Tuthill	Richichi, Schaefer, ten Brummelaar, Ridgway, Ireland	Pilot Study for Lunar Occultations at CHARA (PAVO/Classic)
	P9/C7	vonBraun/Boyajian	van Belle, Henry, Kane, Schaefer, Giguere, Fischer, Brewer. Wana. Huber. Ireland. Pewett. Ciardi. Mann.	Radii of late type dwarf, exoplanet hosts, and exoplanet host candidates
	P10/M8	White, R	Baron, Martens, Egland, Monnier, Roettenbacher, Jones	Star Spots and Radii of Solar Analogues
20	17A			
20		Gordon	Gies, Schaefer	Fundamental Properties of O- and B-type stars
	P2	Huber	White, Collet, Ireland, Tuthill, Bedding	Measuring Limb-Darkening at Visible Wavelengths with PAVO
	P3	Ireland	Karovicova, White, Gilmore, Jofre, Worley, Hourihane, Lind, Bergemann, Wittkowski, Asplund, Chivassa, Thevenin	Interferometric Observations of benchmark stars for calibrating large stellar surveys of the Milky Way
	P4	Jones	White, Boyajian, Schaefer, Baines, Ireland	How old are the nearest A-stars
	P5	Quinn	Jones, White	Open Cluster Ages from Sizes of Giants and A-stars
	P6/C2	Tuthill	Richichi, Farrington, Schaefer, ten Brumelaar, Ridgway, Tasuya	Lunar Occultations at CHARA
	P7//V10	Huber	White, Creevy, Boyajian, Ireland, Tuthill, Bedding, Stello, Aguirre, Nardetto, Mourard	Paving the way for Galactic Archeology: Angular diameters of Oscillating Red Giants
	P8/C3	Boyajian/vonBraun	ten Brummelaar, Farrington, McAlister, Gies, van Belle, White, Jones, Ireland, Huber, Fischer	Diameters and Temperatures of Main-Sequence FG stars
	P9/C4	von Braun/Boyajian	van Belle, Henry, Kane, Schaefer, Giguere, Fischer, Brewer, Wang, Huber, Ireland, Pewett, Ciardi, Mann, Feiden	Radii of late type dwarfs, exoplanet hosts, and exoplanet host candidates

PAVO Science: Diameters for Galactic Archeology

White, Huber, Karovicova, Ireland, Bedding + VEGA Team

















Galactic Archeology

→ Asteroseismology (Kepler/K2, TESS)

→ Spectroscopic surveys (APOGEE, GALAH, RAVE, Gaia-ESO)

but ... all hinges on fundamental calibration: radii for asteroseismology, T_{eff} for spectroscopy!











CHARA 2017: Year 13 Science Review – Adaptive Optics and Open Access

Bright Kepler/K2 Stars







Interlude: Cross-Beam Combiner Calibrations















EXETE











nature astronomy

LETTERS PUBLISHED: 2 MARCH 2017 | VOLUME: 1 | ARTICLE NUMBER: 0056

Two massive rocky planets transiting a K-dwarf 6.5 parsecs away

Mass $(M_{\odot})^*$ Radius $(R_{\odot})^{\dagger}$ Effective temperature (K)[†] 0.81±0.03 0.778±0.005 (0.6 %) 4,699±16 (0.3 %)



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Effective temperature (K) ⁺	4,699±16	(0.3 %)

Solution: dedicated (simultaneous?), multibeam combiner observations of the same target/calibrators?

(mas)



(mas)





PAVO Science: Limb Darkening

White, Huber, Ireland, Tuthill, Bedding



















Optics and Open Access

- 4 solar-type stars with 2nd lobe 2T fringes
- systematically lower limbdarkening than expected from 1D atmospheres

White et al., in prep







PAVO Science: Lunar Occultations

Tuthill, Richichi, Farrington, ten Brummelaar, ...





















CHARA 2017: Year 13 Science Review – Adaptive Optics and Open Access



Georgia State University













EXETER



Summary

- 2T PAVO routine, 3T progress stalled by lack of (human) resources
- Cross-Beam Combiner calibrations should become priority in open access CHARA era
- Key Science (4+ papers in prep):
 - Asteroseismology (White/Huber)
 - Gaia Benchmarks (Karovicova/White)
 - Limb Darkening (White/Huber)
 - Lunar Occultations (Tuthill/Richichi)













