

### Long baseline interferometry in the visible: Recent progress on FRIEND project

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## Outline

- Short presentation of FRIEND prototype
- Coupling efficiency and limiting magnitude
- Instrumental visibility
- Closure phase
- Adaptive optics and injection















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# Short presentation of FRIEND prototype



Recombination module



 $<|V|^2>=\frac{< E_{HF}>}{\int_{\Lambda\lambda}\kappa(\lambda)\int_{\Lambda x}< P_1(\lambda,x)P_2(\lambda,x)>dxd\lambda}$ 







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#### Observations

#### 2016-09-29

- E1E2W2
- $r_0 = 10 \text{ cm}$
- 1 Cal:
  - HD3360: mR = 3.74
- 1 Sci:
  - HD5394: mR = 2.32

#### 2016-09-30

- E1W2W1
- r<sub>0</sub> = 14 cm
- 3 Cal:
  - HD3360: mR = 3.74
  - HD2905: mR = 4.02
  - HD3240: mV = 5.076
- 2 Sci:
  - HD5394: mR = 2.32
  - HD11415: mR = 3.4



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# Coupling efficiency and limiting magnitude



 Limiting magnitude defined by flux in photometric channels













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# Coupling efficiency and limit magnitude

- $r_0 = 14 \text{ cm} @ 550 \text{ nm} => \rho_{\text{expected}} = 2,6\%$
- Measured flux and magnitude
- t<sub>CHARA</sub> = 1.7%
- $t_{\text{FRIEND+OCAM}} = 31.5\%$

#### => p<sub>estimated</sub> = 2.8% with no adaptive optics













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# Coupling efficiency and Imiting mag defined by photometric channels

- For faint stars, flux at 0 or less
- Photometric sorting:  $P_i > n \sigma_{dark}$ , n = 1,5





# Instrumental visibility

 Optimization : OPD rejection method (low spectral resolution)

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# Instrumental visibility 2016-09-29 2016-09-30





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- Stability of Base 1 on 09-29
- Determination of V<sub>instr</sub> stable with or without HD2905 (relative variation between fits < 1%)
- V<sub>instr</sub> does not seem night-dependent

	09-29	09-30
Base 1	$0.72 \pm 0.05$	$0.82 \pm 0.12$
Base 2	$0.62 \pm 0.05$	$0.63 \pm 0.05$
Base 3	$0.83 \pm 0.06$	$0.73 \pm 0.27$

Table 3: Estimation of  $V_{instr}^2$  per base and per night.













# Instrumental visibility

- Decrease on other bases: due to alignment drifts of pupils ?
- Stability with respect to atmospheric condition:

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- No strong variation of the seeing during and between the nights
- Systematic errors unseen ?
- Keep investigating



### Closure phase

 $Re(B_0)$ 

- Bispectrum:  $B_{0,u,v} = B_{1,u,v} \alpha \left( |C_u|^2 + |C_v|^2 + |C_{u,v}|^2 \right) + \beta N$
- $\alpha = \arg \min_{\alpha} \left( Var(Bg(B_{1,u,v})) \right)$ , Bg=background Re(B<sub>1</sub>) Re(B<sub>0</sub>)



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• Closure phase:  $CP = \arctan\left(\frac{\Sigma_{Im}}{\Sigma_{Re}}\right)$ 

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•  $\sigma_{\Sigma} = \sqrt{N} \sigma_{bg}$ ,  $\sigma_{CP}$  determined by Monte-Carlo





#### Closure phase

- 2016-09-29:  $slope = -0.01 \pm 0.01$
- 2016-09-30:  $slope = 0.16 \pm 0.06$  with 1 cal  $slope = 0.14 \pm 0.06$  with 2 cal => Stability of  $CP_{instr}$  with respect to the calibrators => Investigating on possible instrumental effects















### Adaptive optics and injection Observations

(with T.Ten Brummelaar, M.Anderson, J.Sturmann, L.Sturmann, C.Farrington, N. Vargas, VEGA team)

#### 2017-03-08

- S2
- $r_0 = 7 \text{ cm}$
- Target: Sirius
- AO modes:
  - Labao
  - Open loop
  - Labao + own TT
  - Fast TT

#### 2017-03-10

- S1S2
- $r_0 = 8 \text{ cm}$
- Target: Procyon
- AO modes:
  - Labao
  - Open loop













#### Adaptive optics and injection



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2017-03-08

 $<\rho>_{open \ loop} = 0.65\%$   $r_0 = 7 \ cm \Rightarrow \rho_{th} = 0.65\%$   $\rho_{labao} = 0.76\% \Rightarrow$   $r_{0,eq} = 7.43 \ cm$ 

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2017-03-10

 $r_0=8\ cm \Rightarrow \rho_{th}=0.88\%$ 

$$<
ho>_{S2,open\ loop}=0.58\%$$

$$<
ho>_{S1,open\ loop}=0.50\%$$



# Conclusion and prospects

- Limiting magnitude: characterized and quantified
- Instrumental visiblity:
  - Stable with respect to magnitude and night-dependent
  - Very sensitive to pupil drifts?
- Closure phase:
  - First measurements of CP quite accurate
  - Keep characterizing (magnitude, AO)
- AO and Injection:
  - AO improves a bit injection in FRIEND
  - Correlate FRIEND and AO data
- Martinod et al 2017 in prep.











