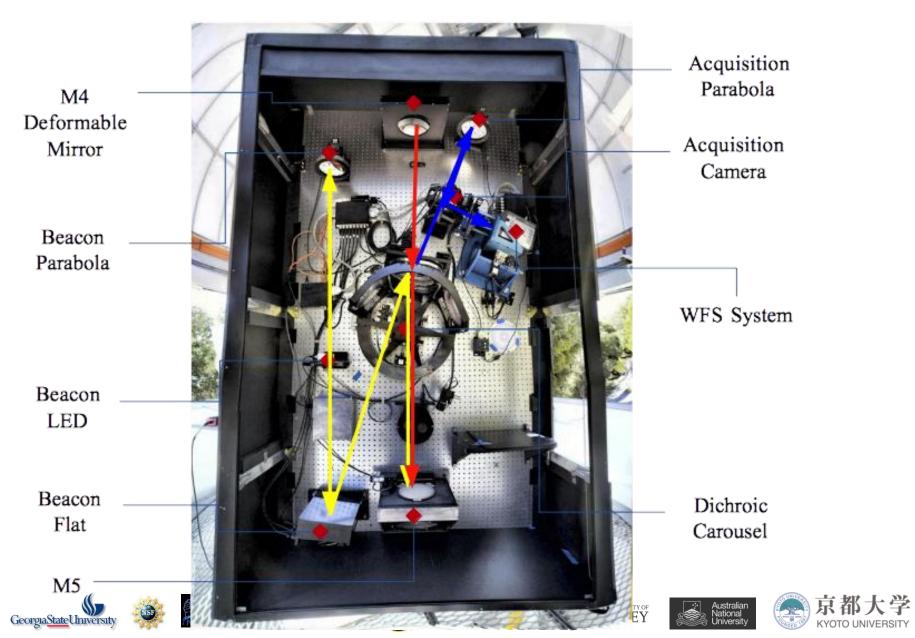


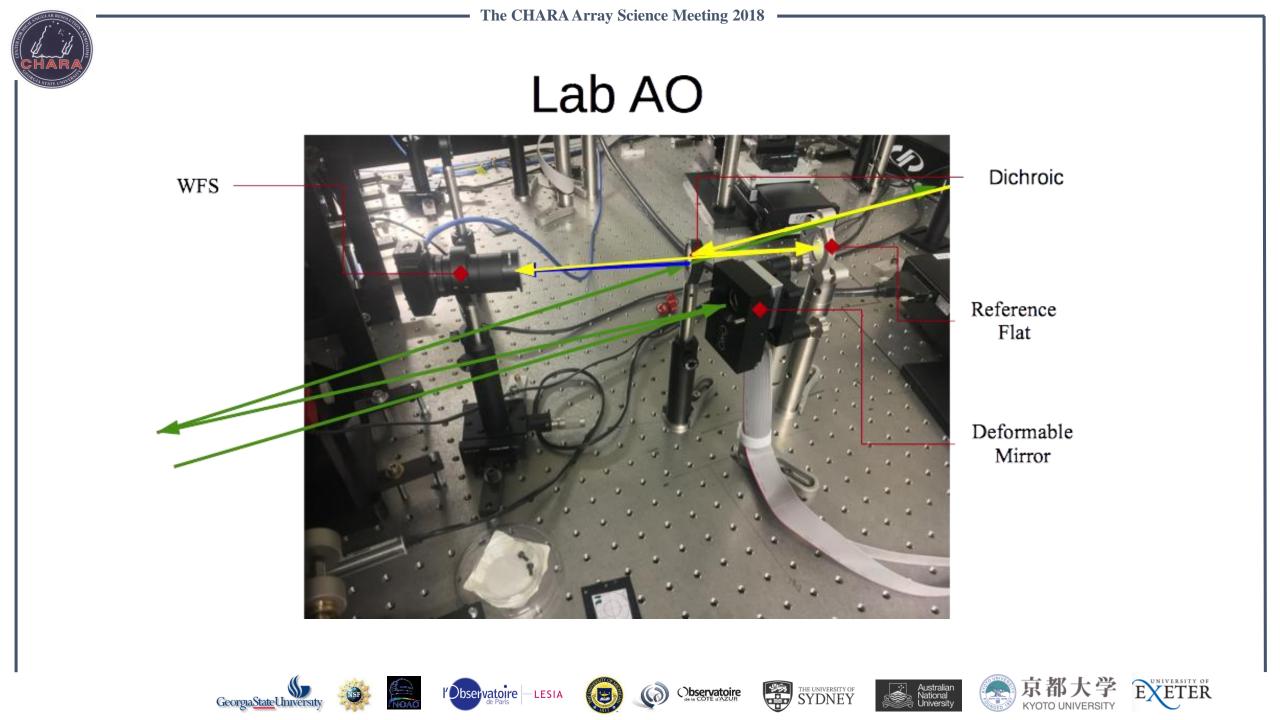


AO at the Telescopes

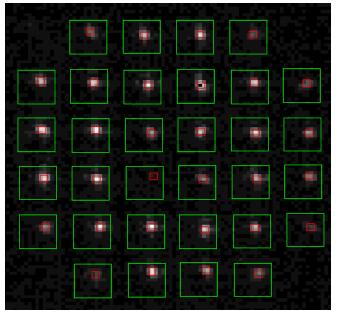


Your Talk Title Here



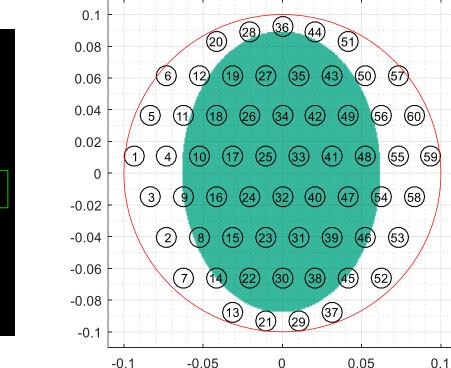


— The CHARA Array Science Meeting 2018 — WFS & DM Specifics



Labao

- Deformable Mirror
 - 37 Actuator OKO
 - 15mm Diameter
 - Max deflection: 9 micron
- Wavefront Sensor
 - 0.15mm pitch



- Deformable Mirror
 - 60 Actuator ALPAO
 - 125x177mm (Active Area)
 - Max deflection: 29 micron
- Wavefront Sensor
 - 0.11 mm pitch







Tao







ETER

Phase | Status













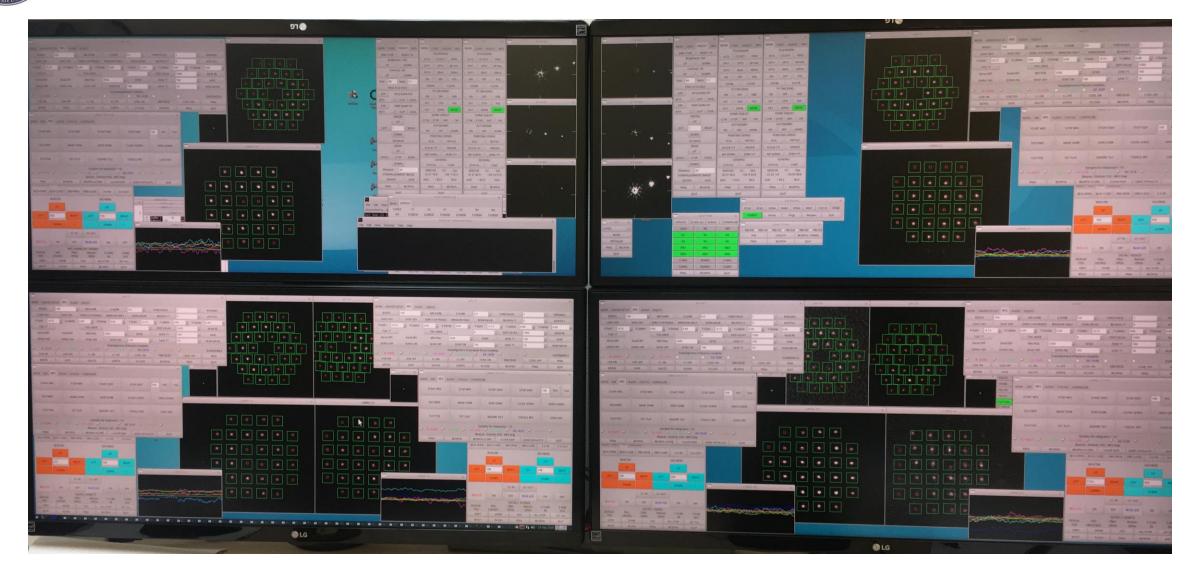




CHARA

The CHARA Array Science Meeting 2018

Phase I Complete











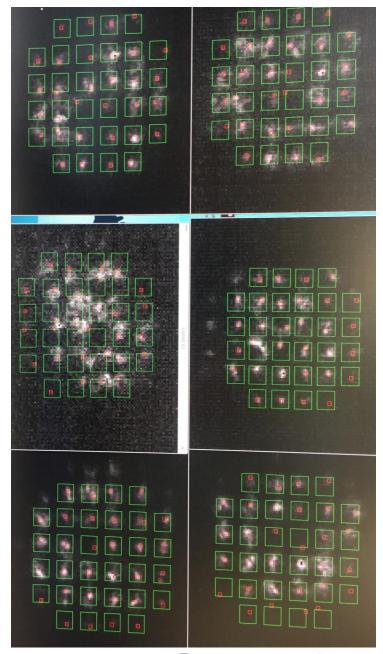






LABAO Current Situation

- Labao used daily for alignment
 - Remote
 - Objective
- Star-Flat Procedure routinely used
 - Flat is measured on nearby bright star (V<4)
 - Data is taken
 - New flat measured when moving to new target
 - Current procedure is on the CHARA wiki
 - /Operating Procedures/Using Labao with Starlight
 - Good results for MIRCX
 - Poor results for CLIMB
- Star-Flat Procedure reveals static aberration sources





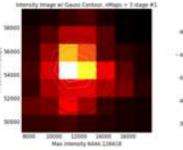


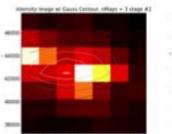




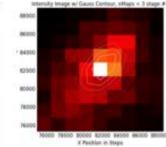


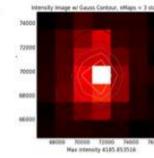
Labao Improvements for MIRCX

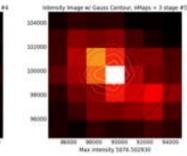


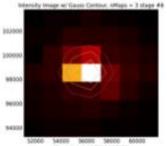






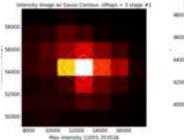


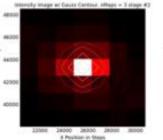


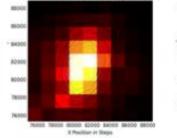


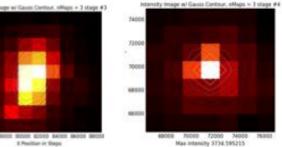
Max interaity 17882.837991

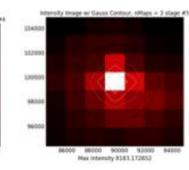
internity image a/ Gauss Century, nMaps = 3 stage 4

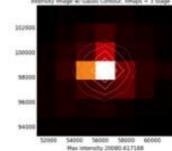












- Closed loop Labao on sky
 - V < 2.5
- Correction speed: 40Hz
- Lab tip-tilt guiding •
- DM is saturating
 - A reason for optimism!

Your Talk Title Here









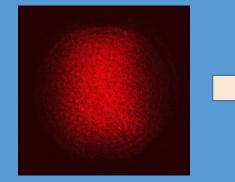






ACTUATOR VOLTAGES vs DM SURFACE

All Lab-DMs were tested with a Michelson setup in the lab before installation J. S. June 2015 Actual wave-fronts reflected from a Lab-DM





0 V applied to all 37 actuators,
~ one fringe across → ~ flat surface

GeorgiaStateUniversity

Uniform V applied to all actuators → smooth concave surface

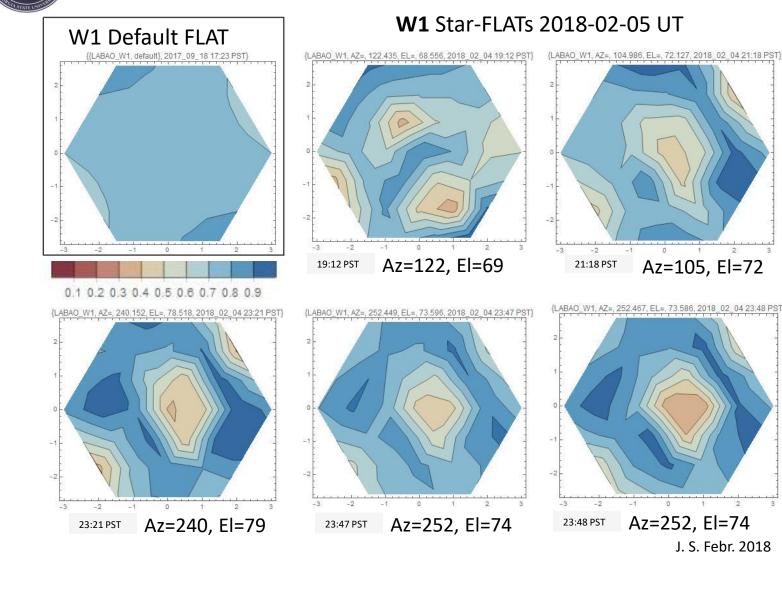
Observatoire

THE UNIVERSITY OF

The following contour plots show the actuator voltages of the default FLAT and Star-FLATs on a single night.

Observatoire LESIA

Applied voltages are scaled from 0 to 1. Contours shown in 0.1 increments. The DM actuator response is some monotonic function of the applied voltages. **Sets of applied voltages relate to** DM shapes, and in turn relate to the **aberrations** being corrected.



GeorgiaStateUniversit

The aberrations being corrected do not seem to rotate on the labwfs, and fairly constant throughout the night.

Only the focusing looks a bit off somewhere.

National

THE UNIVERSITY OF

2

J. S. Febr. 2018

Az=252, El=74



-2

Az=105, El=72



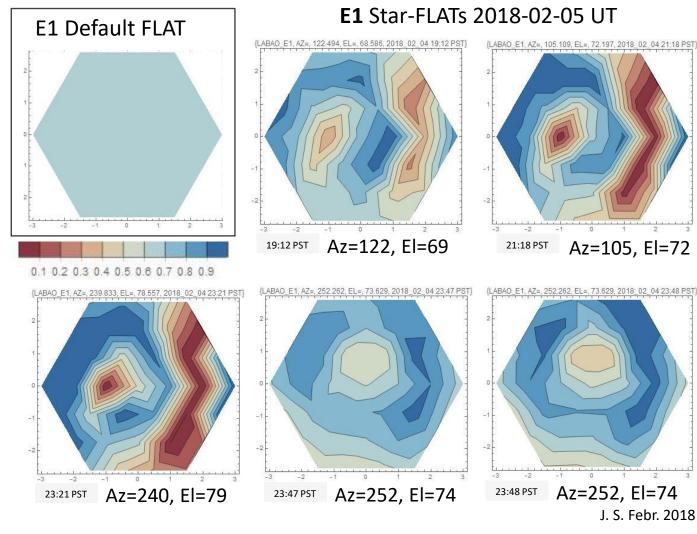


ETER

KYOTO UNIVERSITY







The aberrations being corrected do not seem to rotate on the lab-wfs, no correlation with telescope AZ indicates negligible telescope astigmatism.

Possible correlation with **OPLE** cart position. This looks like beam shear at some places, which may be fixed by better OPLE rail alignment.



Observatoire

THE UNIVERSITY OF









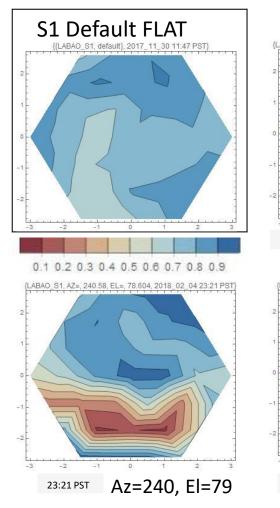




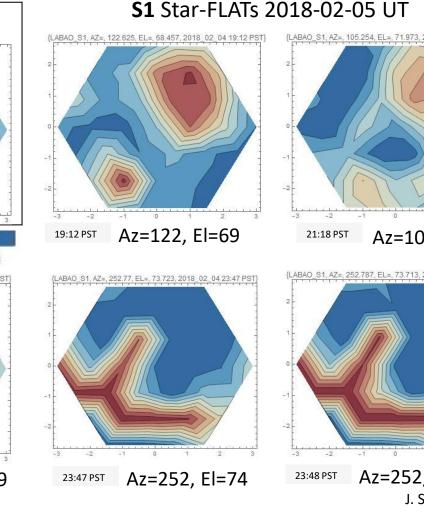


GeorgiaStateUniversit



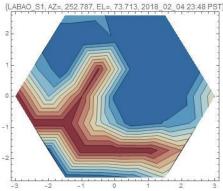


GeorgiaStateUnivers



vatoire

Az=105, El=72



Az=252, El=74 J. S. Febr. 2018

Observatoire

DM corrections seem extreme in all cases, caused probably by multiple problems:

- The extreme ridges may be due to **vignetting** seen by LABAO wfs.

- There could be problems with the LABAO wfs itself, this one was left untouched for the longest time.

- DM shapes seem to rotate with telescope AZ suggesting telescope astigmatism.

Your Talk Title Here







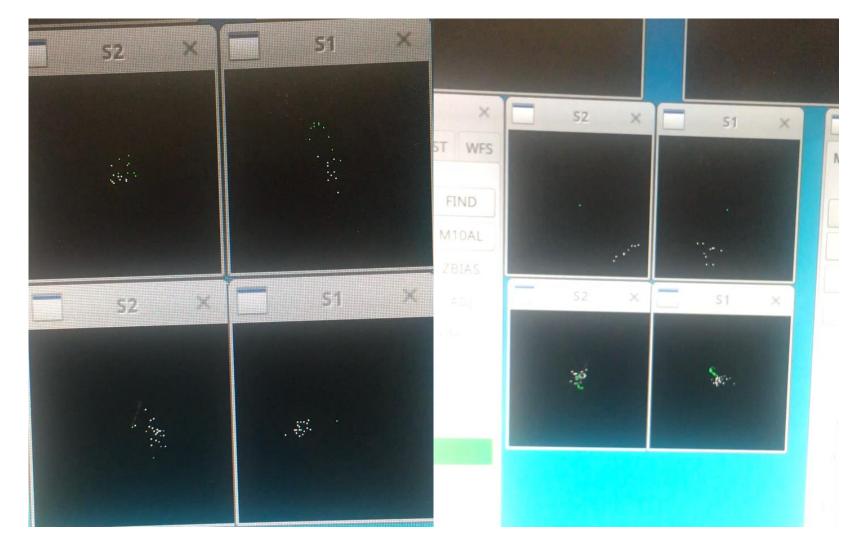






Telescope Based Tip-Tilt

- TWFS Tip-Tilt works ۲
- TWFS offset from lab •
 - Auto M7 •
- Programs can use different Tip-Tilt settings
- Telescope Tip-Tilt has been used to V \sim 11.5 (NGC 4151)















14

Phase II Status







 (Ξ)

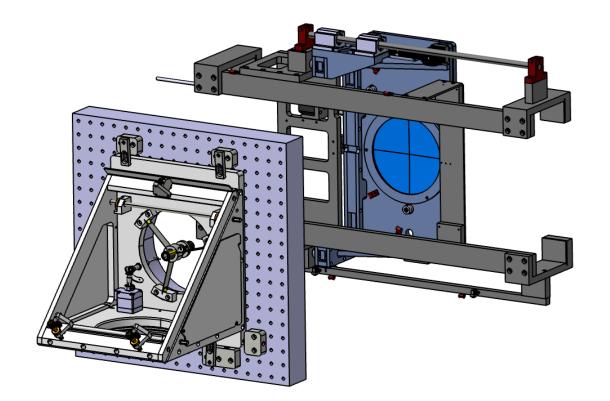








Phase II Telescope Hardware Changes / Additions



- All telescopes require
 - mounts for the DMs
 - DM calibration light sources
- DM mounts
 - Kinematic adjustments
 - "Dummy DM" flat mirror
- DM calibration light source
 - Fiber fed light source
 - Measure DM reconstructor
 - Kinematic Adjustments
- All other hardware is already in place per Phase I



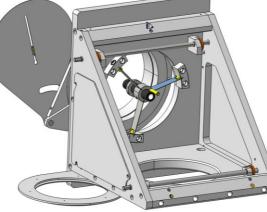


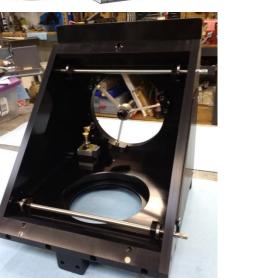




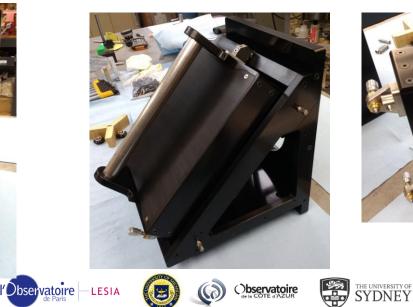


A prototype mount for the DM and a "dummy" DM are in hand.

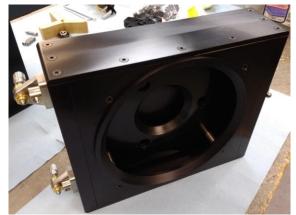




Georgia<u>State</u>University



1



"dummy" DM









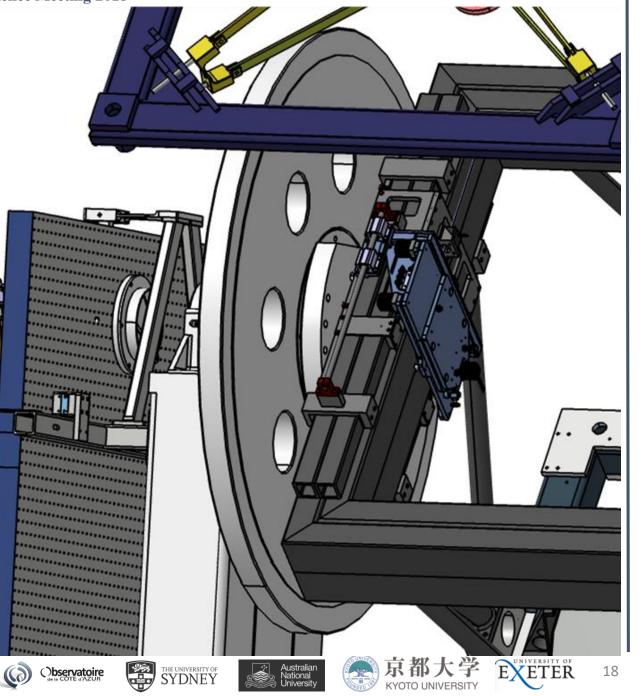
17

DM Calibration Source

- Clamps to telescope midframe
 - Midframes vary up to 0.25" (6.3 mm) •
- Actuators for Tip-Tilt / Focus
- Must clear beam path
 - Linear actuator •
 - < 2" (50mm) from midframe

GeorgiaStateUnivers

Prototype has shipped from Atlanta





servatorre

The first Telescope DM Arrives

Observatoire LESIA



GeorgiaStateUniversit

















Moving Forward

- Prototype DM is onsite
 - Testbed is ready ٠
 - Waiting on software and time ٠
- S2 will be ready for on sky testing with the DM soon ٠
 - M4 Mount being installed •
 - Calibration source nearly ready ٠
- E1 is likely the 2nd telescope to get a DM ٠
 - Recoating in May ٠
- Others to follow in schedule dead time •
- Ongoing AO engineering time has been integrated into the observing • schedule
 - Thursday Nights
 - AO engineering runs •
- There is much software development to do ٠
 - GUI Development
 - Automation Development ٠







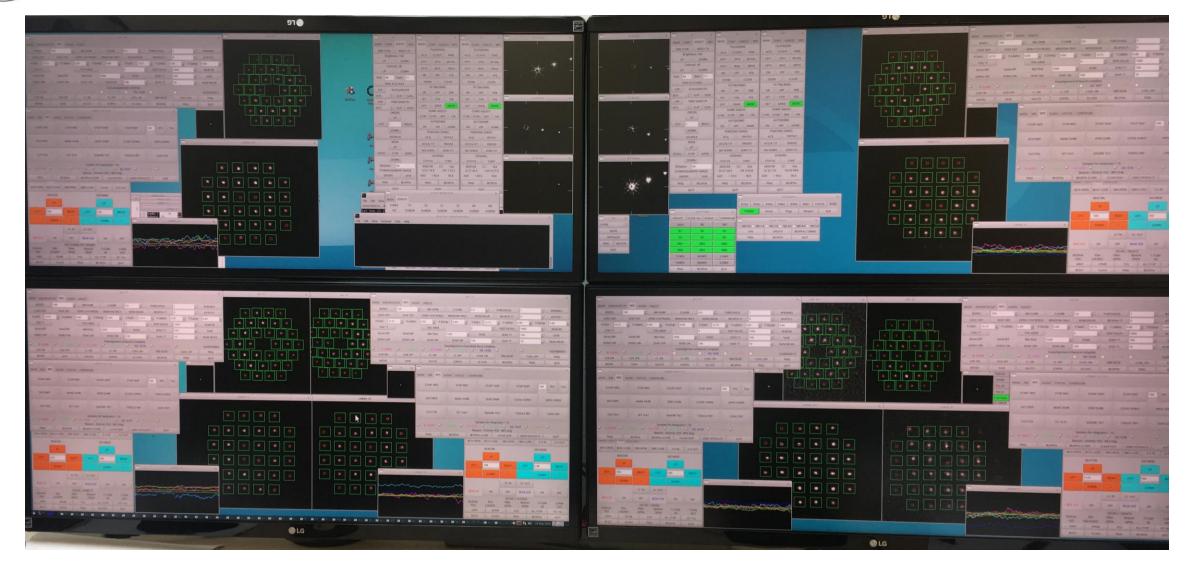








The Great GUI Problem



6

Observatoire

Observatoire - LESIA

THE UNIVERSITY OF SYDNEY

Australian National University



GeorgiaStateUniversity