Characterization of the new e-APD camera of MIRCX

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Bias voltage in the multiplication region accelerate the electron ⇒Collision ionization ⇒Avalanche effect



Characterized by :

- Avalanche Gain
- Excess Noise Factor (ENF)













Camera characteristics

- Instrument MIRCX
- Developed by First Light Imaging (FLI)
- H band

Manufacturer characteristics :

- Avalanche gain up to 300
- System Gain : 0.77 ADU/e-
- ENF around 1.3
- FPS max of 3.5 kHz
- Dark current < 200 e-/pixel/s
- Read out noise < 1 e-/pixel/read

















Photon counting : statistical model

Simulation input :

- Gain = 150 ADU e-
- Flux = 1.5 e-/frame
- Background noise = 18 ADU





- At low ENF : separated peaks
- At high ENF : not separated peaks

Your Talk Title Here

















normalized histogram for statistic model, ENF = 1.09316445616



Calibration of unexpected behavior

- Extract temporal sequence of one pixel (photometric channel)
- Subtract median value of several not illuminated pixels of the same line ⇒Take off the sinusoidal signal
- Build histogram of the temporal sequence values





- No clearly separated peaks => no photon counting
- A break between 0 photon events and the others at low flux
- Background histograms not symmetrical













Photon distribution

MIRCX



Poisson distribution for the measured flux not consistent with histograms : Simulation 10% -> data ≈ 0%

> As F (e-) = F (ADU) / G \Rightarrow Total gain is false

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Camera modelization

Free parameters :

- Gain (G)
- ENF
- Flux (F)

Objective : Fit illuminated data histograms

Determination of F by the ratio between 0 photon events and total events



Determination of G by F and the mean of the histogram (H) : $H = G \cdot F$

Determination of ENF by the spreading of the histogram

















1) Representation of each p photon events by a Dirac at the position G . p

2) Convolution of the Dirac with M, p times for an event of p photons : Where M is the Gain distribution 0 photon => Dirac in 0
1 photon => Convolution of a Dirac in G with M
2 photons => Convolution of a Dirac in 2G with M, then a convolution of the result with M

3) Normalization of p photon events with the corresponding rate of the Poisson distribution of a given F

4) Addition of the different p photon events

5) Convolution of the result buy the Background histogram















Gain distribution

Gain distribution (M) : **Define Gain and ENF** Model dependent

Spread : $G = \langle M \rangle$ Shape : ENF = $\frac{\langle M^2 \rangle}{\langle M \rangle^2}$

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Chi2 map



- Exploration of the parameter space to perform χ^2 maps
- A single solution found in the parameter space

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Chi2 map



- Degeneracy of solution at high flux (F > 3 e-/frame/pixel)











Observation of the different parameters



- Result consistent for the same requested gain at different flux
- Some outliers for ENF, but from unconstrained high flux
- ENF ≈ 1.6

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Photon distribution



Poisson distribution for the measured flux not consistent with histograms ⇒Total gain is false

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Photon distribution



Poisson distribution for the flux from the model is consistent with histograms

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Discussion and perspective

Discussion :

- Gain distribution dependence of the model
 - Test different Gain distribution
- High flux result not constrained
 - Combination CHI2 cubes to constrain high flux result

Perspectives :

- Study massive stars with this new camera
 - Observations already began : faintest star H = 6.5 (previous limit H = 5.5)
- Use the model to try "photon counting" on sky data
- Sub Poisson noise may have a physical explanation => Fano Effect ?

Conclusions

- New model works :
 - Obtain characterizations that explain the data
 - Validation of the characterizations by an independent method
- Validation of a real Gain of half the expected Gain
- ENF of 1.6, unexpected
- Characterization still good enough to improve previous MIRC performance
- Classical photon counting impossible
- Further study needed to explain unexpected behavior of the camera (sub Poisson noise)













