

VLTI update

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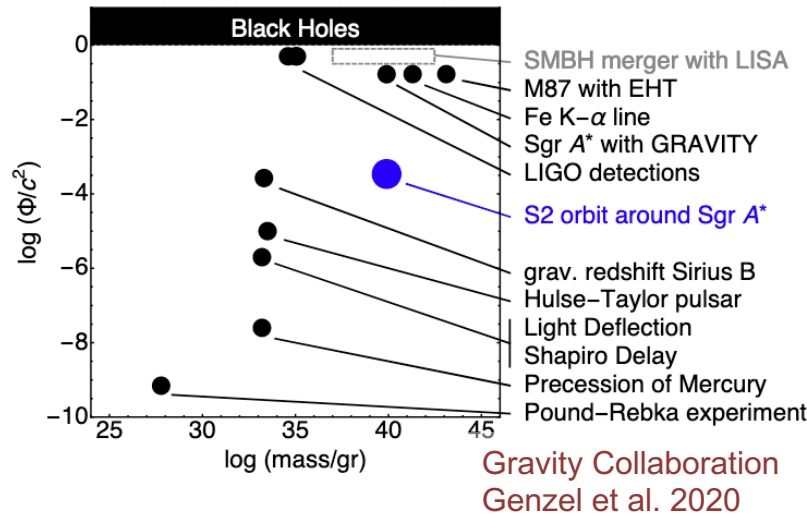
Impact of COVID

- Operations highly disturbed
 - Paranal is a remote site (commuting)
 - We lost a semester
- Travel EU -> Chile not possible
 - All science visitors observe remotely (eavesdropping)
 - Some technical activities / commissioning cancelled
 - Larger remote control room in Garching for commissioning activities

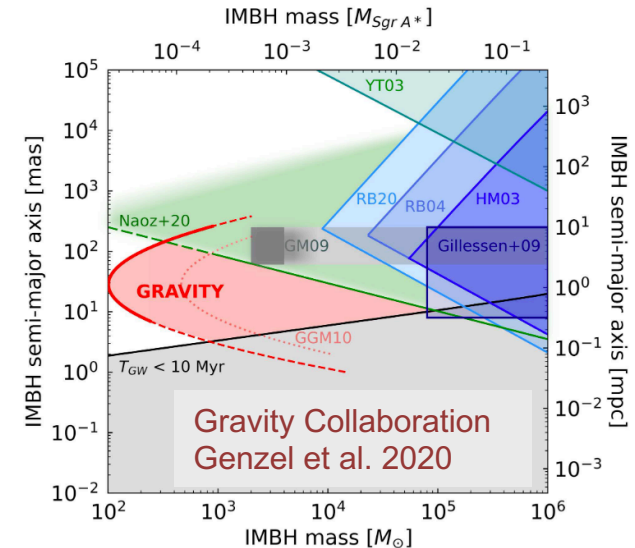


Galactic Centre

■ S2 precession as test of relativity

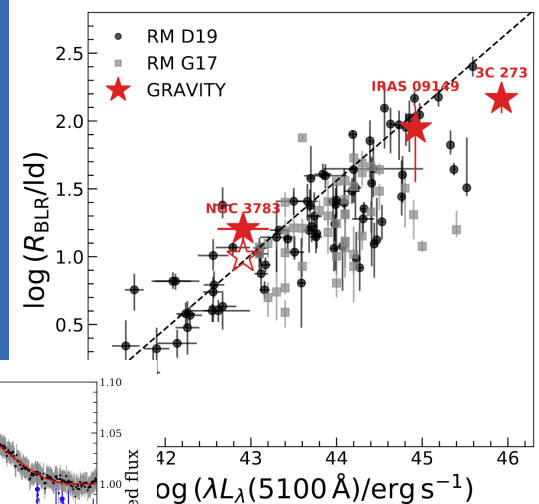
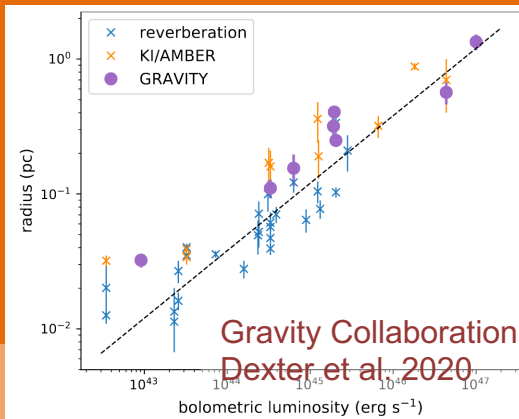
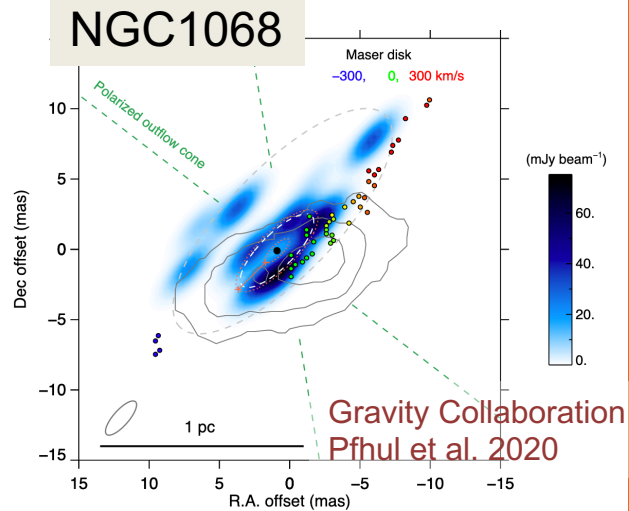


■ Excluding other BH close to Sgr A*



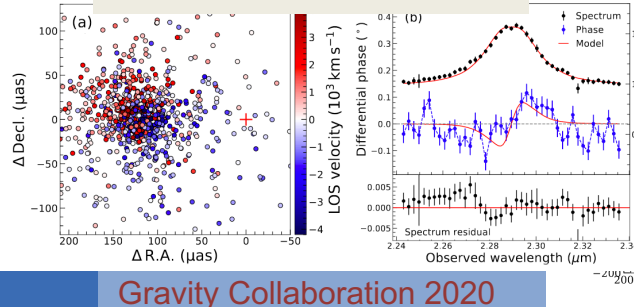
Extragalactic science

AGN



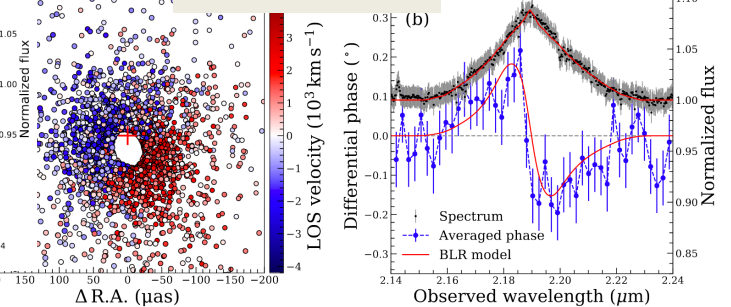
QUASARS

IRAS 09149-6206



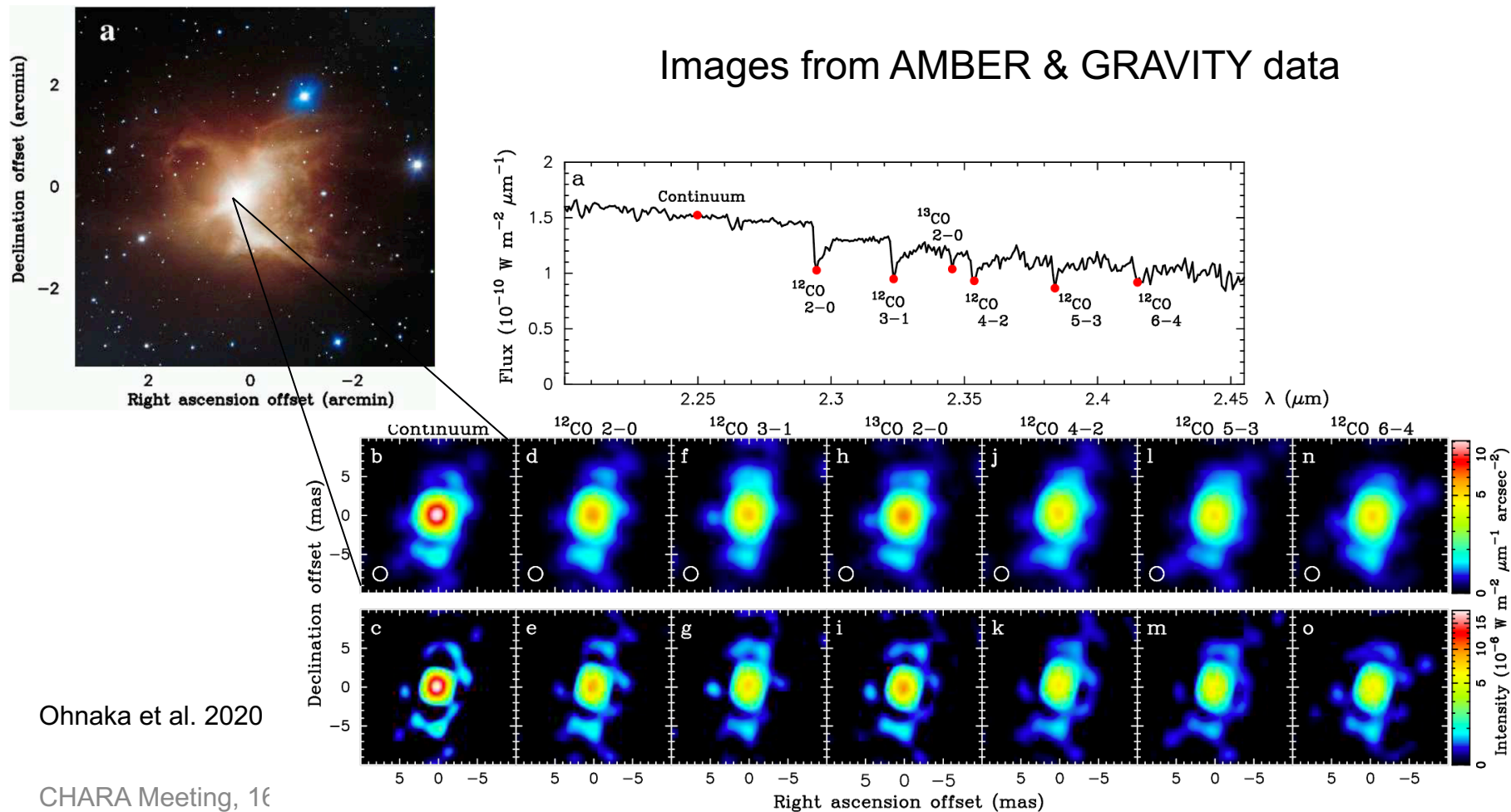
Gravity Collaboration 2020

NGC 3783

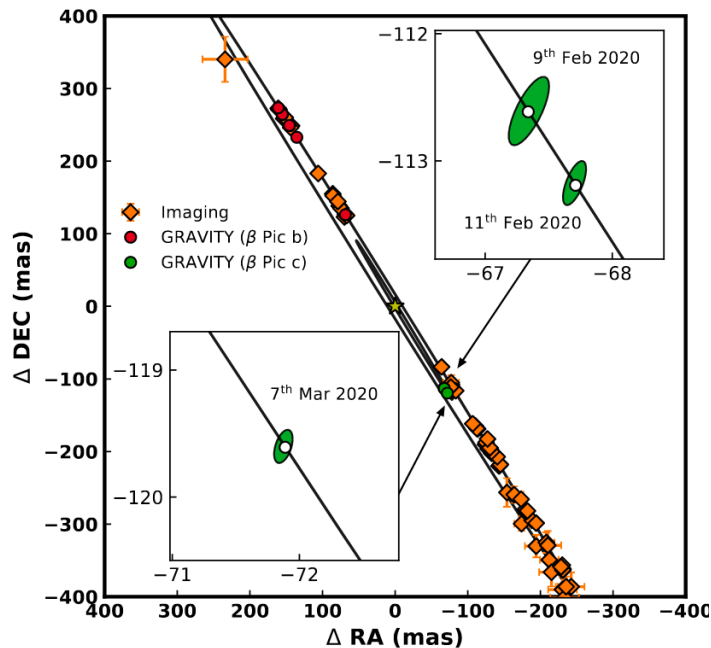


Gravity Collaboration
2021

Imaging the central source of the bipolar nebula IC2220

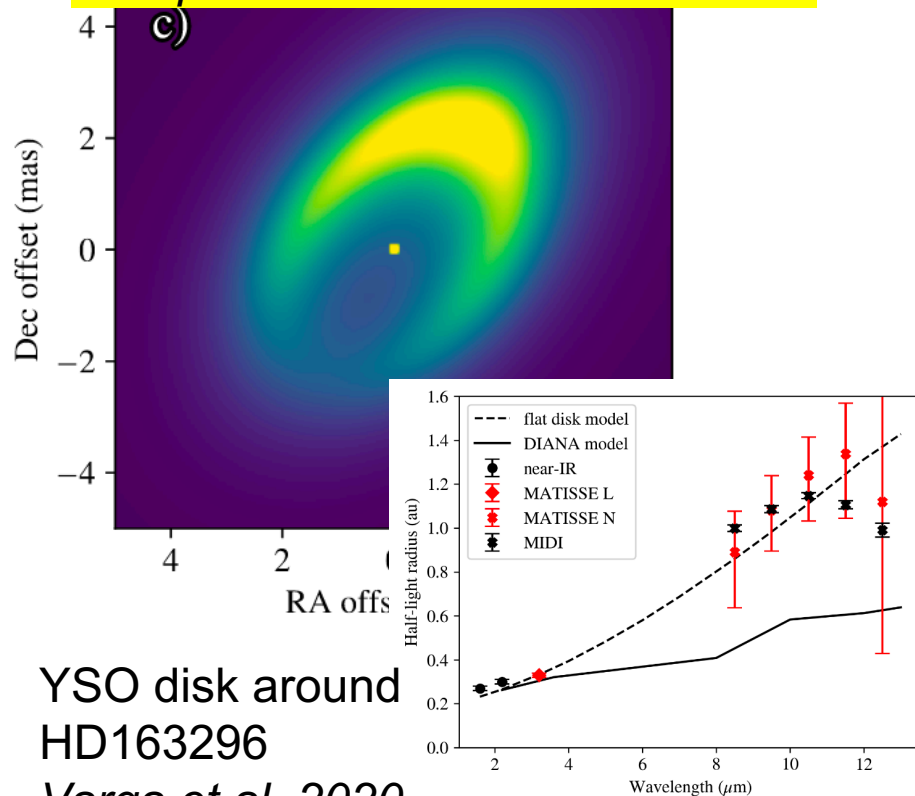


Exoplanets & YSOs



Direct detection of radial-velocity planet Beta pic C using GRAVITY
Nowak et al. 2020

First published result for MATISSE!



YSO disk around HD163296
Varga et al. 2020

Imaging Optimisation

■ Images requirements:

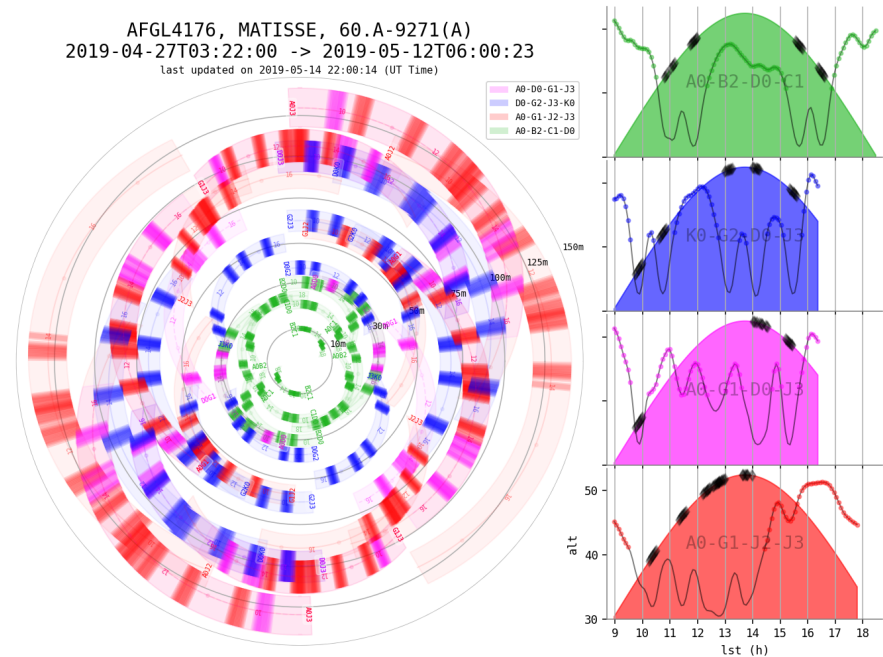
- Multiple AT configurations
- Multiple pointings per configuration, different LST
- Within a limited amount of time ($\sim < 1\text{mo}$)

■ Historically,

- observations over constrained (explicit config and LST range)
- Mix of Visitors / Service mode
- Fixed schedule for AT movements

■ New scheme:

- Reserved “only service” time for images, cycling through all AT configurations
- Flexible constraints + u,v monitoring



VLT expertise centres

- Goal: end-to-end support for users
 - ESO does not support everything: e.g. observation strategy, data interpretation, etc.
 - ALMA has great success thanks to end-to-end community support (ALMA ARC nodes)

- VLT ECs:
 - EU funded until 2025
 - Face-to-face support with traveling funds
 - Close coordination with ESO



The present network of VLT Expertise Centres includes three partners from the OPTICON Horizon 2020 networking activity:

- **Jean-Marie Mariotti Centre (JMMC) - Service aux Utilisateurs du VLT**, (SUV) France
- **Portuguese VLT Expertise Centre**, Portugal
- **University of Exeter**, United Kingdom

two interferometry JRA (Joint Research Activities; WP8) lead partners:

- **Observatoire de la Côte d'Azur**, France (cf. contact at **SUV** page);
- **KU Leuven**, Belgium

and two new nodes from the **OPTICON/RadioNet Pilot** (ORP) program:

- **Leiden Observatory**, The Netherlands
- **Konkoly Observatory**, Hungary

Visitors wishing to travel to the above centres to reduce their VLT data or prepare observations are encouraged to use the **Fizeau Programme**.

<https://european-interferometry.eu/centres-network/>

Technical activities: past year

■ MATISSE

- Last Commissioning activities affected by COVID

■ GRA4MAT

- Use GRAVITY as a fringe tracker for MATISSE (increase spectral coverage in L and M): MATISSE GRISM failure interrupted the commissioning, will be fixed traveling permitted

■ NAOMI (AO on ATs)

- Heaters installed on DM to avoid adverse cold temperature effects

■ Doubling of delay line pathlength

- Postponed due to COVID (fabrication)



- GRAVITY Consortium + new collaborators
- Improve sensitivity / sky coverage of GRAVITY+
 - Updated AO for high strehl
 - Better of-axis FT
 - LGS on all UT
 - Goal of K=22
- 3 main science cases:
 - Galactic Centre
 - AGNs
 - Exoplanets
- Currently in Phase A (design study)

Future: GRAVITY+

Top Level Requirements

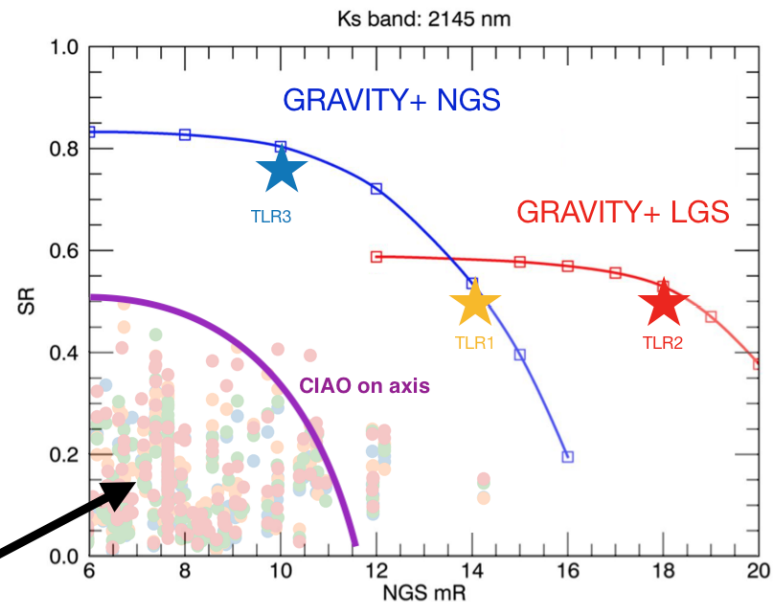
Table 1: Top-level performance specifications for GRAVITY+, the acronyms are given in appendix 1.

	Science case	AO mode	FT mode	AO star R mag	SR* in K-band	FT star K mag	FT rms OPD for K = 10	SC-FT distance	SC star K mag*
1	Galactic Center	LGS	off-axis	14	> 50%	10	< 100 nm	Up to 30"	22
2	Extragalactic, faint galactic	LGS / NGS	on/off-axis	18 (LGS) 10 (NGS)	> 50%	13 (goal 15)	< 100 nm	Limited by STS ** Text	22
3	Exoplanet & high contrast	NGS	on-axis	10	> 75%	10	< 100 nm	Limited by GRAVITY ***	22

* When operating close to the guide-star. For off-axis operation, the performance will be reduced by atmospheric anisoplanatism.

** The available patrol field of the current PRIMA STS is about 1' radius

*** Separation between exoplanet and host star limited by GRAVITY dual field FoV of 2" UT

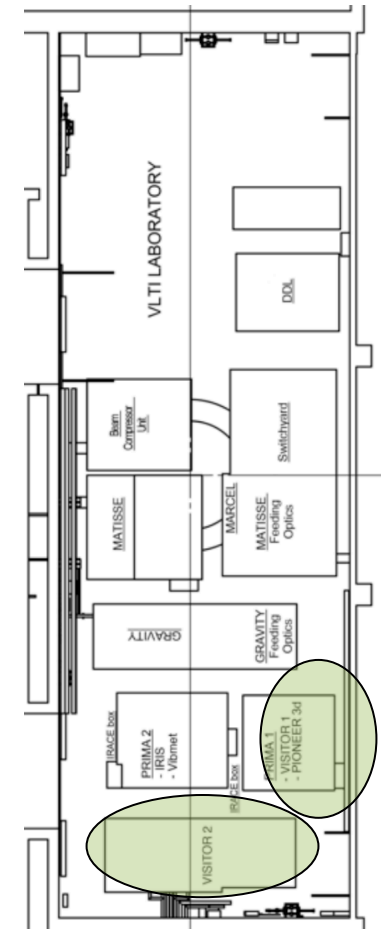


Current MACAO
measured performances

- 200+ registered, >50% non-interferometrists
- Fascinating extra galactic science case
 - 100x more AGNs accessible, up to $z \sim 7$
 - BH in globular clusters
 - ...
- Strong exoplanet science case
- Stellar community want higher spatial resolution and higher spectral resolution...

Visitor instruments

- VI and not facility instruments:
 - Build with less constraints
 - Operated by PI team
 - Focused science case, not offered to community
- Removal of AMBER, FINITO and PRIMA liberated space in the focal lab
- Many ideas presented in a 2019 workshop
 - <https://www.eso.org/sci/meetings/2019/VLT2030/program.html>
 - 2 funded by ERCs!



Funded VIs (ERC)

Hi-5 high contrast at 3-5um

<http://www.biosignatures.ulg.ac.be/ddefrere/hi5.php>

- PI: Denis Defrère (KU Leuven, BE)
- Science case: find exoplanets by nulling
- Timeframe 2021-2025
- They are hiring!
<https://fys.kuleuven.be/ster/vacancies>

BIFROST J spectrograph

<http://www.skraus.eu/presentations/kraus.GAIA-BIFROST.pdf>

- PI: Stefan Kraus (Exeter, UK)
- Science cases: YSOs, binary survey
- Timeframe 2021-2025
- Talk tomorrow at 13:10 PDT by S. Kraus

Conclusions

■ Exciting new sciences cases:

- AGN, Quasars
- Galactic Centre
- Microlensing
- Exoplanets
- ...

■ Generalisation of AO, Spectroscopy, Fringe Tracking

■ Ongoing call for proposal (P108):

- <http://www.eso.org/sci/observing/phase1/p108/CfP108.pdf>
- Deadline 25 March 2021, 12:00 noon CET
- NEW: GRA4MAT, imaging runs
- Ask Expertise Centre for help!

■ Improving Support

- Expertise Centres
- Imaging optimisation

■ Future is bright:

- Strong instruments: GRAVITY and MATISSE
- Community support: Expertise Centres + JMMC
- Projects: GRAVITY+, Hi-5, BIFROST