Welcome to the Univ. of Toledo CHARA Community Workshop

Douglas Gies

Director Center for High Angular Resolution Astronomy Georgia State University

www.chara.gsu.edu

















Your hosts

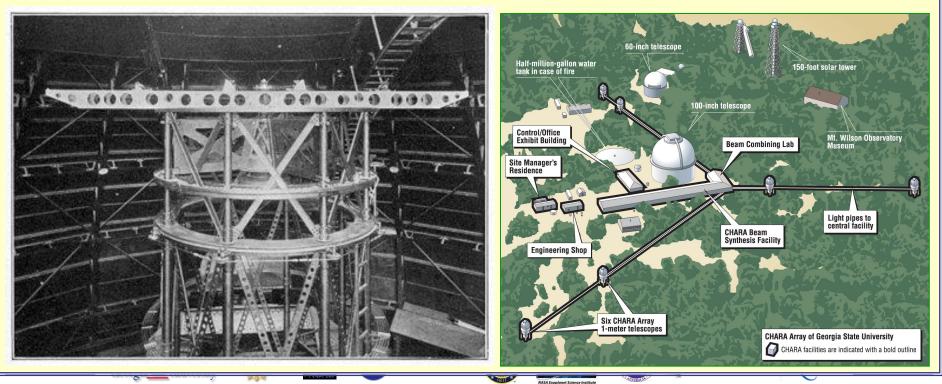
- Douglas Gies, CHARA Director, Dept. of Physics and Astronomy, Georgia State University, Atlanta, Georgia
- Gail Schaefer, Visitor Support Scientist, GSU CHARA Array, MWO, California
- Noel Richardson, Department of Physics and Astronomy, University of Toledo
- THANKS to University of Toledo and NSF!

bservatoire

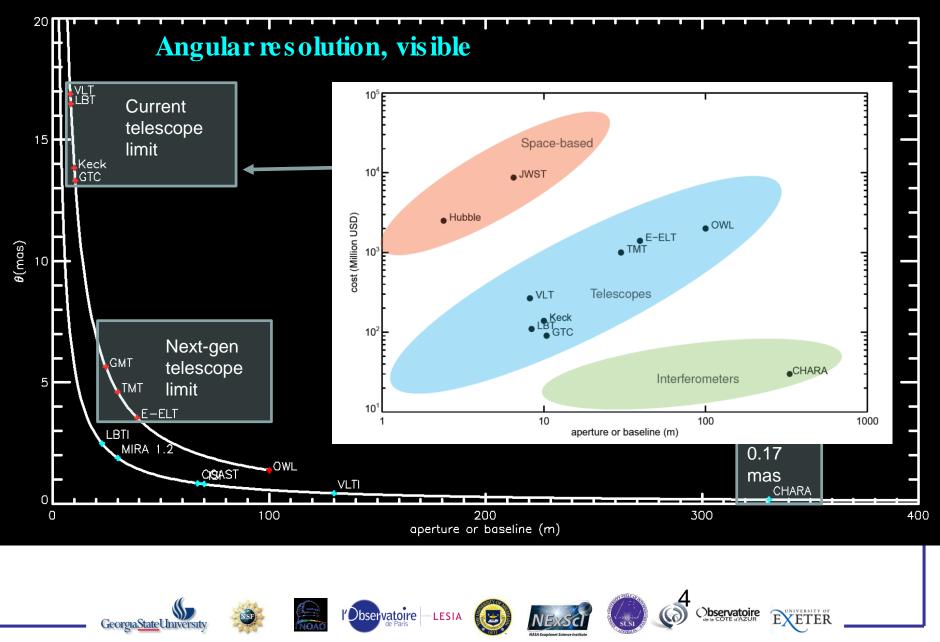


Stellar interferometry: measuring star size

- Michelson and Pease (1920): Betelgeuse
- Required technology (lasers, optics, computers)
- Culmination at Mount Wilson in CHARA Array: resolves most stars visible to human eye









Underground to listen to Divisiand Jack, reg-and rock bands, scarf up souvenir AOC T-rts, and see a fireworks display at 8.45 p.m. "It will be quite spectacular," said Laurie Ol-an AOC volunteer. "According to the guy

a is doing it, it's the biggest fire it's ever been done in Atlanta The 7:49 a m. announcement of Atlanta's vis

Observatoire

The CHARA Legacy of Prof. Hal McAlister



Founded in 1984 GSU Center for High Angular Resolution Astronomy with goal to build a world-class instrument.

Realized with the CHARA Array, the best of its kind in the world.

Ground breaking July 13, 1996.

First "fringe" November 1999.

Scientific observations since 2004.

Hal retired (Aug 2015), now emeritus professor at GSU.

















10000 km

VLBA spans the Earth and has the highest angular resolution of any observatory



Image © 2005 EarthSat

Google[.]

Pointer 33"08'16.94" N 108"28'03.21" W











Streaming |||||||| 100%

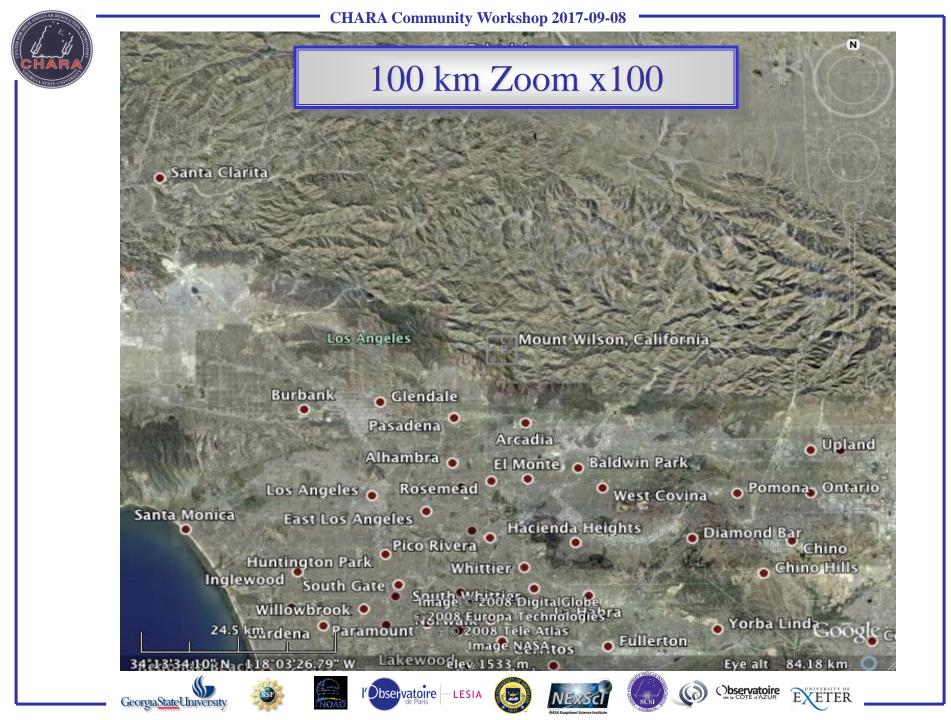




Eye alt 4014.55 mi

EXETER







1 km Zoom x10000

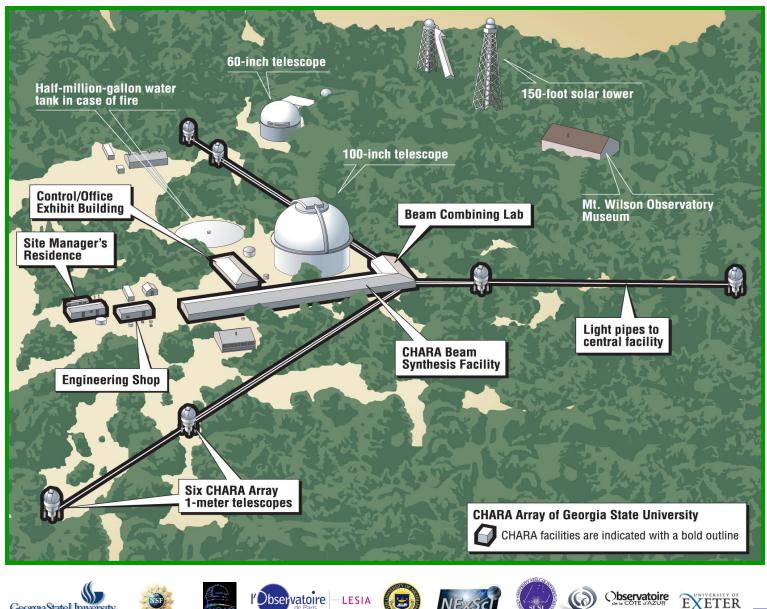
By using NIR and Visible light instead of radio waves, we can achieve the same angular resolution as VLBA but with a much smaller interferometer





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CHARA Community Workshop 2017-09-08 Layout of the CHARA Array





A short visit to Mount Wilson























Vacuum Light Tubes Feed Light from Each Telescope to the **Central Lab**











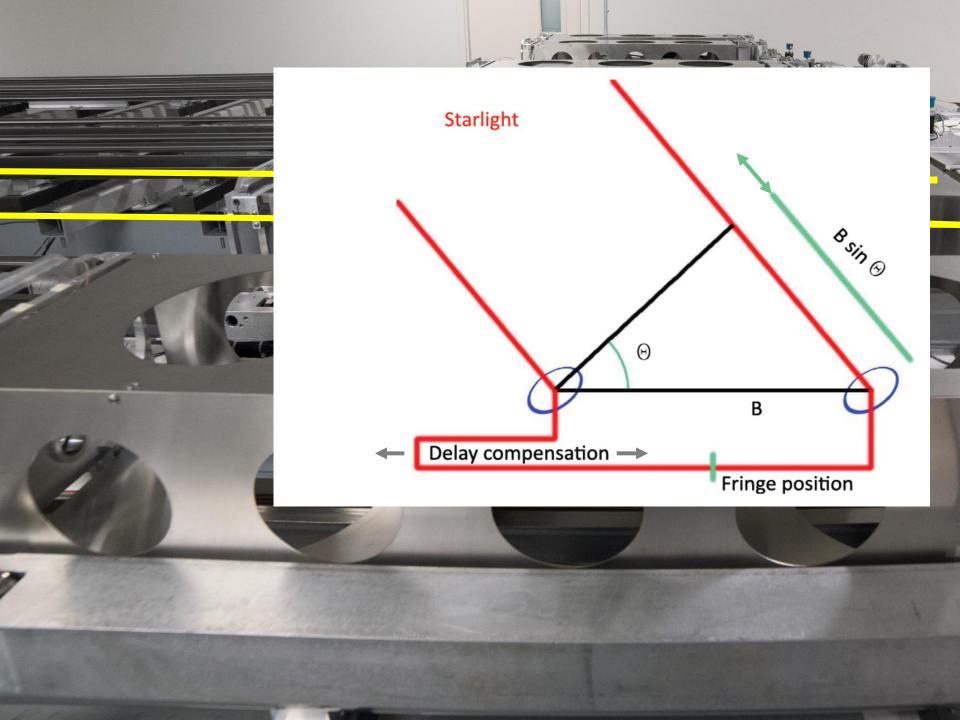


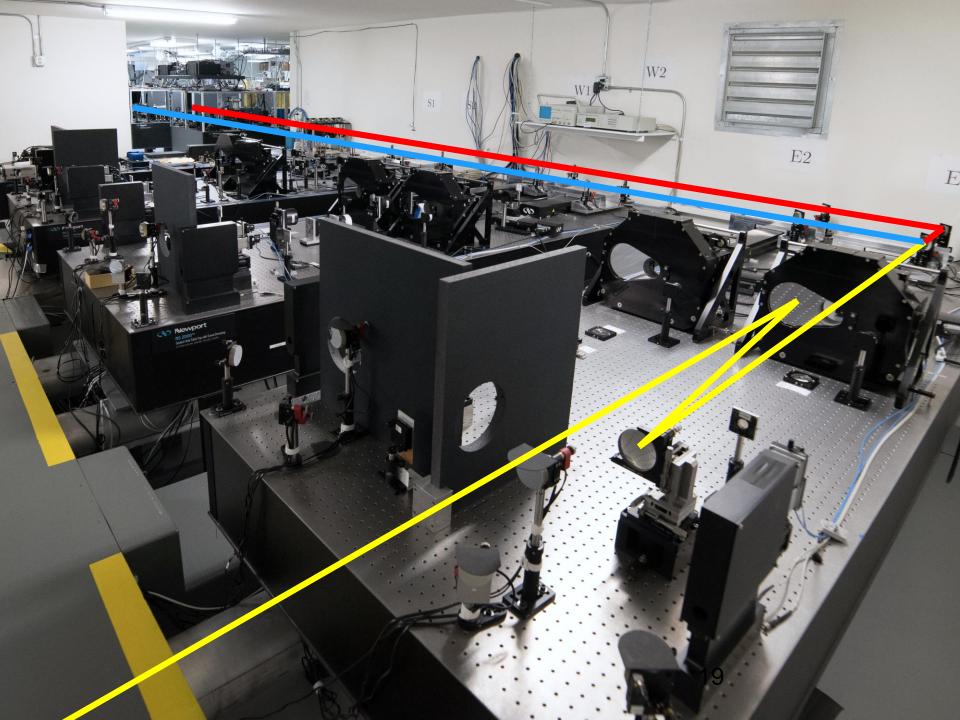






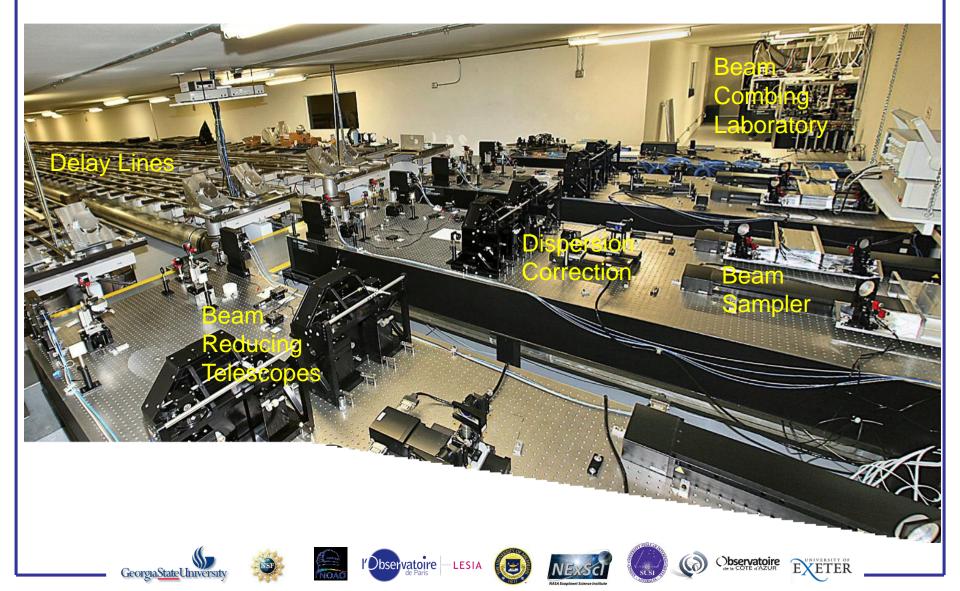






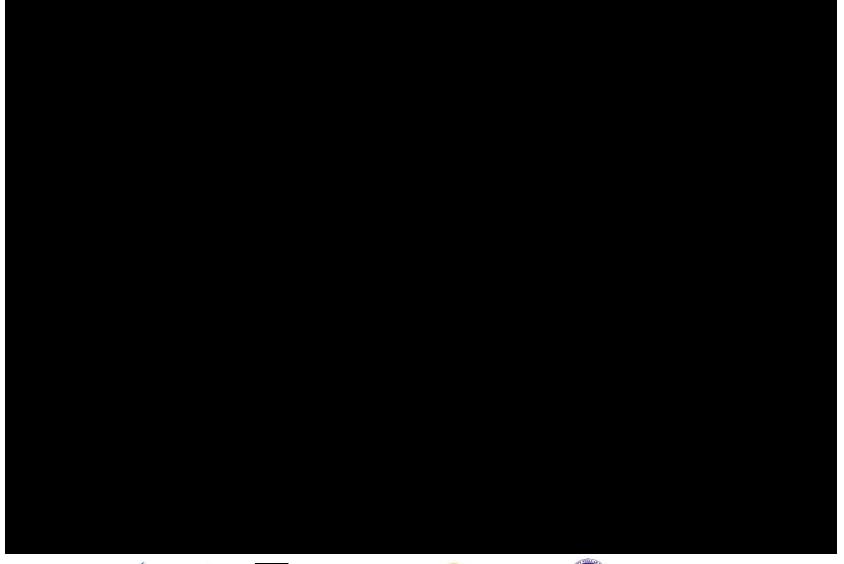


Optics Laboratory





The 30 second CHARA tour.















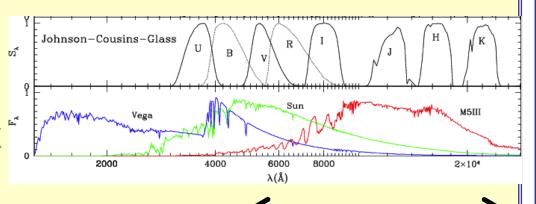


Beam Combiners: 0.5 - 2.2 microns

- Classic (GSU)
 2-tel open-air J,H,K
- CLIMB (GSU) Dual 3-tel open-air J,H,K
- JouFLUOR (Paris)
 2-tel fiber-based K band
- MIRC (Michigan)
 6-tel fiber-based imager H
- VEGA (Nice)
 4-tel open-air V,R,I R=6000, 30000
- PAVO (Sydney)
 2-tel aperture-plane V,R,I

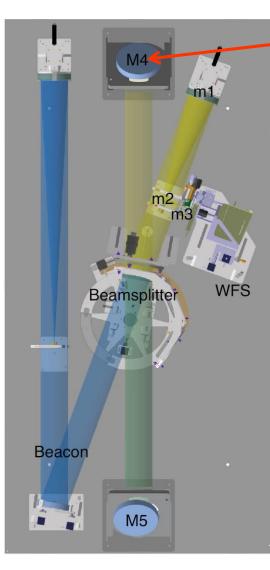
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• In progress: AO plus MIRCX, MYSTIC, SPICA



CHARA-AO: Telescopes and Lab





We will replace M4 with a deformable mirror at each telescope. This will enable us to correct for atmospheric seeing and increase scientific throughput.





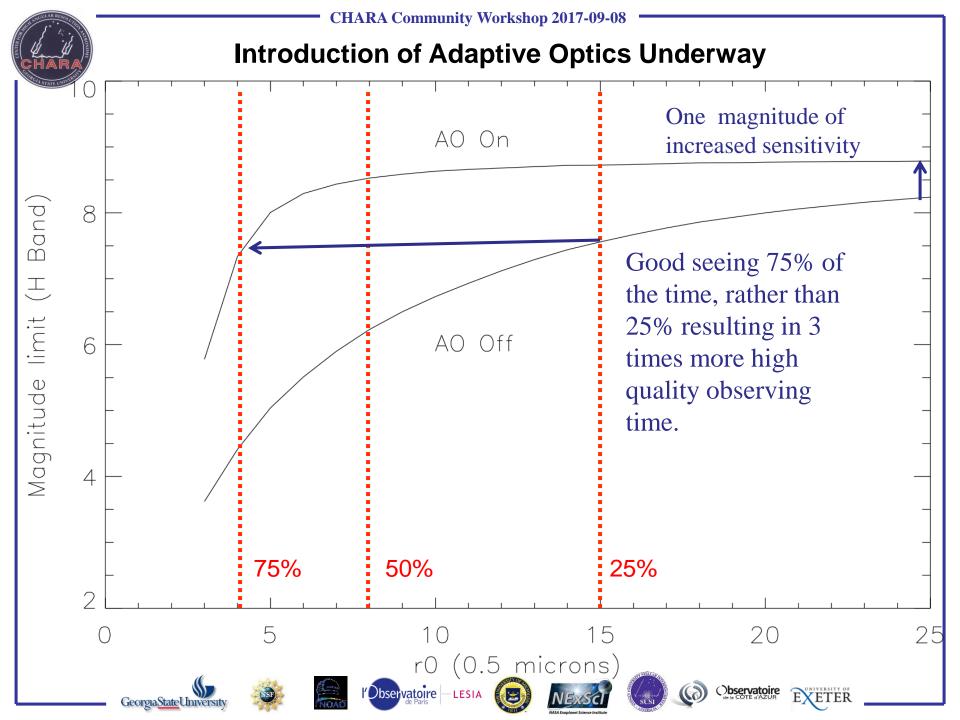






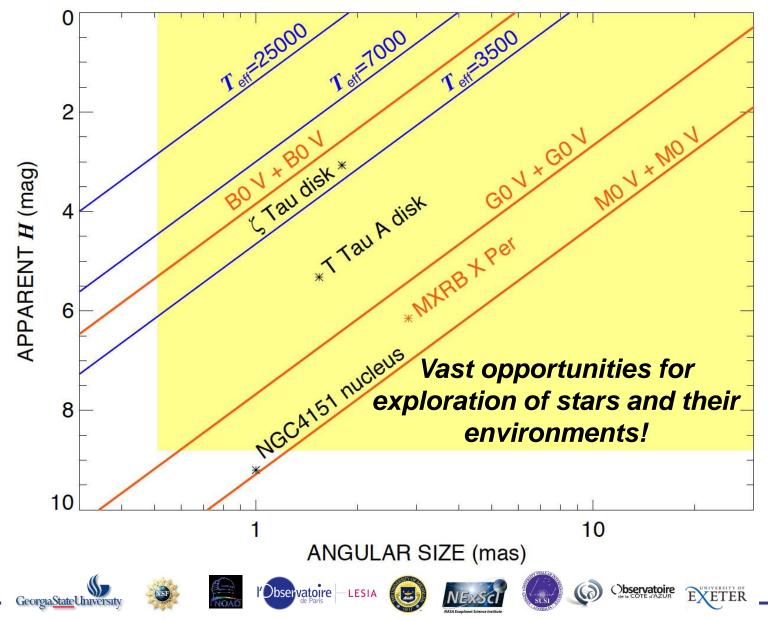






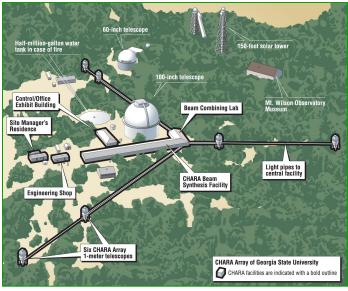


Working limits for Classic BC: stellar diameters, disk diameters, binary star separations (P=10d)

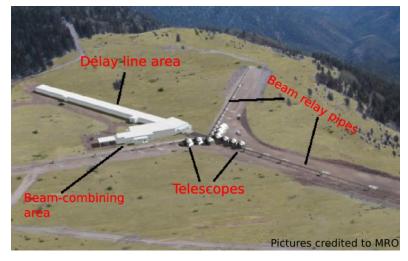




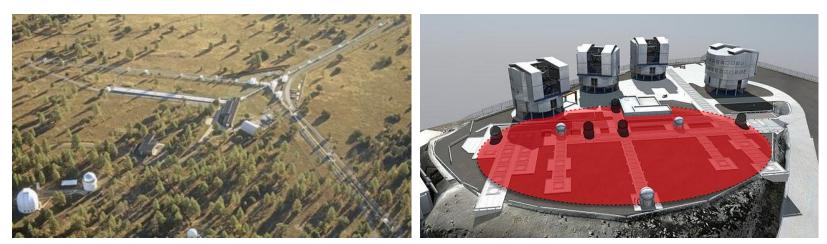
Other Long Baseline Interferometers



CHARA



MROI (under construction)

















VLTI



Goal: provide background for you to propose observations with the CHARA Array. *Morning Program*

Time	Торіс	Speaker
11:00	Overview of the CHARA Array	Douglas Gies
11:30	Interferometry & Science Review	Gail Schaefer
12:10	Applying for time at CHARA	Douglas Gies
12:40	Applying for time at VLTI	Jean-Baptiste Le Bouquin
1:00	Lunch	

















Afternoon Program

Time	Торіс	Speaker
2:00	Observing strategies and planning software	Gail Schaefer
2:20	Data format and modeling/imaging software	Gail Schaefer
2:40	Science topic: Imaging results with MIRC	John Monnier
3:00	Science topic: Massive stars	Noel Richardson
3:20	Science topic: Disk physics	Aaron Sigut
3:40	Open discussion	All
4:00	Optional visit to Ritter Observatory	Noel Richardson













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