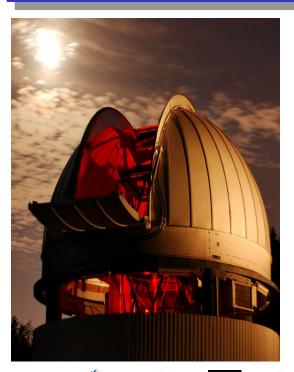


Observing Strategies and Planning Software



Gail Schaefer

The CHARA Array of Georgia State University

Mount Wilson, CA

With contributions from: Laurent Bourgès and Christopher Farrington























Outline

- Selecting a Beam Combiner
- Selecting Telescopes and Baselines
- Selecting Calibrator Stars
- Selecting Delay Settings (POP Configuration)
- Time Needed for Observations























Beam Combiners

Combiner	Num Tel.	Band	Typical Mag	Best Mag	Spec. Res.	Advantages	
CLASSIC	2T	H or K	7.0	8.5	Broad	Sensitivity	
CLIMB	3T	H or K	6.0	7.0 Broad		Sensitivity	
JouFlu	2T	K	4.5	5.0	Broad	Precision	
MIRC	6T	Н	4.5	6.0	40	Imaging	
PAVO	2T	630-900 nm	7.0	8.0	30	Sensitivity	
VEGA – HiRes	2-3T	2 bands (7nm) in 480-850 nm	4.0	5.0	30000	Spectral Res	
VEGA – MedR	2-3T	2 bands (35 nm) in 480-850 nm	6.5	7.5	6000	Spectral Res	

Limit for acquisition and tiptilt tracking: V = 10-12 mag















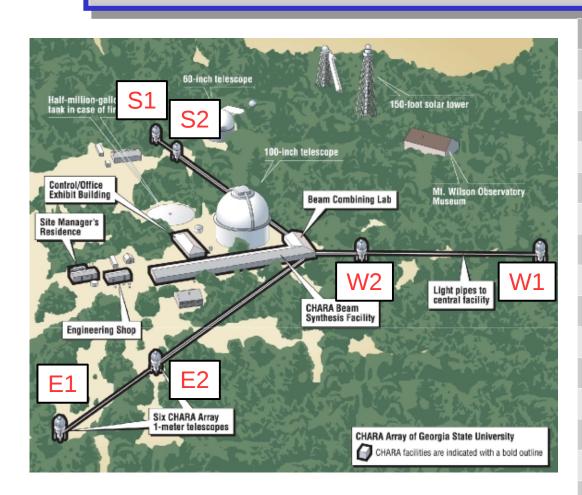








Telescopes and Baselines



Baseline	Length (m)
Daseillie	Lengur (m)
E1-S1	331
W1-E1	314
E1-S2	302
E2-S1	279
W1-S1	279
W1-E2	251
W1-S2	249
E2-S2	248
W2-S1	211
W2-E1	222
W2-S2	177
W2-E2	156
W2-W1	108
E2-E1	66
S2-S1	34























Selecting Beam Combiner and Baselines

- Angular Resolution: 0.5 λ /B
 - 0.66 mas in K-band (2.13 μm)
 - 0.52 mas in H-band (1.67 μm)
 - 0.20 mas in visible at 650 nm
- Simple diameter:
 - Single baseline (two telescopes)
- Imaging complex sources: Rapid rotators, binaries, stellar surfaces
 - Multiple baseline projections
 - Sample beyond the first null (at 1.22 λ /B)















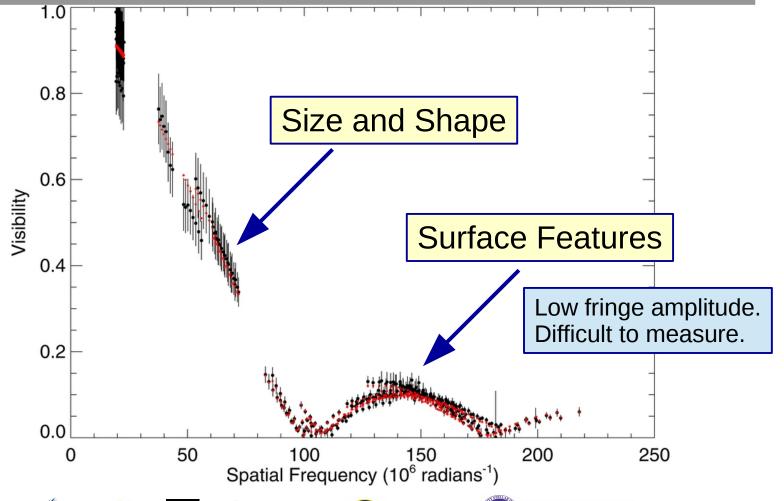








Selecting Beam Combiner and Baselines

























Wavelength Coverage

- Broad-band
 - Higher sensitivity for faint objects
 - Bandwidth smearing
- Spectrally dispersed visibilities
 - Increase u,v sampling by measuring fringes in different wavelength channels
 - Emission/absorption line studies
- Coherence length (width of fringe packet)
 - Sets the interferometric field of view

















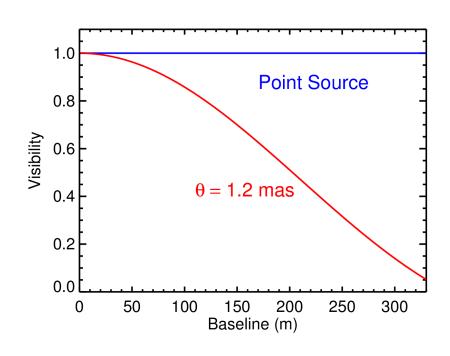








Calibrator Stars



- Unresolved point source:
 - Visibility = 1.0
- However, instrumental and atmospheric effects will cause a loss in coherence, causing a drop in the measured visibility.
- Observe unresolved calibrator stars to define the true visibility of the target.























Selecting Calibrators

- Unresolved stars or stars with a known angular diameter.
- Within 5-10 degrees on the sky from the science target.
- Within 1-2 mag in brightness from science target and similar in color.
- Avoid binary stars, rapid rotators, emission line stars.
- Minimum of two calibrators per object, three is better.
 - Discovery of unknown binaries























Selecting Calibrators

- SearchCal developed by JMMC
 - http://www.jmmc.fr/searchcal_page.htm
- getCal developed by NexSci
 - http://nexsciweb.ipac.caltech.edu/gcWeb/gcWeb.jsp























SearchCal





SearchCal [c1]											000		
File Edit Query Calibrators Interop Help													
Query Parameters													
1) Inst	trumental	Configu	ration 2)	Science Object —				3) 5	earchCal	Parameter	s		
Magnitude Band : H ▼ Name : Q ▼ HD 69897 ⊗ Min. Magnitude (H) : 3.0													
Wavelength (H) [µm]: 1.65 RA 2000 [hh:mm:ss]: 08 20 03.86158 Max. Magnitude (H): 5.0													
Max. Baseline [m]: 300.0 DEC 2000 [+/-dd:mm:ss]: +27 13 03.7416 Scenario: ® Bright ○ Faint												Coint	
Magnitude (H): 3 942												Failit	
Magnitude (H): 3.942 RA Range [mn]: 24													
DEC Range [deg] : [20.0													
Progress: Get Calibrator													rators
													Tators
Index	dist A	HD	RAJ2000	DEJ2000	vis2	LDD	UD V	UD H	UD K	SpType	V	Н	K [1
1	5.21E-6	69897	08 20 03.8602		0.374	0.701	0.662	0.686	0.689	F6V	5.13	3.942	3.868
2	2.975	67542	08 09 35.1816		0.622	0.503	0.468	0.488	0.49	GOII	6.47	4.699	4.621
3	3.383	67544	08 09 24.8645	+24 49 34.0716	0.619	0.509	0.468	0.49	0.493	G8III	7.29	4.966	4.908
4	3.595	71730	08 29 40.0634	+24 20 40.9452	0.581	0.542	0.496	0.52	0.523	KOIII	7.05	4.872	4.81
5	3.977	73080	08 37 22.1112	+28 17 39.8328	0.554	0.555	0.52	0.541	0.543	G5	6.63	4.702	4.591
6	4.945		08 10 54.7320	+22 43 43.1904	0.514	0.588	0.548	0.573	0.575	K0	8.026	4.909	4.698
7	6.121	<u>65471</u>	07 59 42.6055	+23 10 58.4652	0.506	0.594	0.554	0.579	0.581	K0	6.92	4.73	4.562
8	6.902	<u>75216</u>	08 49 45.3118		0.534	0.581	0.529	0.557	0.561	K2III	7.38	4.868	4.712
9	7.14	63138	07 48 28.8108		0.519	0.592	0.542	0.568	0.572	KOIII	6.86	4.694	4.605
10	7.303	<u>75646</u>	08 52 00.4543		0.568	0.553	0.504	0.53	0.534	K2III	7.54	4.983	4.834
11	7.524		08 52 09.6634	+29 51 13.3848	0.515	0.588	0.548	0.572	0.575	K0	7.08	4.742	4.631
12	7.677 7.731	75783	08 32 54.2333 08 53 00.0972		0.544 0.564	0.565 0.549	0.524	0.549 0.533	0.552 0.536	K2 K2	7.52 7.35	4.934 4.982	4.798 4.813
13 14	7.731	74198	08 43 17.1466		0.774	0.349	0.351	0.362	0.362	AliV	4.66	4.788	4.638
15	7.842	64092	07 53 01.0094		0.557	0.553	0.515	0.538	0.541	K0	7.05	4.85	4.755
16	7.845	0.1002	08 03 34.1340		0.599	0.519	0.486	0.506	0.508	G5	7.03	4.836	4.742
17	8.769	67482	08 09 39.7601	+35 42 08.5032	0.58	0.535	0.498	0.52	0.523	KO	7.3	4.952	4.839
18	8.815	64602	07 56 01.9399	+34 22 10.4160	0.572	0.541	0.505	0.527	0.53	K0	7.57	4.972	4.827
19	8.917		07 48 06.8957	+32 51 25.0308	0.557	0.552	0.518	0.539	0.541	G5	7.204	4.952	4.548
20	10.159	60204	07 34 31.5922		0.519	0.583	0.547	0.569	0.571	G5	6.66	4.605	4.498
21	10.402	<u>77694</u>	09 04 51.4817	+24 36 18.5040	0.561	0.559	0.509	0.536	0.539	K2III	7.8	4.982	4.838
	1												• 3
Filters			*************										
				D. C /							, .		
Reje	ct stars fa	arther th	an: Maximum	RA Separation (m	n) : 10.0		ма	XIMUM DE	C Separat	ion (degre	e) : 10.	0	
Reje	ct stars w	ith magr	nitude: below	0.0			a	nd above	: 10.0				
Reje	ct Spectra	al Types	(and unknowns	:):						V O V E	B ∠ A [∠ F ∠ G	■K ■M
Reje	ct Lumino	sity Clas	ses (and unkno	owns):						I	□ II □	III 🗹 IV	V V VI
∠ Reje	ct Visiblity	y below :	vis2: 0.5										
Reje	ct Visibilit	y Accura	cy above (or ur	ıknown) : vis2Err	/vis2 (%) :	2.0							
	ct Variabi	-				-							
✓ Reject Multiplicity													
✓ Reject Invalid Object Types													
			aximum chi squa	are: 2.0			Maximur	n relative	error (%)	: 100			
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sear	ching calil	brators	. done.									Provided	by JMMC



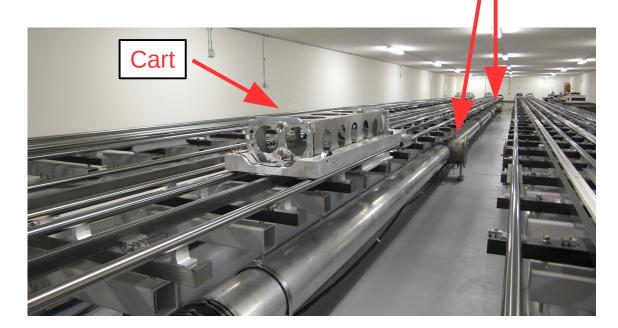


Delay settings (sky coverage)

Delay settings to equalize optical path length

Fixed increments: PoPs

Variable delay: carts



















PoPs







Planning Software

- ASPRO2 developed by JMMC
 - http://www.jmmc.fr/aspro_page.htm
- CHARA_PLAN2 developed by CHARA
 - http://www.astro.gsu.edu/~theo/chara_reduction/index.html

















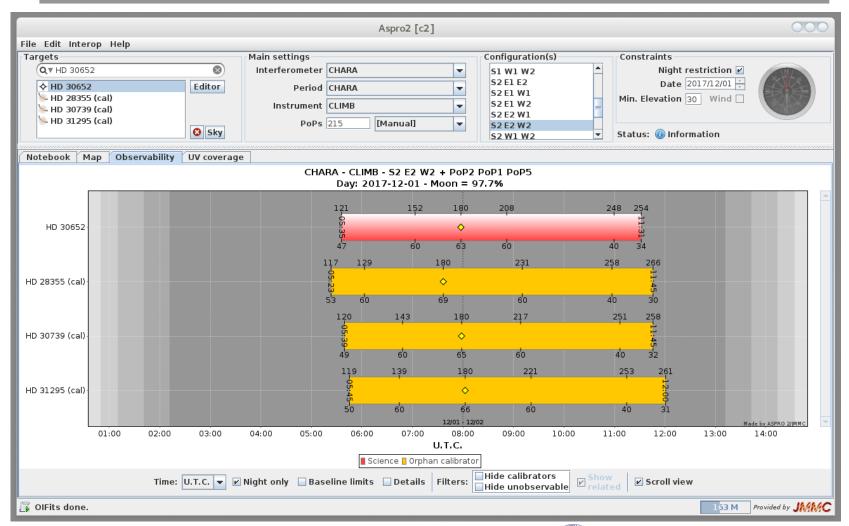






ASPRO 2 JA





















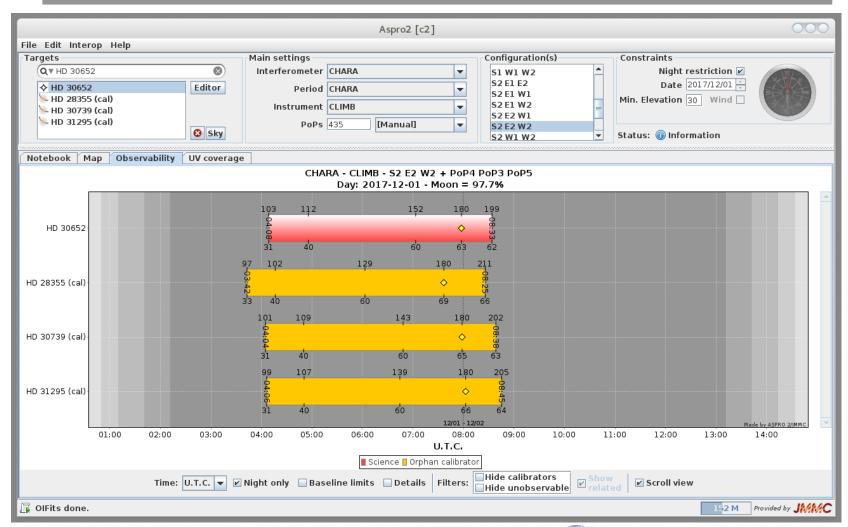






ASPRO 2 JA





























			Aspro2 [c1]				000
le Edit Interop Help argets		Main settings		Configuration(s)		Constraints	
Q ▼ HD 56537	S	Interferometer	CHARA	▼ 51 W1 W2		Night restriction 🗹	N
♦ HD 30652	Editor	Period	CHARA	S2 E1 E2 S2 E1 W1		Date 2017/12/01	
► HD 28355 (cal)		Instrument	CLIMB	S2 E1 W2		Min. Elevation 30 Wind 🗌	ZXX
№ HD 30739 (cal) № HD 31295 (cal)		PoPs		S2 E2 W1 S2 E2 W2			S
	3 Sky	FOFS	213	S2 W1 W2	-	Status: 📵 Information	
otebook Map Observability UV	coverage OIFits	viewer					
Instrument mode	coverage Offics	viewei	CHARA - CLIMB -	S2 E2 W2 + PoP2 PoP1 Po	P5		
н 🔻				2-01 - Source: HD 30652			
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8	0.90		250		450		
	0.85		250		400		
	0.75		200		350		
40	0.70		150		250		
Total Integration time (s)	0.65		100	-\ 	200		
300	0.60			WHITH I	150		
-6.00	0.55		50	- Huge	100	<u> </u>	— S2-
HA max 3.58	0.50		0 (MX)	HILLY MANUEL	0	(m) -	— E2-
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8	0.35		-100	Million	-20		
	0.30		-150		-25)	
3	0.25		-200		-30		
[3]	0.20				-35		
E compare on its data	0.15		-250		45		
	0.10		-300 -250 -200 -150		by ASPRO 2/JMMC -50)	
	0.00		-300 -250 -200 -150	U (Mλ)	200 250 300		
OIFits done.	0.00						Provided by JMA























CHARA PLAN 2

							CHAR	A_PLAN						000	
S1	POP1		BEAM1	~ □ S1	UT Yea	ar: 201	7	רט	UT Month : 12			UT Day :	01		
31	POPI		BEAMIT	21	SLEW TO	CAL 1	WHEN	FIND	IRC		ŀ	HD_28355			
S2	POP2		BEAM1	∨ ✓ S2	NUM	43449	IRC		HR	1414	HD	28355	SAO	93960	
32	SZ POPZ BE	DEAIVIT	32	CI	EAR	RA: 04 28	50.1641 DE	C: +13 02 51	.369 Vmag: 5	5.01 Kmag: 4.	.53 Type: A7V				
E1	POP1	~	BEAM1	→ E1	SLEW TO	OBJECT	WHEN	WHEN FIND IRC			HD_30652				
	rori		BLANT		NUM	48290	IRC	+10071	HR	1543	HD	30652	SAO	112106	
E2	E2 POP1 Y	BEAM1	∨ ⊘ E2	CI	EAR	RA: 04 49	50.4106 DE	C: +06 57 40	.592 Vmag: 3	1.19 Kmag: 1.	.60 Type: F6V				
	1011		DLAITI		SLEW TO	CAL 2	WHEN	FIND	IRC		H	HD_30739			
W ₁	POP1	~	BEAM1	~ □ W1	NUM	48490	IRC		HR	1544	HD	30739	SAO	112124	
	1011		DE/II 11		CI	CLEAR RA: 04 50 36.7217 DEC: +08 54 00.633 Vmag:					.35 Kmag: 4.	17 Type: A1Vn			
W2	END	ND Y	BEAM1	√ ⊘ W2	SLEW TO	CHECK	WHEN	FIND	IRC		·	HD_31295			
	2.15		22,11.12		NUM	49700	IRC		HR	1570	HD	31295	SAO	94201	
Ι.,	Reference	Cart	W2	~	CI	EAR	RA: 04 54	53.7279 DE	279 DEC: +10 09 02.999 Vmag: 4.66 Kmag: 4.42 Type: AOV						
					JOB QUEUE: 0 START JOB QUEUE			JE STOP J	STOP JOB QUEUE CLEAR JOB QUEUE PI:				Prog:		
	1: HD_30 HARA 482			HR_1543 HD_	30652 SAO_1	112106.									
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CHARA PLAN 2

							CHAR	A_PLAN						000	
S1	POP1	~	BEAM1	∨ □ S1	UT Yea	ar: 201	7	רט	UT Month : 12			UT Day :	01		
51	POPI		BEAMIT	31	SLEW TO	CAL 1	WHEN	FIND	IRC		ŀ	HD_28355			
S2	POP2	~	BEAM1	√ ♂ S2	NUM	43449	IRC		HR	1414	HD	28355	SAO	93960	
52	POPZ		BEAMIT	52	CI	EAR	RA: 04 28	50.1641 DE	C: +13 02 51	.369 Vmag: 5	5.01 Kmag: 4.	.53 Type: A7V			
E1	POP1	~	BEAM1	∨ □ E1	SLEW TO	EW TO OBJECT WHEN FIND IRC			ŀ	HD_30652					
E1	FOFI		DEAMI		NUM	48290	IRC	+10071	HR	1543	HD	30652	SAO	112106	
E2	POP1	~	BEAM1	∨ ⊘ E2	CI	EAR	RA: 04 49	50.4106 DE	C: +06 57 40	.592 Vmag: 3	3.19 Kmag: 1.	.60 Type: F6V			
	FOFI		BLANT		SLEW TO	CAL 2	WHEN	FIND	IRC		ŀ	HD_30739			
W ₁	1 POP1 V	BEAM1	√ □ W1	NUM	48490	IRC		HR	1544	HD	30739	SAO	112124		
***	1011		BEAPIT		CI	CLEAR RA: 04 50 36.7217 DEC: +08 54 00.633 Vmag: 4.3					.35 Kmag: 4.	17 Type: A1Vn			
W2	W2 END Y	~	BEAM1	√ ⊘ W2	SLEW TO	CHECK	WHEN	FIND	IRC		ŀ	HD_31295			
""			DE, II T		NUM	49700	IRC		HR	1570	HD	31295	SAO	94201	
١.,	Reference	Cart	W2	~	CLEAR RA: 04 54 53.7279 DEC: +10 09 02.999 Vmag: 4.66 Kmag: 4.42 Type: A0V										
					JOB QUEUE: 0 START JOB QUEUE			JE STOP J	STOP JOB QUEUE CLEAR JOB QUEUE PI:				Prog:		
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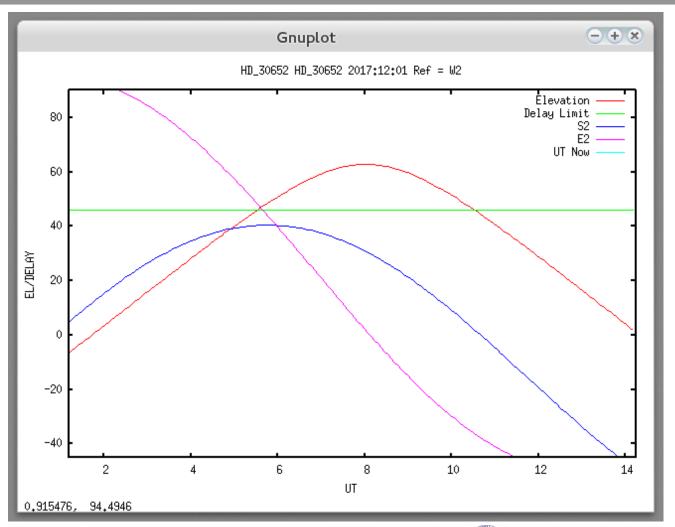








CHARA PLAN 2

























How much time is needed?

- Calibration Strategies:
 - Cal1 Obj Cal2 Obj Cal1 ...
 - Cal1 Obj Cal2 Cal1 Obj Cal2 ...
- Time to collect data on single object (star acq. + data)
 - Seeing and brightness dependent
 - Fast instruments (CLASSIC, CLIMB, PAVO, JouFlu):
 - 5 10 minutes
 - VEGA: 5 20 minutes
 - MIRC: 45 60 minutes
- Cal-Sci-Cal will take between 30 120 min





















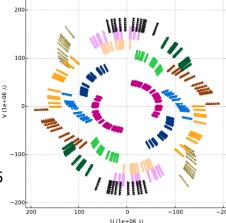


How much data is needed?

- Diameters Several brackets of data per baseline on two separate nights.
- Binaries Minimum of three brackets or observations on at least three baselines.

Imaging – Many brackets on multiple baselines during

the night to fill in the sky coverage.



Kloppenborg et al. 2015

















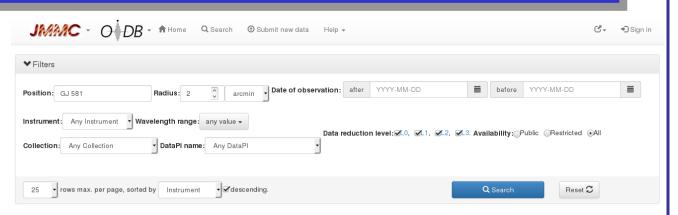






Do observations already exist?

- OI Database
- Query and download data (OIFITS)
- CHARA observation logs for Classic, CLIMB, VEGA only



Results

Meta-data will try to follow VO4OI proposal and Ivoa:ObsCore document (get metadata description in the associated doc 33 observations from 1 oifits files (0 private)

SELECT ALL * FROM oidb AS t WHERE (CONTAINS(POINT('ICRS', t.s_ra, t.s_dec), CIRCLE('ICRS', 229.8617625, -7.7222806, 0.03333333333333333))=1) ORDER BY instrument

Page 1 / 2 Next Last

				(Luit query)				
O +	target_name	access_url	t_min	instrument_name +	wlen_min	wlen_max	nb_channels	datapi
₩	HIP_74995	-	2008-05-16T09:38:52	CLASSIC	1.96000000	2.31000000	-	Baines 🔀
Q +	HIP_74995	-	2010-03-30T08:09:35	CLASSIC	1.53000000	1.82000000	-	Boyajian 🔀
Ø +	HIP_74995	-	2010-03-30T08:31:12	CLASSIC	1.53000000	1.82000000		Boyajian 🔀
Ø +	HIP_74995	-	2010-03-30T09:44:38	CLASSIC	1.53000000	1.82000000		Boyajian 🔀
₩	HIP_74995	-	2010-03-30T10:13:26	CLASSIC	1.53000000	1.82000000	-	Boyajian 🔀

http://oidb.jmmc.fr/index.html























On the night of observation

- Observations will be carried out by CHARA staff
- Visitors are encouraged to travel to the Array to participate in the observations
 - Real-time input from PI on decisions that could impact the science objectives and priorities
- Guide to planning observations available on the CHARA website:
 - http://www.chara.gsu.edu/observers/planning-an-observation



















