# Planning Observations and Software Tools



### **Gail Schaefer**

The CHARA Array of Georgia State University

Mount Wilson, CA

With contributions from: Fabien Baron, Laurent Bourgès, Christopher Farrington, and Jeremy Jones



- Visitor Support Scientist to help with planning observations
- Observations carried out by CHARA staff
- Visitors are encouraged to travel to the Array to participate in the observations
  - Real-time input from PI on decisions that could impact the science objectives and priorities





- CHARA staff supports data reduction to OIFITS format
  - Data reduction software available for those interested in reducing and calibrating data
- OIFITS Format:
  - Data exchange standard for Optical Interferometry
  - Description of format (target info, UV coord, VIS2, T3)
  - Resources available for reading OIFITS files (C, IDL, OIFITS Explorer)



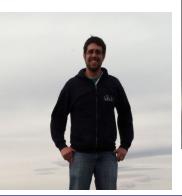
# Data Access and Searchable Archive

- Under Development
- Located at GSU Data Center
- 3 Virtual Machines on CHARA server:
  - Database/Archive Machine
  - Data Reduction Machine
  - Remote Observing Machine

Data Scientist

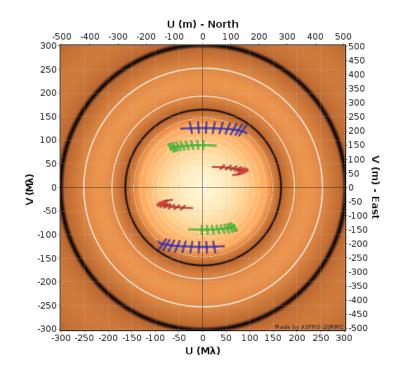
Jeremy Jones

- Active Mode
- Passive Mode



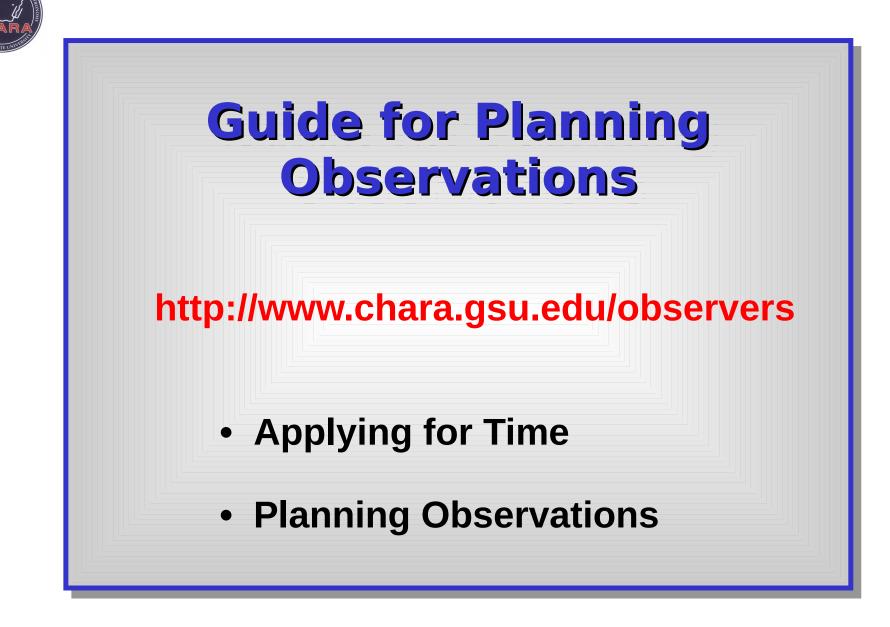


# **Data Analysis**



- Interferometers measure the Fourier Transform of the brightness distribution
- Sparse sampling
- Geometric model fitting
- Physical models
- Image reconstruction

http://www.chara.gsu.edu/observers/data-analysis-software





Combiner	Num Tel.	Band	Typical Mag	Best Mag	Spec. Res.	Science
CLASSIC	2T	H or K	7.0	8.5	Broad	Diameters
CLIMB	3T	H or K	6.0	7.0	Broad	Binaries, disks
JouFlu	2T	К	4.5	5.0	Broad	Diam, precision
MIRC	6T	Н	5.0	6.0	40	Stellar imaging, binaries, disks
PAVO	2T	630-900 nm	7.0	8.0	30	Diameters
VEGA – HiRes	2-4T	2 bands (7nm) in 480-850 nm	4.0	5.0	30000	Spectral studies
VEGA – MedR	2-4T	2 bands (35 nm) in 480-850 nm	6.5	7.5	6000	Spectral studies, diam.

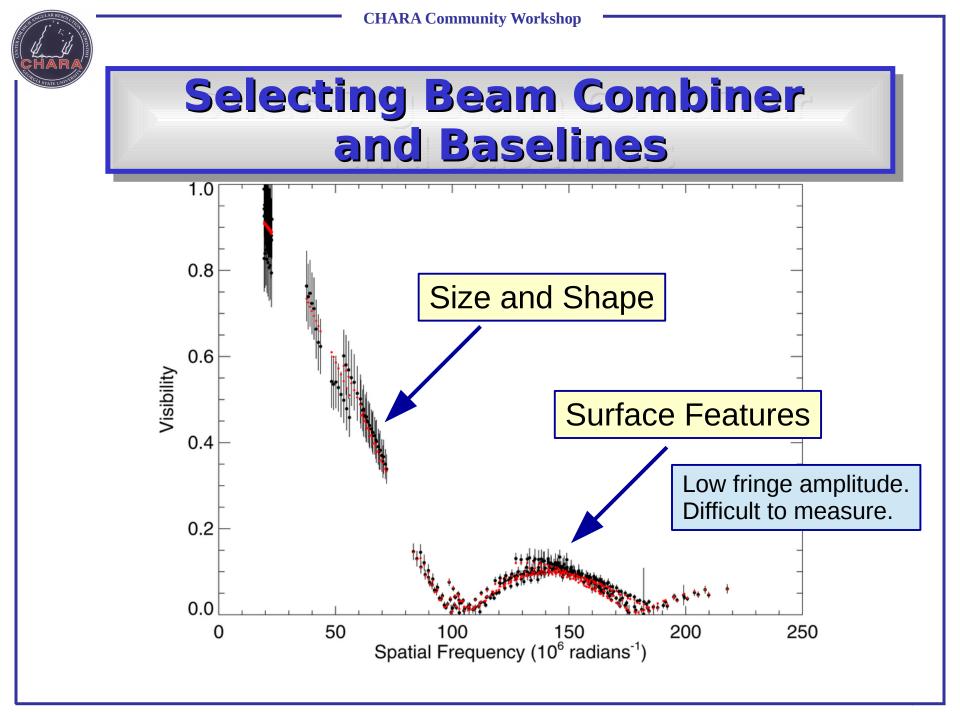
Limit for acquisition and tiptilt tracking: V = 10-12 mag

## Selecting Telescopes and Beam Combiner

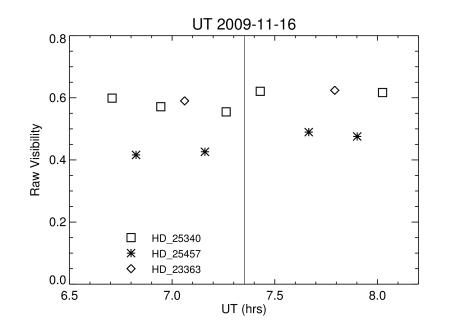
- Angular Resolution on Longest Baseline: 0.5  $\lambda$  /B
  - 0.66 mas in K-band (2.13 µm)
  - 0.52 mas in H-band (1.67 µm)
  - 0.20 mas in visible at 650 nm

## Selecting Telescopes and Beam Combiner

- Angular Resolution on Longest Baseline: 0.5  $\lambda$  /B
  - 0.66 mas in K-band (2.13 µm)
  - 0.52 mas in H-band (1.67 µm)
  - 0.20 mas in visible at 650 nm
- Resolving stellar diameters
  - One or two baselines long enough to resolve star
- Binaries
  - Ideally three or more telescopes (2D separation)
- Imaging stellar surface features
  - Multiple baselines (all 6 telescopes)
  - Sample beyond the first null (at 1.22  $\lambda$ /B)







- Instrumental and atmospheric effects will cause a loss in coherence, causing a drop in the measured visibility.
- Observe unresolved calibrator stars to define the true visibility of the target.
  - Within 5-10 degrees on sky
  - Within 1-2 mag in brightness

# How much time is needed?

- Calibration Strategy:
  - Cal1 Obj Cal2 Cal1 Obj Cal2 ...
- Time to collect Cal-Sci-Cal set:
  - Seeing and brightness dependent
  - CLASSIC, CLIMB, PAVO, JouFlu: 15 45 minutes
  - VEGA: 30 60 minutes
  - MIRC: 90 minutes for Cal-Sci set
- Collect many repeated calibration sets
  - Improve detection, test systematics
  - Increase u,v coverage on the sky



### Using planning software tools:

### www.chara.gsu.edu/workshops/instructions



Center for High Angular Resolution Astronomy

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Workshop Tutorial - Planning CHARA Observations

1. Pick a target - estimate the angular diameter (or binary separation, disk size, etc).

JMMC Stellar Diameters Catalogue available through VizieR.

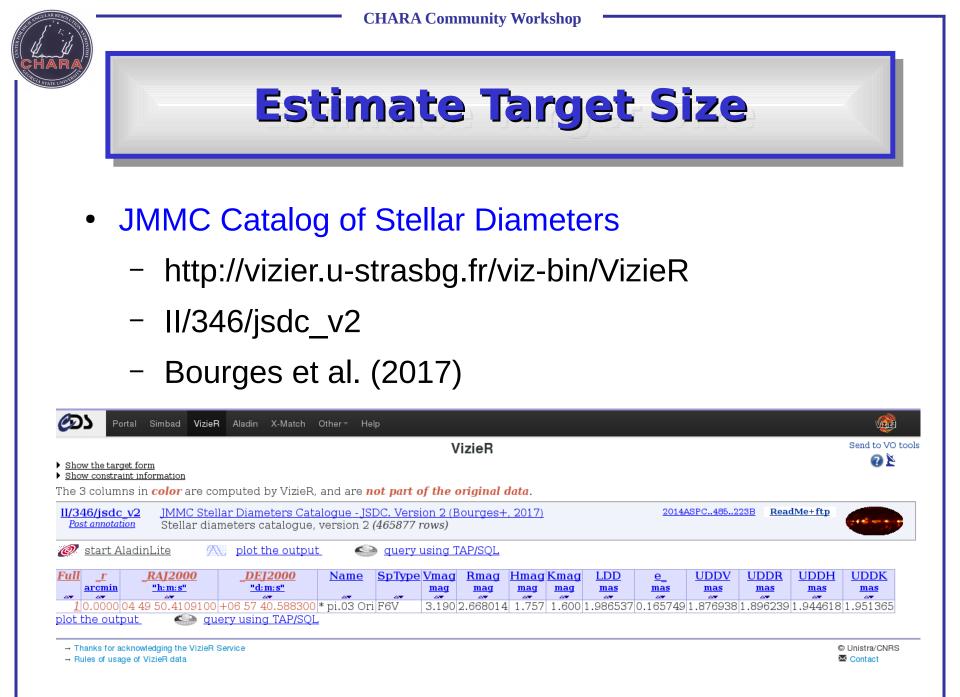


• Select favorite target

• Or use example target:

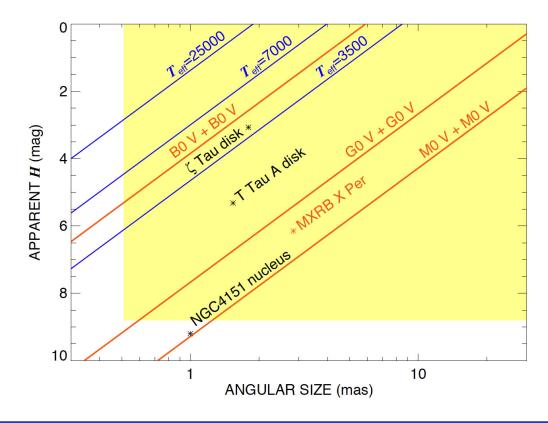
HD 30652: F6V, V=3.2, H=1.8, K=1.6

(resolved by Boyajian et al. 2012)





- Color and temperature relations
- Look up tables



# **Telescopes and Baselines**

A CARLES AND	Baseline	Length (m)	Res (mas)
60-inch telescope	E1-S1	331	0.52
Half-million-galle S1	W1-E1	314	0.55
	E1-S2	302	0.57
100-inch telescope	E2-S1	279	0.62
Control/Office Exhibit Building Beam Combining Lab Mt. Wilson Observatory Museum	W1-S1	279	0.62
Site Manager's Residence	W1-E2	251	0.69
	W1-S2	249	0.69
W2 Light pipes to central facility W1	E2-S2	248	0.69
Engineering Shop	W2-E1	222	0.78
	W2-S1	211	0.82
E1 E2	W2-S2	177	0.97
Six CHARA Array 1-meter telescopes	W2-E2	156	1.10
CHARA Array of Georgia State University	W2-W1	108	1.61
	E2-E1	66	2.65
Resolutions listed for H-band (1.6 µm)	S2-S1	34	5.08

### Selecting Calibrators: SearchCal

i www.jmmc.fr/searchcal\_page.htm

JAAAAC

#### www.jmmc.fr/searchcal\_page.htm

#### JEAN-MARIE MARIOTTI CENTER Infrared and Optical Interferometry for Astronomy

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Who are we ?	SearchCal: the JMMC Evolutive Search Calibrator Tool 🔊 subscribe to SearchCal feed	
Who was JMM ? Partners Structure	SearchCal is a tool developed by the JMMC Working Group "catalogue of calibration sources" to assist the astronomers in this calibrator selection process for long baseline interferometric observations.	
Working Groups	The science behind SearchCal is described in the 2016 paper "Pseudomagnitudes and differential surface	
Training	brightness: Application to the apparent diameter of stars." by Chelli A., Duvert G., Bourgès L. et al., 2016,	
Proposal Preparation	A&A, 589, 112.	
ASPRO GetStar SearchCal VLTI Proposals	The SearchCal user interface (check download link below) give access to the complete JSDC catalog for stars having a spectral type estimate and to a less precise estimate for the other 2 million stars of the Tycho list for which the spectral type has not been measured (yet).	
Data Processing	Additional help in using SearchCal results can be found on the SearchCal Wiki Page.	
AMBER PIONIER OIFits Explorer	Last Minute: expect some problems with the few stars identified by SIMBAD as planet-hosting stars (confusion). This will be corrected ASAP.	<u> </u>
Oifits Validator	Acknowledgement	C
Data Analysis LITpro Olmaging	If this software was helpful in your research, please use this acknowledgement and reference paper.	D
	A previous version of SearchCal was based on the following works:	<b>D</b>
Databases BadCal CalEx OiDB JMDC JSDC	<ul> <li>Bright calibrator request scenario : Bonneau D., Clausse JM., Delfosse X., et al., 2006,A&amp;A 456, 789</li> <li>Faint calibrator request scenario : Bonneau, D., Delfosse, X., Mourard, D., et al., 2011,A&amp;A 535, A53</li> </ul>	<u></u>
VO tool AppLauncher		
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Publications	Download Application Installation Help	
Job Offers		

Click on "Download Application"



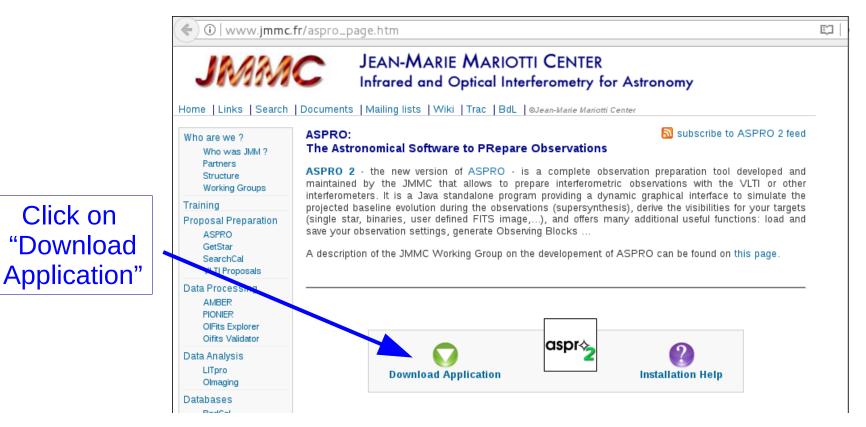
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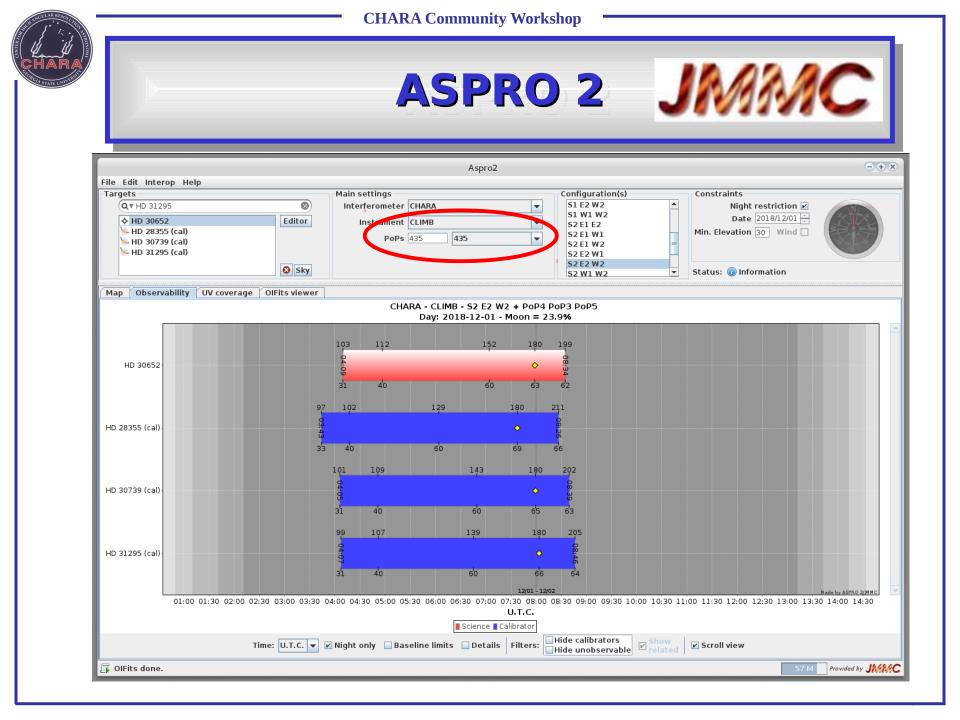
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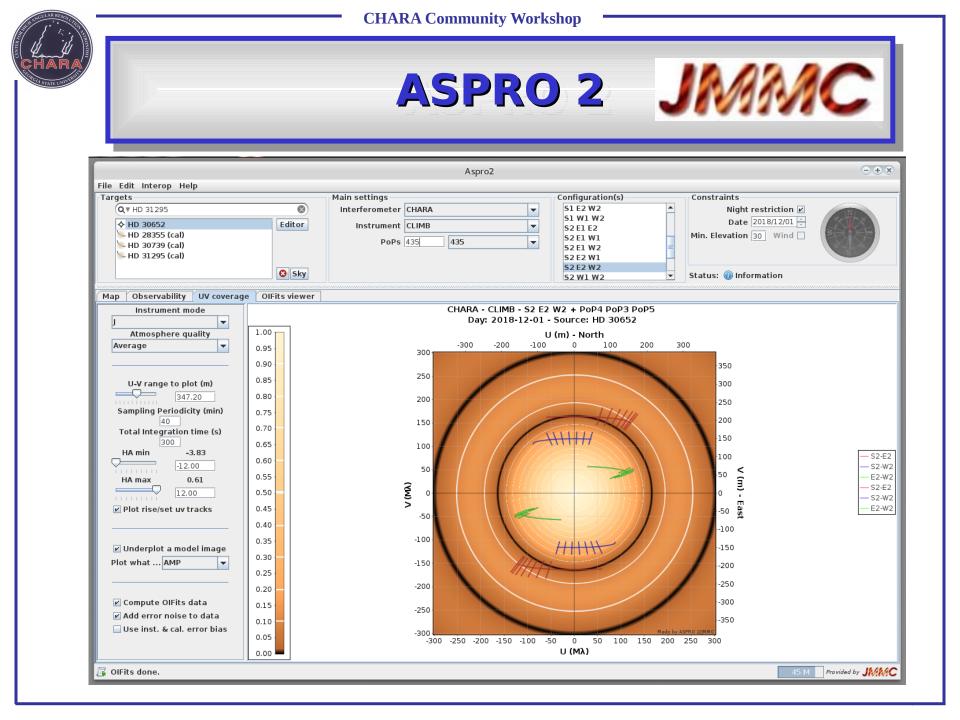
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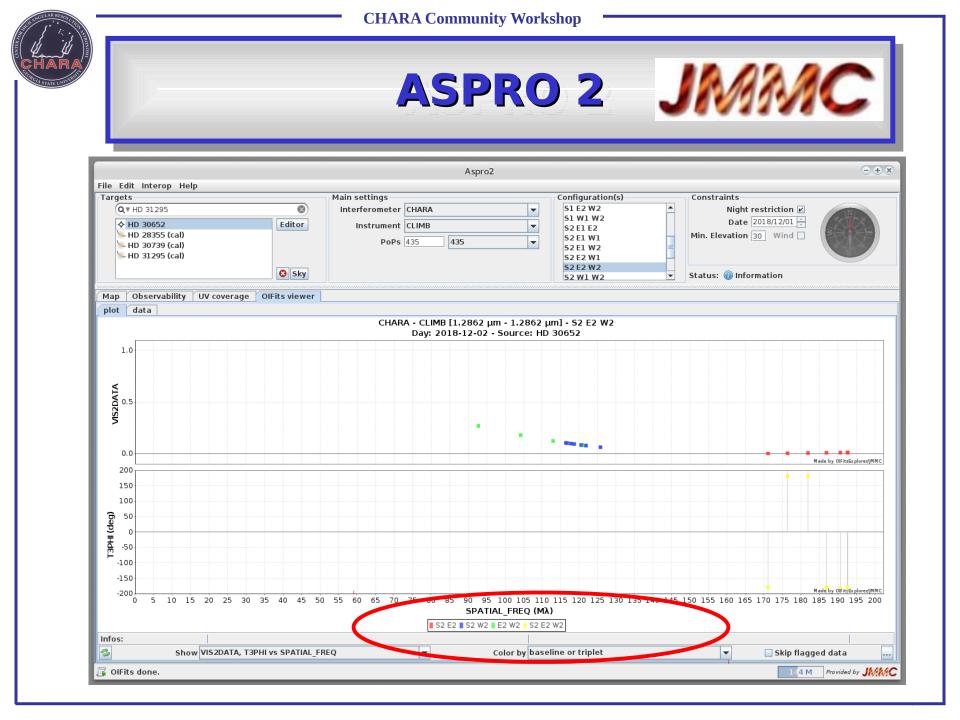


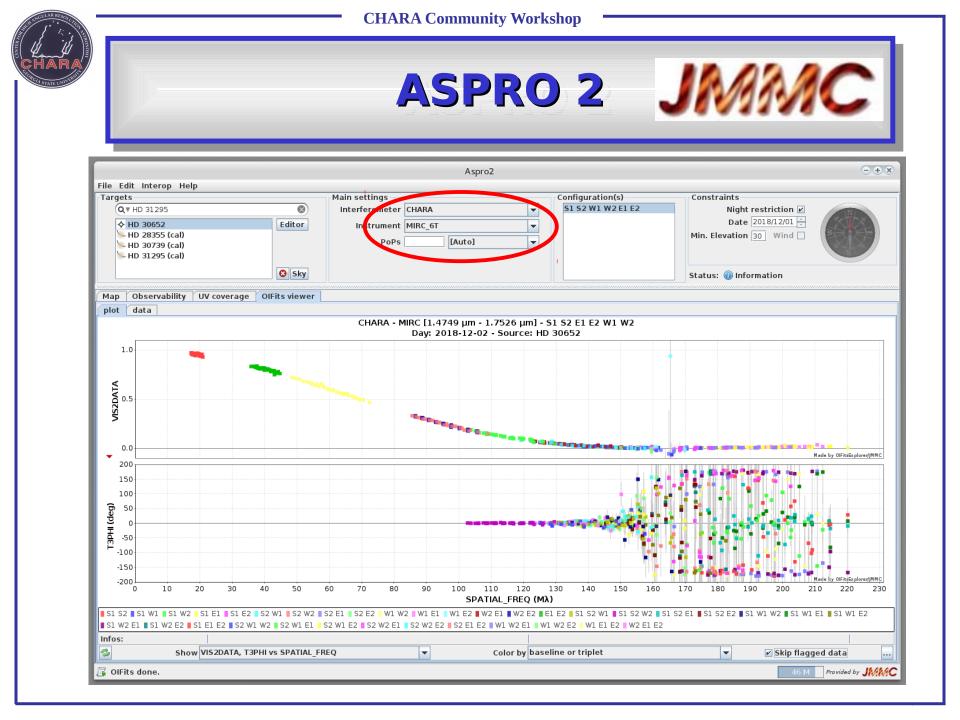
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