

ICCD SPECKLE OBSERVATIONS OF BINARY STARS. XI. MEASUREMENTS DURING  
1991–1993 FROM THE KITT PEAK 4 m TELESCOPE

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ABSTRACT

One-thousand one-hundred ninety-seven observations of 730 binary star systems, observed by means of speckle interferometry with the 4 m telescope on Kitt Peak, are presented. Included in these binary stars are new interferometric companions to five visual binaries. These measurements, made mostly during the period 1991 to 1993, comprise the 11th installment of results stemming from our speckle program at the 4 m class telescopes on Kitt Peak, Cerro Tololo, and Mauna Kea.

1. INTRODUCTION

This paper is the 11th in our current series of reports on the continuing effort to provide high accuracy, high angular resolution measurements of binary star systems by speckle methods. We here present measurements from the KPNO 4 m telescope, obtained mainly during 1991 April/May, 1991 November, 1992 April, and 1993 March. The data published in this and in our previous series of speckle measurements comprise over 80% of all published interferometric measurements of binary stars.

2. NEW MEASUREMENTS

Our basic ICCD speckle camera remained identical to that described earlier in this series (see McAlister *et al.* 1987), and all data were reduced utilizing the “directed vector autocorrelation” algorithm described by Bagnuolo *et al.* (1992). Absolute calibration of our results was made using a double-slit pupil mask, as described by McAlister *et al.* (1987). An independent  $\theta$  calibration for each run was determined by trailing a star across our acquisition field with the telescope drive off.

The GSU speckle camera was scheduled for a total of 17 nights during these four KPNO runs. One full night and a few partial nights were lost to clouds and/or poor seeing, but in the remaining time we were able to obtain nearly 3700 observations. Some of these observations, made of stars in the Hyades, Pleiades, and IC 4665 clusters, have already been published in Papers VII and X in this series (see Mason *et al.* 1993a, b). After removing these stars, plus observations of calibration objects and unresolved stars, and after averaging observations of stars observed using more than one filter,

we are left with 1162 observations. Added to this are 16 other ICCD measurements and 12 unresolved measures made between 1982 and 1990, as well as 7 unresolved measures made with a photographic speckle camera on the KPNO 2.1 m telescope between 1976 and 1979 (see McAlister *et al.* 1984 and references therein). These bring the totals reported here to 1197 observations of 730 stars.

Five newly discovered companions, each an additional component to a known binary system, are included in these totals. Table 1 gives a brief description of these systems and an estimate of their orbital periods. In determining these periods, circular orbits at the discovery separation are assumed and both components are assumed of equal brightness and mass [based on spectral type and estimated from tables in Allen (1973)]. Distances are derived from  $V$  and spectral type or, in the case of the nearby star BD+39 2376, from a parallax value given by Woolley *et al.* (1970).

The new measurements of binary stars are presented in Table 2, (this table is presented in its complete form in the ApJ/AJ CD-ROM Series, Vol. 3, 1994), where we use the same condensed format of previous papers in this series. The coordinates in Table 2, which also serve as the *Washington Double Star Catalog* (WDS) number, are for equinox 2000.0, but position angles have not been corrected for precession

TABLE 1. New multiple star components.

$\alpha, \delta$ (2000)	ADS	CHARA Number	HD/BD	$V$	Spectral Type	Discovery Epoch	Discovery Separation	$P_{est}$ (yrs)
08095+3213	6623	190 Aa	67501	6.7	F2	1991.3210	0".148	120
10453+3831	7915	191 Aa	+39 2376	9.23	dM2	1983.4277	0.381	25
11150+3735	8102	192 Aa	97731	8.2	G5	1987.2666	0.093	30
13197+4747	8862	193 Aa	115953	8.54	K0	1992.3126	0.109	25
17215+2845		194 Aa	157358	6.35	G0III	1991.3247	0.052	90

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## Notes to TABLE 2

- 02366+1266=HR 763=McA 7:** These three measurements of 31 Ari were based on reprocessing of archival video data. A combined spectroscopic/speckle orbital analysis of this very close system is now in progress, in collaboration with F. C. Fekel.
- 04404+1631=vB 185=CHARA 154:** This confirms the discovery observation of this Hyades binary, made in 1991 (Mason *et al.* 1993a).
- 05271+1758=ADS 4038=McA 19 Aa:** All archival data for this system were reprocessed using more powerful algorithms than earlier available, resulting in four new unresolved measures as well as the new 1991 measurement. In addition, new  $\rho$  and  $\theta$  values have been determined for the observations originally published from 1986.9–1988.3 (McAlister *et al.* 1989), 1988.7–1989.2 (McAlister *et al.* 1990), and 1990.3–1990.8 (Hartkopf *et al.* 1992). The date of the 1988.2518 measurement was incorrectly given as 1988.2490 in McAlister *et al.* (1989). Orbital analysis of this system is in progress; a preliminary orbit gives a period of  $\sim 15.3$  years and a semi-major axis of 0".080.
- 05373+6642=Mir 314:** Published orbital elements for this system (see Table 3) yield very large residuals, especially in  $\theta$ . A new set of elements, including an inclination considerably larger than that found by either Baize or Muller, has been determined; these elements are included in a paper in preparation.
- 06017+2224=HR 2116=CHARA 161:** This confirms the discovery observation, made in 1988 (McAlister *et al.* 1993).
- 06212+2932=HR 2272=CHARA 165:** This confirms the discovery observation, made in 1988 (McAlister *et al.* 1993).
- 06418+4031=ADS 5332=A 218:** The orbital elements of Baize (see Table 3) do not seem to agree with published visual or speckle data; nor do they reproduce Baize's ephemeris.
- 07277+2127=ADS 6089=McA 30 Aa:** A combined spectroscopic/speckle orbit of this system has been undertaken in collaboration with F. C. Fekel.
- 08017+6019=HR 3109=McA 33:** The orbit by Baize (see Table 3) yields rather large  $\theta$  residuals for these and earlier speckle data as well. A new set of orbital elements has been determined for a paper in preparation.
- 08095+3213=ADS 6623=CHARA 190 Aa:** This is a new component to the wide (2".8) pair STF 1187, discovered by F. G. W. Struve in 1829.
- 10453+3831=ADS 7915=Ho 532 AB+CHARA 191Aa:** The Aa pair is a new component to the 0".7 pair Ho 532. This new component was actually first resolved in 1983 (McAlister *et al.* 1987) but it was believed at the time that the measurement was of the known pair (the wide component laid outside the processor "window" of our old autocorrelator). Reprocessing of our archival video data allowed us to resolve both components in the 1983 and 1984 data sets, although both early measurements of the Aa component are quite weak.
- 11150+3735=ADS 8102=CHARA 192 Aa:** This is a new component to the 0".6 pair STT 232, discovered by Otto Struve in 1843. Reprocessing of archival data [whose 1987 and 1988 measurements for the wide pair were originally published in McAlister *et al.* (1989)] has allowed us to confirm this discovery and also to note that the system appears to be accelerating toward periastron.
- 11182+3132=ADS 8119=STF 1523 AB:** An orbital analysis of  $\xi$  UMa, including both the AB pair, the astrometric companion, and the newly discovered Bc companion, has been submitted.
- 13175-0041=HR 5014=Fin 350:** An orbital analysis of this system is in progress. This binary is comprised of a pair of F0V stars of near-zero  $\Delta m$ , so there is considerable uncertainty as to whether the orbit is of short-period/high-eccentricity ( $P \sim 9$  yr,  $e \sim 0.6$ ) or long period/low eccentricity ( $P \sim 18$  yr,  $e \sim 0.0$ ). The rms residuals to the speckle observations are slightly smaller for the long-period orbit, but mass sums ( $a^3/P^2$ ) for the two orbits differ by just 2%, insufficient to rule out either solution on the basis of plausibility.
- 13197+4747=ADS 8862=CHARA 193 Aa:** This appears to be a new component to the 1".5 49-year period pair of K/M dwarfs Hu 644. This is the only one of the current crop of new interferometric binaries which has not yet been confirmed.
- 15245+3723=ADS 9626=CHARA 181 Aa:** This confirms the discovery observation of this pair, made in 1988 (McAlister *et al.* 1993).
- 16317-0215=ADS 10095=A 693:** Residuals to the speckle data suggest that the orbital period of this system is perhaps 10–15 years longer than the 91 years found by Baize (see Table 3) and that periastron passage occurred about 1990.
- 17088+6543=HR 6396= $\zeta$  Dra:** Our new speckle observations of this very close interferometric binary are poorly fit by the estimated orbital elements of Zulević (see Table 3). Clearly more data are needed before a meaningful orbit of this pair can be determined.
- 17215+2845=HR 6466=CHARA 194 Aa:** This is a new close companion to the 0".7 pair Kui 80. The primary of this system is listed in the preliminary version of the fifth edition of the *Bright Star Catalogue* (Hoffleit & Warren 1992) as a G0 III spectroscopic binary.
- 17240-0921=RST 3972:** The 30-year period orbit of Scardia (see Table 3) gives a poor fit to recent speckle data, as do the earlier orbits of Heintz (1981) and Starikova (1983). A 15-year period orbit ( $\Delta m$  for this system is  $\sim 0$ , so there is a 180° ambiguity in both the visual and speckle data) seems to fit the data considerably better. However, masses derived from these elements (using a distance derived from the spectral types) are much too large, while those resulting from the longer period orbit are more reasonable.
- 17572+2400=HR 6697=McA 50:** All our CCD speckle data for this close system have been reprocessed and examined. In addition, we have listed 7 negative measures dating from observations made at the KPNO 2.1 m in the 1970's. A combined spectroscopic/interferometric orbit is now in preparation.
- 17584+0427=+04 3562=Kui 84:** A measurement published by McAlister *et al.* (1987) dating from 1985.4872 was misidentified as being of Kui 84; the measurement actually was of Kui 89 (WDS 18594–1250). In addition, an observation of Kui 84 dating from 1988.6655 ( $\theta = 270".6$ ,  $\rho = 0".134$ ) was published by McAlister *et al.* (1990). These data have been reprocessed, and we now believe the 1988.6655 measurement was spurious.
- 18594-1250=HR 7166=Kui 89:** This measurement was incorrectly attributed to Kui 84 (WDS 17584+0427) by McAlister *et al.* (1987).
- 19419+4015=HR 7499=Kui 94:** The orbital elements published by Baize (see Table 3) give quite large  $\theta$  residuals to our most recent speckle data. New elements have been determined for a paper in preparation.
- 20158+2749=HR 7744=CHARA 94 Aa:** These two observations confirm the discovery measurement of this close component, made in 1985 (McAlister *et al.* 1987), and indicate considerable motion during this interval.
- 20462+3358=ADS 14274=CHARA 100 Aa:** This finally confirms the discovery observation of this component, made in 1983 (McAlister *et al.* 1987). A number of unresolved observations of this very close system, dating from 1979 to 1981, were published by Hartkopf & McAlister (1984).
- 20474+3629=ADS 14296=McA 63 Aa:** Orbital analysis of  $\lambda$  Cyg is in preparation; a preliminary orbit gives a period of 11.8 years and a semi-major axis of 0".049.
- 21145+1001=ADS 14773=STT 535 AB:** A combined spectroscopic/speckle orbit of this system has been undertaken in collaboration with F. C. Fekel.
- 23460+4625=HR 9003=McA 75 Aa,c:** With this measurement we have finally confirmed the discovery observation of this component, made by McAlister *et al.* (1984) in 1981!

\*Table 2 is also presented in its complete form in the AAS CD-ROM Series, Volume 3, 1994.

Finally, in Table 3 we give  $O-C$  residuals for orbital elements published within the past five years, as tabulated by Worley (1994) in his computer supplement to the *Fourth Catalogue of Orbits of Visual Binary Stars* (Worley & Heintz 1983). Dates are repeated for stars having more than one published set of orbital elements. Predicted  $\theta$  and  $\rho$  values (enclosed in parentheses) are given in lieu of  $O-C$  values for the unresolved observation dates listed for WDS 17572 +2400=McA 50. Also, two possible sets of orbital elements

are given by Hartkopf *et al.* (1989) for WDS 18117 +3327=B 2545 AB, hence the two entries in Table 3 for the same date and paper. The coded references in the last columns (usually based on the first three letters of the first author's name, plus the publication date) are given in full in the reference list below.

As always, we are indebted to the efforts of the telescope operators in maintaining the highest observing efficiency. We

TABLE 3. Residuals to recently published orbits.

WDS $\alpha, \delta$ (2000)	Discoverer Designation	Date (1900+)	$\Delta\theta$ ( $''$ )	$\Delta\rho$ ( $''$ )	Orbit Reference	WDS $\alpha, \delta$ (2000)	Discoverer Designation	Date (1900+)	$\Delta\theta$ ( $''$ )	$\Delta\rho$ ( $''$ )	Orbit Reference
00335+4006	Ho 3	91.9016	5.6	0.040	BAI91a			91.3267	-3.5	-0.012	BAI89b
00352-0336	Ho 212 AB	91.8934	-0.3	0.000	HAR89			91.9052	-2.1	-0.009	
00507+6415	McA 2	91.8934	16.2	0.018	BAL89			92.3068	-1.2	-0.007	
00516+2238	A 1808	91.8936	4.3	0.004	BAI89a			92.3125	-4.0	-0.010	
00550+2338	STF 73 AB	91.8936	6.2	-0.005	D&C90a			93.2026	-3.9	0.001	
00568+6022	Bu 1099 AB	91.8934	6.6	0.001	COL92	10083+3137	Kui 48 AB	91.8945	3.8	-0.006	BAI93b
		91.9043	0.3	0.004				92.3069	1.7	-0.009	
01198-0029	Fin 337 BC	91.8936	-16.4	0.004	BAI91a			93.2025	2.7	-0.006	
01233+5808	STF 115 AB	91.9017	3.3	-0.005	HAR89	10427+0335	A 2768	91.3212	-2.5	0.010	HAR89
01512+2439	Ho 311	91.8936	1.8	0.002	HAR89			91.9053	-0.7	0.005	
		91.8936	-7.2	0.087	ZUL92			92.3097	-0.6	0.009	
02157+2503	Cou 79	91.8936	-1.6	-0.017	HAR89			92.3126	-0.5	0.007	
		93.2047	2.1	-0.004				93.1969	-1.7	0.010	
02231+7021	MLR 377	91.9017	-7.8	0.010	MUL91	10585+1711	A 2375	91.3268	11.2	0.055	BAI89a
02366+1226	McA 7	82.7659	-0.7	0.005	BAL89	11136+5525	A 1353	91.3241	0.0	0.007	D&L94
		84.7046	-11.2	0.004				91.9053	-0.3	0.005	
		88.6554	-13.4	0.010				92.3125	0.1	0.009	
02383+4604	A 1278	91.8937	3.1	-0.005	ZUL92	11191+3811	CHARA 133	93.1970	-3.2	-0.001	McA93
02396-1153	Fin 312	91.8937	-1.6	0.000	HAR89	11308+4117	STT 234	91.3295	0.2	-0.009	COU89
03337+5752	CHARA 117	91.8937	0.3	0.000	McA92			91.9053	0.9	-0.008	
		91.9019	-0.2	-0.003				92.3071	0.6	-0.007	
04361+0813	A 1840 AB	91.9020	-0.5	-0.019	OLE90			93.1970	0.5	-0.007	
04432+5932	A 1013	91.8969	-0.3	-0.006	D&C90c	11520+4806	Hu 731	91.3241	-1.1	0.124	D&L92
05373+6642	MLR 314	91.9049	-25.7	0.004	MUL91	11551+4629	A 1777 AB	91.3213	-4.8	0.030	BAI91a
		93.2023	-31.0	-0.006				91.3269	-6.6	0.029	
		91.9049	-25.5	-0.014	BAI91c			91.9053	-9.7	0.020	
		93.2023	-30.0	-0.028				92.3071	-12.5	0.023	
06171+0957	Fin 331 Aa	91.8969	-0.6	-0.007	BAI91c			93.1970	-17.0	0.025	
06289+2014	BTZ Aa	91.8969	5.5	-0.039	BAI92a	12160+4807	Hu 736	91.3213	-38.3	-0.002	BAI93b
06383+2859	McA 27	91.8969	2.9	0.030	HAR89			92.3073	-43.1	0.002	
		93.2024	5.3	0.046				91.3213	-0.2	-0.018	SCA93
		91.8969	2.4	0.027	BAI89b			92.3073	-4.9	-0.020	
		93.2024	4.3	0.035		12199-0040	McA 37	91.3186	6.1	-0.005	BAI89b
06418+3041	A 218	91.8997	-29.8	-0.189	BAI92a			92.3098	8.6	0.003	
06474+1812	STT 156	92.3121	-2.7	-0.043	BAI92a			91.3186	-2.0	-0.001	HAR92
06503+2410	Cou 768	91.9052	1.7	-0.027	BAI89a			92.3098	-2.2	0.005	
		93.2050	4.2	-0.003		13100+1731	STF 1728 AB	91.3187	-0.3	0.000	HAR89
07352+3058	STT 175 AB	91.8943	1.5	0.003	HAR89			91.3269	-0.1	0.000	
		92.3068	1.8	-0.002				92.3098	-0.3	-0.001	
		93.1967	1.6	0.001				92.3127	-0.2	-0.001	
07518-1352	Bu 101	91.9052	0.1	0.006	HAR89			93.1973	-0.4	-0.003	
		93.1968	-0.1	0.004		13202+1747	A 2166	91.3187	1.1	-0.009	BAI89c
07561+2342	Cou 929	92.3068	-3.6	0.008	BAI92b			92.3098	1.1	-0.009	
		93.1967	-4.4	0.002				92.3127	1.0	-0.008	
07573+0108	STT 185	91.9052	2.0	-0.004	D&L92			93.1971	1.1	-0.013	
		93.1968	-3.5	-0.007		13258+4430	A 1609 AB	91.3187	-0.2	-0.007	HEI91
08017+6019	McA 33	91.3265	15.1	-0.003	BAI93a			92.3101	0.3	-0.006	
		91.8943	11.3	-0.005				92.3129	0.1	-0.004	
		92.3124	10.2	-0.007		13320+3109	Wor 24	90.2704	4.0	-0.006	BAI91a
08214-0136	STF 1216	91.3239	-0.4	0.006	D&L91			91.3270	-4.9	0.028	
08231+2001	Ho 525 AB	91.9025	1.7	0.099	BAI91a	14138+3100	Cou 606	91.3188	0.9	0.016	D&L92
		91.9025	2.7	0.079	BAI93a			92.3099	2.1	0.020	
08267+2433	A 1746 BC	91.3265	-1.0	-0.006	D&L93	14323+2641	A 570	92.3073	-2.7	0.004	HEI91
		92.3068	-4.1	-0.004				93.1975	-2.2	0.006	
		93.1967	-5.4	-0.012		14403+2158	McA 40	92.3074	-6.5	0.003	BAI89b
08468+0625	SP AB	92.3068	-1.0	-0.003	HAR89			93.1975	-5.1	-0.002	
		93.2026	1.0	-0.006				92.3074	-9.2	0.008	BAL89
08531+5458	A 1584	91.3267	-1.1	0.014	HEI91			93.1975	-6.1	-0.002	
		91.8944	-3.3	0.018		14534+1543	STT 288	91.3217	0.3	-0.025	ZUL90
		92.3068	-2.4	0.016		15136+3453	Ho 60	91.3297	-1.7	-0.008	BAI93b
09008+4148	Kui 37 AB	91.3210	-2.0	-0.023	HAR89			92.3101	2.1	-0.007	
		91.8944	-2.6	-0.017		15307+3810	Hu 1163	91.3271	0.8	0.007	COU90
		92.3068	-2.9	-0.018				92.3102	9.3	0.010	
		93.2025	-3.5	-0.019				93.1976	5.4	0.013	
09123+1459	Fin 347 Aa	91.3267	-4.0	0.000	HAR89	15318+4053	A 1634 AB	92.3102	19.7	-0.009	HAR89
		91.8943	-2.6	0.005				93.1976	9.3	0.004	
		92.3068	-0.3	0.000		15361+3948	STT 298 AB	91.3218	-1.2	-0.011	COU89
09379+4554	A 1765	91.3295	-0.2	-0.033	BAI91c			92.3102	-1.2	0.000	
		91.9027	1.4	-0.044				93.1976	-2.0	0.000	
		92.3070	1.0	-0.048		15370+6426	Hu 1168	91.3270	-7.3	0.029	BAI92a
		93.2051	-3.4	-0.048		15390+2545	Cou 612	91.3219	-2.0	0.034	BAI92b
09474+1134	McA 34	91.3267	20.7	0.026	HAR89			92.3101	0.0	0.016	
		91.9052	6.8	0.024				93.1975	0.5	0.025	
		92.3068	4.4	0.021							
		92.3125	1.5	0.018							
		93.2026	-1.1	0.019							

TABLE 3. (continued)

WDS $\alpha, \delta$ (2000)	Discoverer Designation	Date (1900+)	$\Delta\theta$ ( $''$ )	$\Delta\rho$ ( $''$ )	Orbit Reference	WDS $\alpha, \delta$ (2000)	Discoverer Designation	Date (1900+)	$\Delta\theta$ ( $''$ )	$\Delta\rho$ ( $''$ )	Orbit Reference
15420+0028	A 2176	91.3273	-8.1	0.010	BAI93b	17563+0259	A 2189	91.3302	1.0	-0.002	D&C91
		91.3298	-5.3	0.011		17572+2400	McA 50	76.3674	(161.8)	(0.098)	BAI93b
		92.3075	-4.0	0.005				76.5503	(168.2)	(0.103)	
		93.1974	-1.9	0.005				76.6158	(170.4)	(0.103)	
15428+2618	STF 1967	91.3271	-0.3	-0.001	HAR89			76.6213	(170.6)	(0.103)	
		92.3101	0.2	-0.003				77.3285	(210.1)	(0.056)	
		93.1975	-0.1	-0.002				78.3169	(108.8)	(0.025)	
15492+6032	Hu 912	91.3271	1.2	0.003	BAI92a			79.1904	(176.6)	(0.098)	
		91.3298	1.4	0.006				82.5028	(279.6)	(0.041)	
16057-0617	Fin 384 Aa	92.3075	-24.8	0.004	BAI92a			83.4203	(155.4)	(0.084)	
16137+4638	A 1642	91.3192	-0.7	0.063	BAI92b			83.7098	(166.8)	(0.103)	
		92.3103	-0.5	0.077				84.3759	(196.1)	(0.070)	
16254+3724	CHARA 55	93.2056	-0.3	0.008	McA93			84.3840	5.1	-0.004	
16300+3354	Hu 1173	91.3191	-4.7	-0.010	ZUL91			86.4100	-2.3	0.009	
16317-0215	A 693	91.3298	23.2	0.006	BAI91a			87.2728	(273.1)	(0.040)	
16341+4227	$\sigma$ Her	91.3192	-10.4	0.005	BAI89b			87.7617	(355.2)	(0.035)	
		92.3077	-8.1	0.016				88.2612	0.8	-0.004	
		93.1978	-3.1	0.014				88.6655	-3.0	-0.001	
		91.3192	-6.2	0.005	BAL89			89.2385	(201.2)	(0.064)	
		92.3077	-5.2	0.004				89.7084	82.6	0.026	
		93.1978	-1.3	0.003				90.2733	( 72.3)	(0.017)	
16515+0113	STT 315	91.3300	-0.1	0.002	D&L91			90.7542	5.8	-0.004	
		92.3076	0.3	0.001				91.3193	-3.7	0.008	
		93.2057	0.6	0.006				92.3077	(315.1)	(0.051)	
17075+3810	Cou 1291	91.3192	-3.7	0.013	BAI93c			93.2059	-1.6	0.002	
17081+3555	Hu 1176 AB	91.3192	-1.2	-0.007	HAR89	18043+4205	Cou 1786	91.3276	-1.2	-0.012	COU93
		91.3300	-1.9	-0.005				92.3132	-0.1	-0.013	
		92.3077	-0.9	0.001				93.2059	-0.8	-0.016	
		93.2056	0.6	-0.006				91.3276	-0.3	-0.007	BAI93c
		91.3192	0.4	-0.010	CES90			92.3132	-1.4	-0.002	
		91.3300	-0.4	-0.008				93.2059	-5.1	0.003	
		92.3077	-1.5	0.005		18063+3824	Hu 1186	91.3192	0.8	0.033	BAI92c
		93.2056	1.8	-0.006		18117+3327	B 2545 AB	92.3077	-10.0	0.011	BAI89a
		91.3300	-59.9	0.031	ZUL93			92.3077	4.9	0.004	HAR89
		92.3103	66.2	0.008				92.3077	7.8	-0.009	HAR89
17217+3958	McA 47	91.3192	4.8	0.008	BAI91a	18280+0612	CHARA 71	93.2059	-8.7	0.007	McA93
		92.3104	3.0	0.002		18384-0312	A 88 AB	92.3134	0.7	-0.006	HAR89
		93.2058	1.2	0.001		18413+3018	STF 2367 AB	92.3105	-0.3	-0.008	CES91
		91.3192	-0.8	0.010	SCA94			93.2059	-2.7	-0.024	
		92.3104	-0.2	0.002		18594-1250	Kui 89	85.4872	0.3	0.001	BAI89a
		93.2058	-0.2	-0.002		19419+4015	Kui 94	91.8959	-28.8	0.008	BAI91a
17221+2310	Cou 415	91.3301	-4.1	-0.006	BAI92b	19489+1908	AGC 11 AB	91.8960	0.7	0.000	HAR89
		92.3104	-3.3	0.004		20397+1556	WCK Aa	91.9014	-1.9	-0.006	HAR89
17240-0921	RST 3972	91.3301	-32.4	-0.055	SCA93	20474+3629	McA 63 Aa	91.9014	-2.3	-0.001	BAL89
		92.3131	-2.4	-0.048				91.9014	18.9	0.002	BAI91b
17323+2849	A 352	91.3301	2.6	0.010	D&C90d	21135+1559	Hu 767	91.9014	-1.8	-0.014	HAR89
		92.3104	3.4	0.014		21187+1134	Bu 163 AB	91.9014	0.9	0.021	CES91
		93.2058	-0.1	0.021		21425+4106	Kui 108	91.9014	0.9	-0.004	HAR89
17366+0722	A 1156	92.3104	2.8	0.000	D&C91	21446+2539	Bu 989 AB	91.8932	-0.5	-0.001	HAR89
17399-0039	Bu 631	91.3220	4.0	-0.020	BAI92a			91.8932	0.7	-0.001	CES90
		92.3132	3.3	-0.022		21502+1718	Cou 14	91.9014	0.4	0.004	HAR89
17490+3704	Cou 1145	91.3192	-0.7	0.017	BAI92a	22241-0451	Bu 172 AB	91.9015	-51.1	-0.057	BAI92a
		92.3132	1.9	0.023				91.9015	1.9	-0.003	HE191
		93.2058	2.4	0.033		22408+1432	Ho 296 AB	91.8989	0.0	-0.003	HAR89
17505+0715	STT 337	91.3302	0.0	0.007	D&C90b			91.8989	2.9	0.000	CES90
17543+1108	Fin 381	92.3132	0.3	-0.008	BAL89	22570+2441	Cou 542 Aa	91.8989	-7.8	0.005	COU90
		92.3132	-1.6	-0.007	CES90			91.8989	-0.4	0.007	COU93
		92.3132	-6.9	0.007	BAI92a						

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