

## Abstract

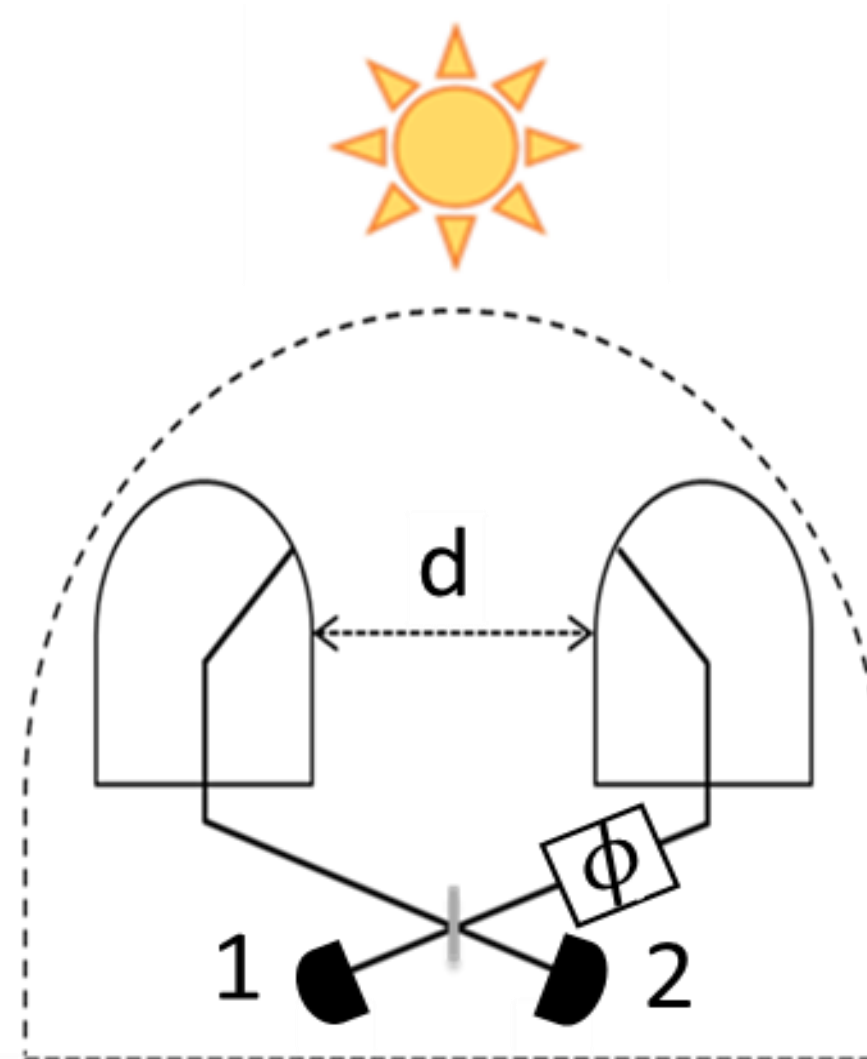
- We demonstrate the underlying mechanism for one version of quantum-enhanced telescope
- The “astronomical” and “terrestrial” source couple into three interconnected Hong-Ou-Mandel interferometers
- Amplitude of the complex visibility of the “astronomical” source is recovered in the presence of turbulence
- Inducing two phase shifts onto our “terrestrial” photon is sufficient to recover the visibility amplitude.

## Background

Interferometric telescope can simulate one massive telescope using an array of smaller receivers.

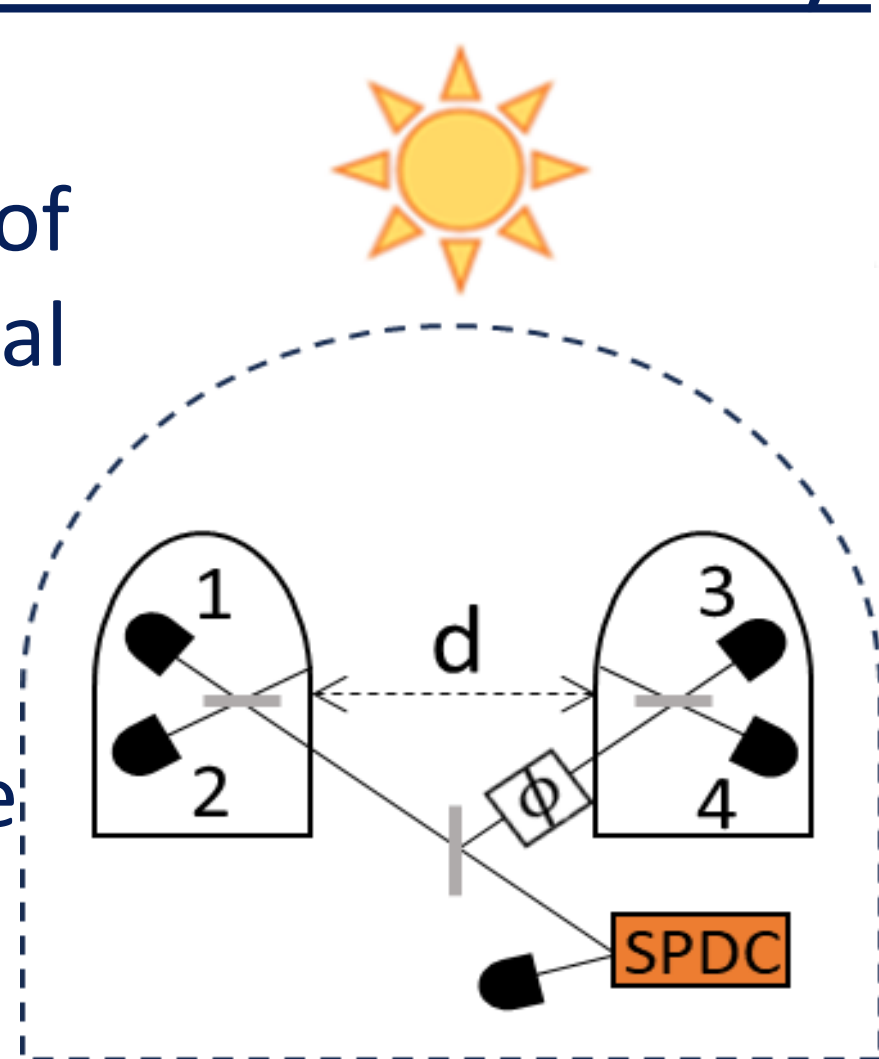
### Classical:

- Dependent on the brightness of the astronomical source
- Direct-detection interference is limited by loss

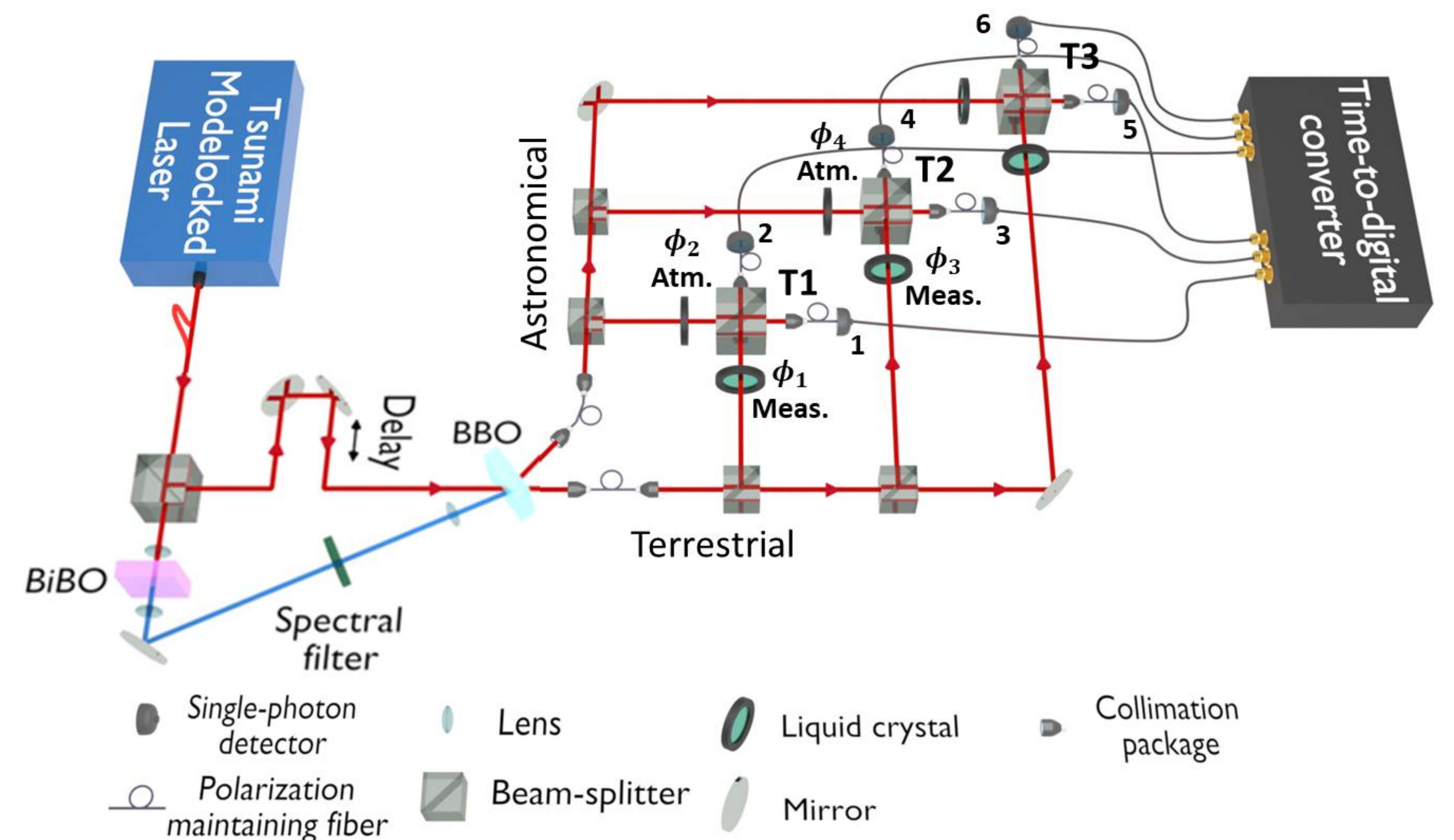


### Quantum-Enhanced (Gottesman et al. Scheme<sup>†</sup>):

- Dependent on the brightness of the astronomical and terrestrial source
- Limited by our ability to create desired single photons



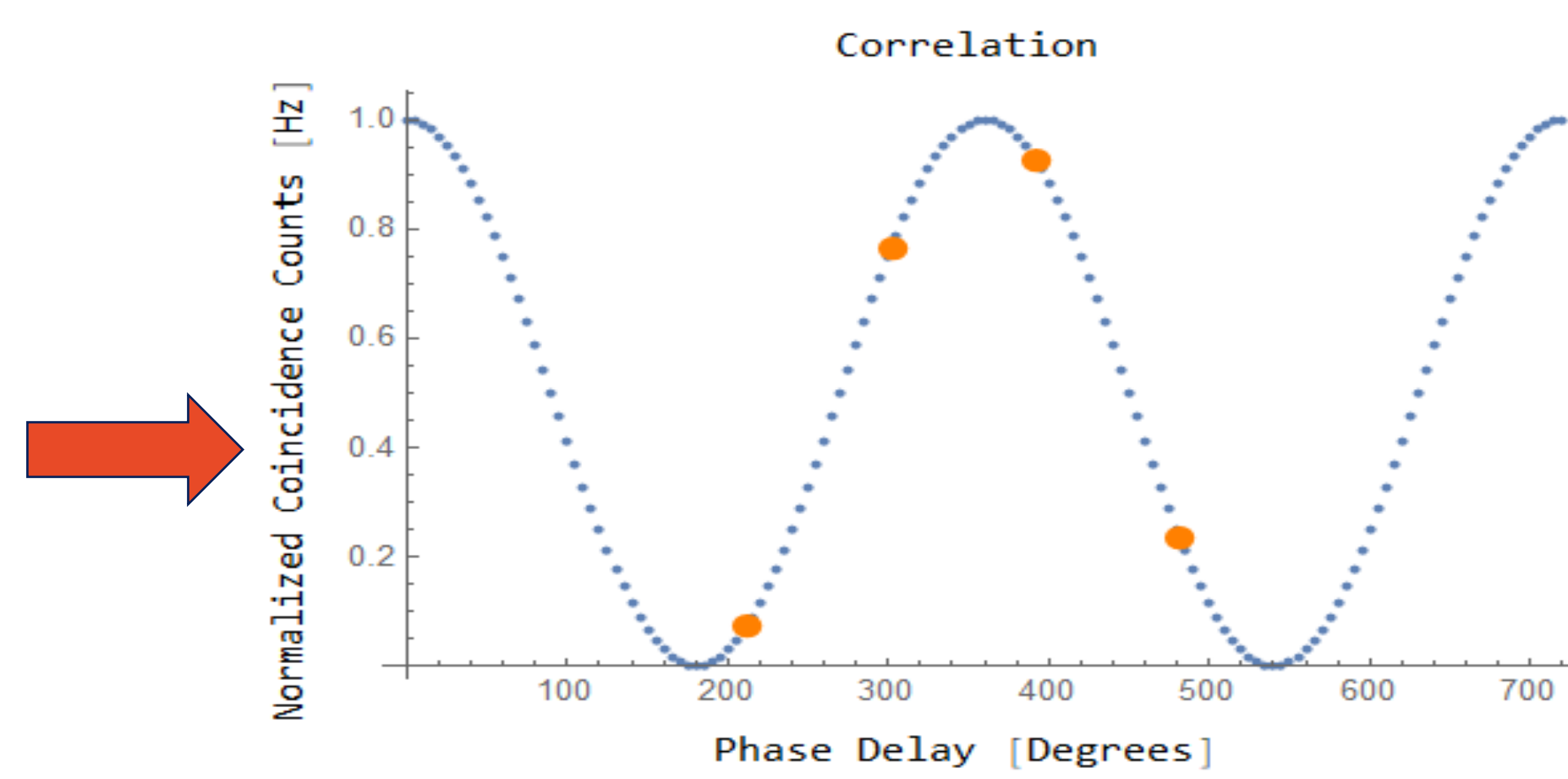
## Experimental Setup



- “Terrestrial” and “astronomical” sources are photons created via spontaneous parametric down conversion
- Liquid crystals in the bottom arms apply the controlled phase input
- Liquid crystals in the left arms mimic atmospheric turbulence
- We look for coincidences across our telescopes for (anti-)correlation measurements, i.e., detectors 1 and 3(4)
- Two-photon interference fringes → visibility amplitude

## 2-Photon Interference

- With limited photon number, we use a two-phase setting scheme (0°, 90°)
- Effectively 4 phases due to both beam splitter outputs
- Shown to be sufficient via numerical analysis

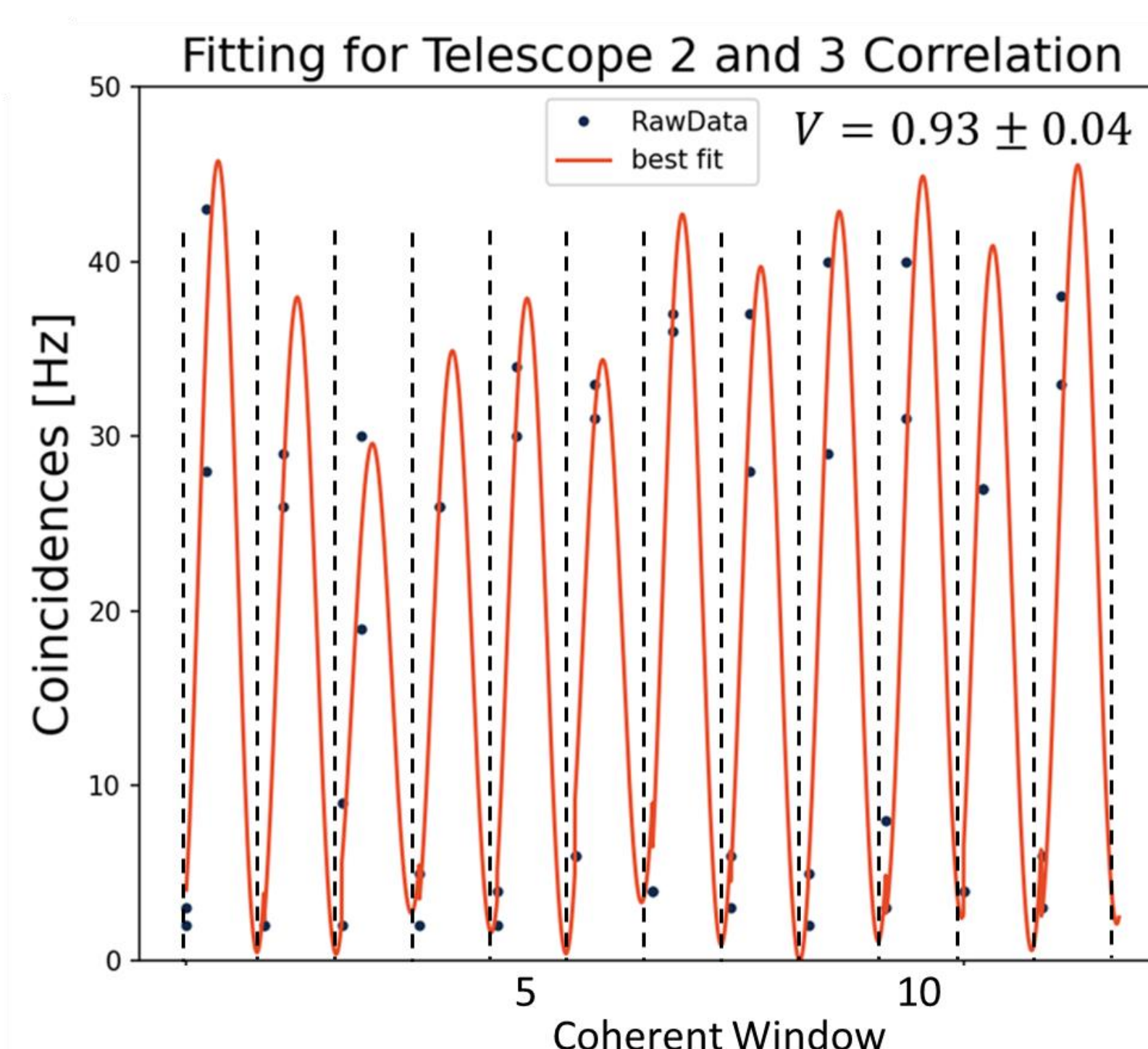


- Fringes obtained over 4 seconds consecutively for 150 total coherence windows

### Fitting Function

$$A_{i,jk} \left( 1 - \mathcal{V}_{i,j} \cos(\varphi_{app} + \varphi_{atm_{i,j}}) \right), \quad i, j \in [1,3], k \in [1,4]$$

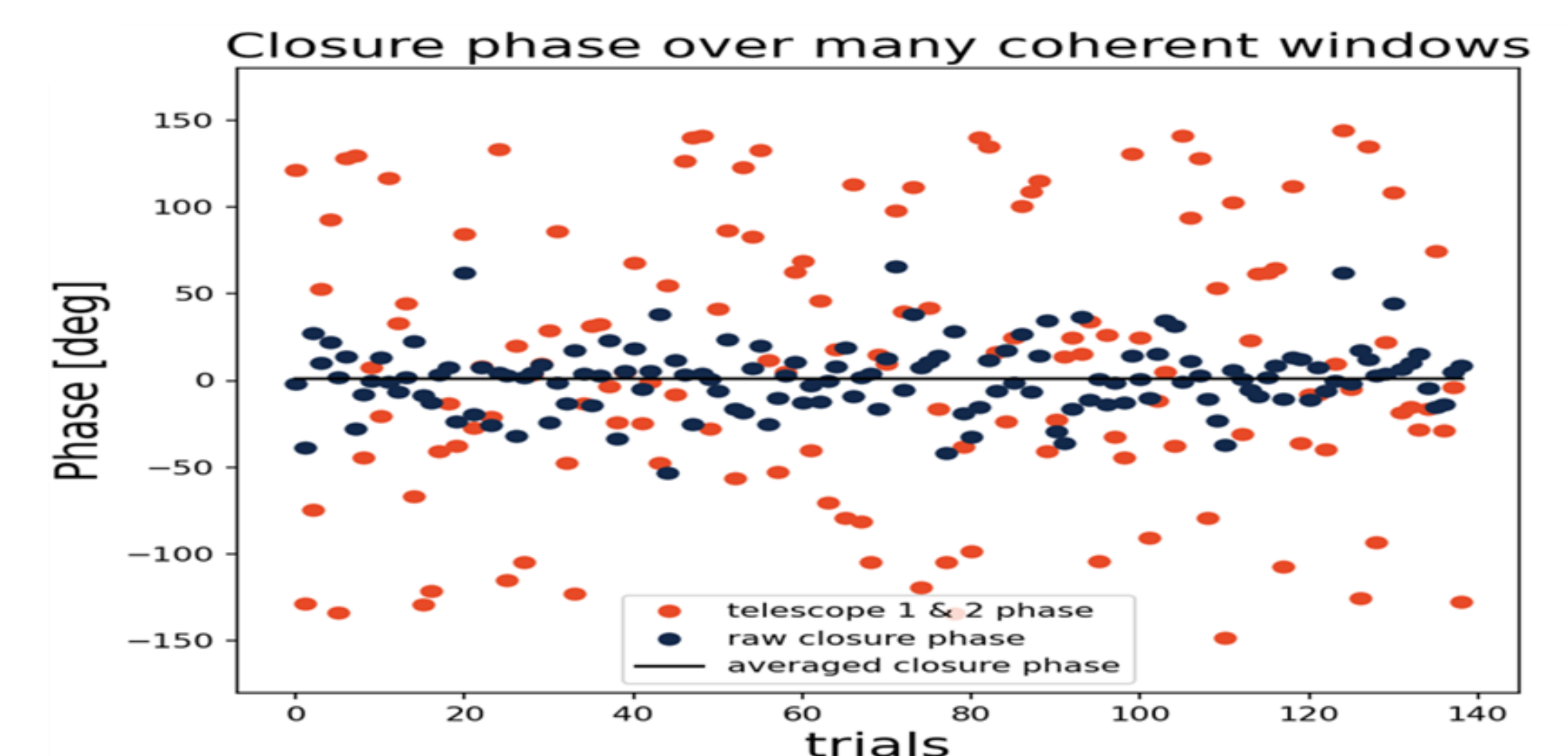
- No relative phases in a point source, accurate to average the visibility amplitudes  $\mathcal{V}_{i,j}$
- Error determined via the standard deviation of the visibility amplitude determined over many coherence windows



## Closure Phase

- Closure phases calculated in sets of three telescopes to diminish turbulence effects on recovering source information
- For a point source, the closure phase is 0°

$$\begin{aligned} \Phi(1,2) &= \Phi_0(1,2) + [\varphi(2) - \varphi(1)] \\ \Phi(2,3) &= \Phi_0(2,3) + [\varphi(3) - \varphi(2)] \\ \Phi(3,1) &= \Phi_0(3,1) + [\varphi(1) - \varphi(3)] \\ \Phi(1,2,3) &\equiv \Phi(1,2) + \Phi(2,3) + \Phi(3,1) \\ &= \Phi_0(1,2) + \Phi_0(2,3) + \Phi_0(3,1) \end{aligned}$$



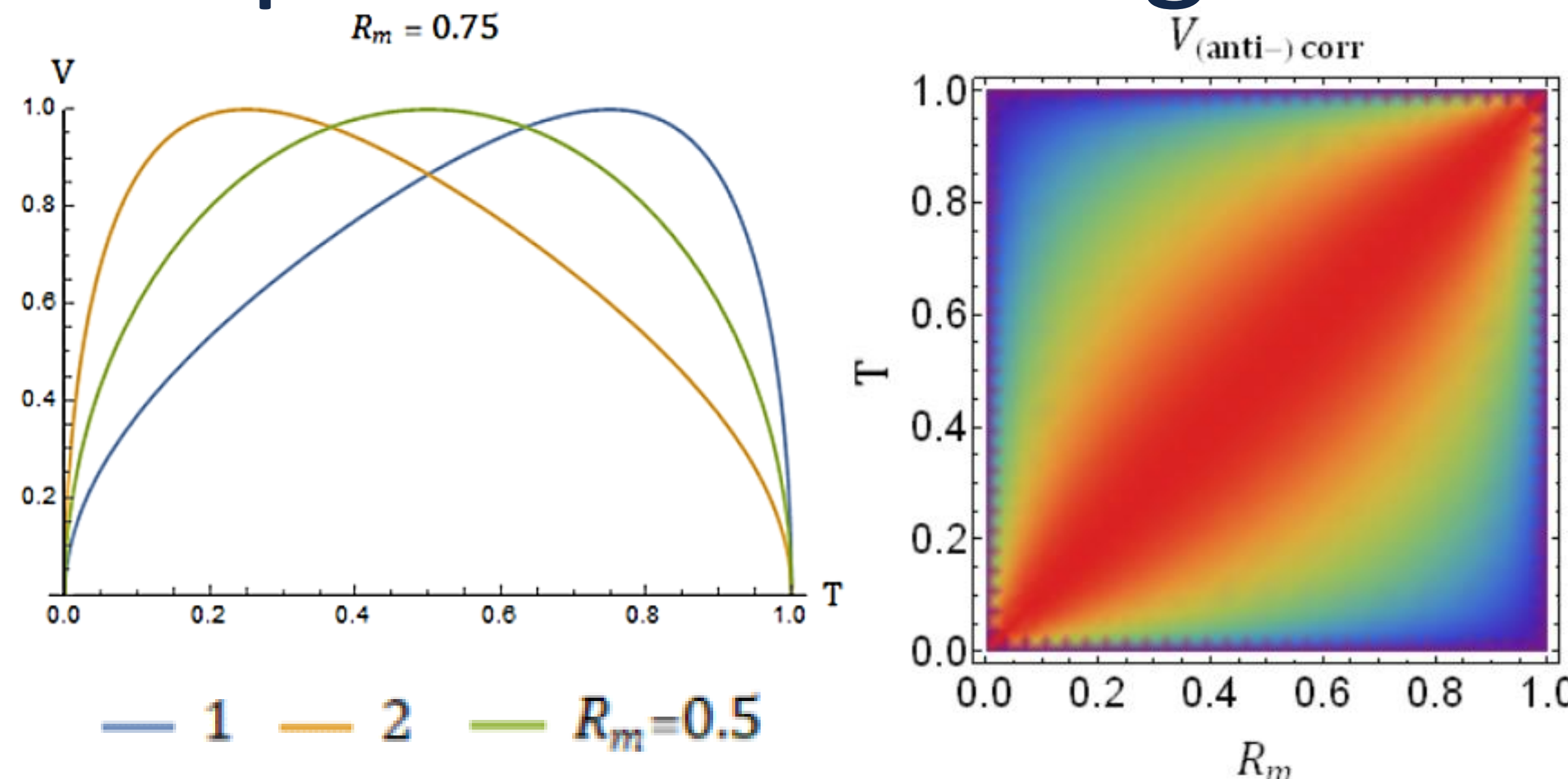
- Extended sources require information on complex visibility  $g_{ij} = \mathcal{V}_{i,j} e^{i\varphi_{ij}}$
- When there is turbulence, we need to take the product  $\mathcal{V}^c = |g_{12} g_{23} g_{31}|$
- If we instead do  $\mathcal{V}^i = |g_{12} g_{23} g_{31}|$  we see a loss of information
- We will use the notation  $\mathcal{V} = \overline{\mathcal{V}_{12} \mathcal{V}_{23} \mathcal{V}_{31}}$ , as our true visibility

- Turbulence will be:
  - S - standard deviation less than 45°
  - L - uniform distribution over 360°

	$\tau(s)$	$\tau(L)$
$\mathcal{V}$	$0.71 \pm 0.17$	$0.66 \pm 0.18$
$\mathcal{V}^c$	0.69	0.59
$\mathcal{V}^i$	0.42	0.0002

## Amplitude Matching

- $R_m = \frac{R_L}{R_R + R_L}$ ,  $R_L(R)$  is photon rate into the left (right) telescope
- T is the transmission of the recombining BS for direct detection or the distributing BS for the single photons



## Acknowledgements

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<sup>†</sup> D. Gottesman, T. Jennewein, and S. Croke, Longer-baseline telescopes using quantum repeaters, Phys. Rev. Lett. 109, 070503 (2012).